

NO-A105 920

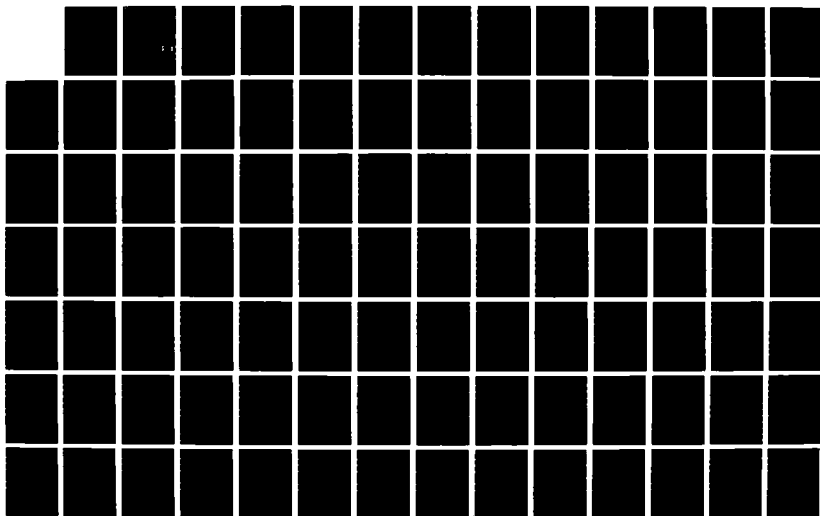
TEXTING THE FORMULARY: A COLLECTION OF EXAMPLES OF THE
USE OF TEX TO PROD. (U) NAVAL RESEARCH LAB WASHINGTON
DC D L BOOK ET AL. 86 OCT 87 NRL-RR-6844

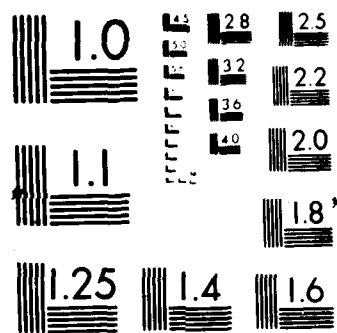
1/2

UNCLASSIFIED

F/G 20/9

NL





(2)



Naval Research Laboratory

Washington, DC 20375-5000

NRL Memorandum Report 6044

AD-A185 920

T_EXing the FORMULARY:
A Collection of Examples of
the Use of T_EX to Produce Ruled
Tables and Displayed Equations

DAVID L. BOOK

Laboratory for Computational Physics and Fluid Dynamics

TODD BRUN

*Physics Department
Harvard University, Cambridge, MA 02138*

ROBERT J. SCOTT

Berkeley Research Associates, Springfield, VA 22150

DTIC
ELECTE
NOV 10 1987
S D

October 6, 1987

Supported by the Office of Naval Research

87

REPORT DOCUMENTATION PAGE		Form Approved OMB No. 0704-0188	
1a REPORT SECURITY CLASSIFICATION UNCLASSIFIED		1b RESTRICTIVE MARKING A185920	
2a SECURITY CLASSIFICATION AUTHORITY		3 DISTRIBUTION / AVAILABILITY OF REPORT Approved for public release; distribution unlimited.	
2b DECLASSIFICATION / DOWNGRADING SCHEDULE			
4 PERFORMING ORGANIZATION REPORT NUMBER(S) NRI. Memorandum Report 6044		5 MONITORING ORGANIZATION REPORT NUMBER(S)	
6a NAME OF PERFORMING ORGANIZATION Naval Research Laboratory	6b OFFICE SYMBOL (If applicable)	7a NAME OF MONITORING ORGANIZATION	
6c ADDRESS (City, State, and ZIP Code) Washington, DC 20375-5000		7b ADDRESS (City, State, and ZIP Code)	
8a NAME OF FUNDING / SPONSORING ORGANIZATION Office of Naval Research	8b OFFICE SYMBOL (If applicable)	9 PROCUREMENT INSTRUMENT IDENTIFICATION NUMBER	
8c ADDRESS (City, State, and ZIP Code) Arlington, VA 22217		10 SOURCE OF FUNDING NUMBERS	
		PROGRAM ELEMENT NO 61153N	PROJECT NO DN280-69
		TASK NO RR011-09-43	WORK UNIT ACCESSION NO
11 TITLE (Include Security Classification) TeXing the FORMULARY: A Collection of Examples of the Use of TeX to Produce Ruled Tables and Displayed Equations			
12 PERSONAL AUTHOR(S) Book, D.L., Brun, T. and Scott, R.			
13a TYPE OF REPORT	13b TIME COVERED FROM _____ TO _____	14 DATE OF REPORT (Year, Month, Day) October 6, 1987	15 PAGE COUNT 133
16 SUPPLEMENTARY NOTATION			
17 CROSS-REFERENCE CODES		18 SUBJECT TERMS (Continue on reverse if necessary and identify by block number)	
FIELD	GROUP	SUB GROUP	
19 ABSTRACT (Continue on reverse if necessary and identify by block number)			
<p>The <i>NRI Plasma Formulary</i> is a collection of formulas and reference data in a handy pocket-sized booklet. This report contains some sixty pages of text, tables, and displayed equations, together with the associated <i>TEX</i> files. An enlarged copy (the output from an IMAGEN ink-jet printer) of each page of the <i>Formulary</i>, identical with that which was photographically reduced for the actual book, is reprinted facing the <i>TEX</i> source code that generated it. Users can browse through the report until they come to a table or displayed equation similar to the one they want to produce, then extract from the facing page the control sequences they need.</p>			
20 DISTRIBUTION STATEMENT OF ABSTRACT X UNCLASSIFIED 1 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 2 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 3 AVAILABLE ONLY TO SPECIFIED DOMAINS 4 AVAILABLE ONLY TO SPECIFIED COUNTRIES 5 AVAILABLE ONLY TO SPECIFIED GROUPS 6 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 7 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 8 AVAILABLE ONLY TO SPECIFIED DOMAINS 9 AVAILABLE ONLY TO SPECIFIED COUNTRIES 10 AVAILABLE ONLY TO SPECIFIED GROUPS 11 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 12 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 13 AVAILABLE ONLY TO SPECIFIED DOMAINS 14 AVAILABLE ONLY TO SPECIFIED COUNTRIES 15 AVAILABLE ONLY TO SPECIFIED GROUPS 16 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 17 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 18 AVAILABLE ONLY TO SPECIFIED DOMAINS 19 AVAILABLE ONLY TO SPECIFIED COUNTRIES 20 AVAILABLE ONLY TO SPECIFIED GROUPS 21 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 22 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 23 AVAILABLE ONLY TO SPECIFIED DOMAINS 24 AVAILABLE ONLY TO SPECIFIED COUNTRIES 25 AVAILABLE ONLY TO SPECIFIED GROUPS 26 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 27 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 28 AVAILABLE ONLY TO SPECIFIED DOMAINS 29 AVAILABLE ONLY TO SPECIFIED COUNTRIES 30 AVAILABLE ONLY TO SPECIFIED GROUPS 31 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 32 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 33 AVAILABLE ONLY TO SPECIFIED DOMAINS 34 AVAILABLE ONLY TO SPECIFIED COUNTRIES 35 AVAILABLE ONLY TO SPECIFIED GROUPS 36 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 37 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 38 AVAILABLE ONLY TO SPECIFIED DOMAINS 39 AVAILABLE ONLY TO SPECIFIED COUNTRIES 40 AVAILABLE ONLY TO SPECIFIED GROUPS 41 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 42 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 43 AVAILABLE ONLY TO SPECIFIED DOMAINS 44 AVAILABLE ONLY TO SPECIFIED COUNTRIES 45 AVAILABLE ONLY TO SPECIFIED GROUPS 46 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 47 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 48 AVAILABLE ONLY TO SPECIFIED DOMAINS 49 AVAILABLE ONLY TO SPECIFIED COUNTRIES 50 AVAILABLE ONLY TO SPECIFIED GROUPS 51 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 52 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 53 AVAILABLE ONLY TO SPECIFIED DOMAINS 54 AVAILABLE ONLY TO SPECIFIED COUNTRIES 55 AVAILABLE ONLY TO SPECIFIED GROUPS 56 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 57 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 58 AVAILABLE ONLY TO SPECIFIED DOMAINS 59 AVAILABLE ONLY TO SPECIFIED COUNTRIES 60 AVAILABLE ONLY TO SPECIFIED GROUPS 61 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 62 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 63 AVAILABLE ONLY TO SPECIFIED DOMAINS 64 AVAILABLE ONLY TO SPECIFIED COUNTRIES 65 AVAILABLE ONLY TO SPECIFIED GROUPS 66 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 67 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 68 AVAILABLE ONLY TO SPECIFIED DOMAINS 69 AVAILABLE ONLY TO SPECIFIED COUNTRIES 70 AVAILABLE ONLY TO SPECIFIED GROUPS 71 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 72 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 73 AVAILABLE ONLY TO SPECIFIED DOMAINS 74 AVAILABLE ONLY TO SPECIFIED COUNTRIES 75 AVAILABLE ONLY TO SPECIFIED GROUPS 76 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 77 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 78 AVAILABLE ONLY TO SPECIFIED DOMAINS 79 AVAILABLE ONLY TO SPECIFIED COUNTRIES 80 AVAILABLE ONLY TO SPECIFIED GROUPS 81 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 82 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 83 AVAILABLE ONLY TO SPECIFIED DOMAINS 84 AVAILABLE ONLY TO SPECIFIED COUNTRIES 85 AVAILABLE ONLY TO SPECIFIED GROUPS 86 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 87 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 88 AVAILABLE ONLY TO SPECIFIED DOMAINS 89 AVAILABLE ONLY TO SPECIFIED COUNTRIES 90 AVAILABLE ONLY TO SPECIFIED GROUPS 91 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 92 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 93 AVAILABLE ONLY TO SPECIFIED DOMAINS 94 AVAILABLE ONLY TO SPECIFIED COUNTRIES 95 AVAILABLE ONLY TO SPECIFIED GROUPS 96 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 97 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 98 AVAILABLE ONLY TO SPECIFIED DOMAINS 99 AVAILABLE ONLY TO SPECIFIED COUNTRIES 100 AVAILABLE ONLY TO SPECIFIED GROUPS 101 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 102 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 103 AVAILABLE ONLY TO SPECIFIED DOMAINS 104 AVAILABLE ONLY TO SPECIFIED COUNTRIES 105 AVAILABLE ONLY TO SPECIFIED GROUPS 106 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 107 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 108 AVAILABLE ONLY TO SPECIFIED DOMAINS 109 AVAILABLE ONLY TO SPECIFIED COUNTRIES 110 AVAILABLE ONLY TO SPECIFIED GROUPS 111 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 112 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 113 AVAILABLE ONLY TO SPECIFIED DOMAINS 114 AVAILABLE ONLY TO SPECIFIED COUNTRIES 115 AVAILABLE ONLY TO SPECIFIED GROUPS 116 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 117 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 118 AVAILABLE ONLY TO SPECIFIED DOMAINS 119 AVAILABLE ONLY TO SPECIFIED COUNTRIES 120 AVAILABLE ONLY TO SPECIFIED GROUPS 121 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 122 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 123 AVAILABLE ONLY TO SPECIFIED DOMAINS 124 AVAILABLE ONLY TO SPECIFIED COUNTRIES 125 AVAILABLE ONLY TO SPECIFIED GROUPS 126 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 127 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 128 AVAILABLE ONLY TO SPECIFIED DOMAINS 129 AVAILABLE ONLY TO SPECIFIED COUNTRIES 130 AVAILABLE ONLY TO SPECIFIED GROUPS 131 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 132 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 133 AVAILABLE ONLY TO SPECIFIED DOMAINS 134 AVAILABLE ONLY TO SPECIFIED COUNTRIES 135 AVAILABLE ONLY TO SPECIFIED GROUPS 136 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 137 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 138 AVAILABLE ONLY TO SPECIFIED DOMAINS 139 AVAILABLE ONLY TO SPECIFIED COUNTRIES 140 AVAILABLE ONLY TO SPECIFIED GROUPS 141 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 142 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 143 AVAILABLE ONLY TO SPECIFIED DOMAINS 144 AVAILABLE ONLY TO SPECIFIED COUNTRIES 145 AVAILABLE ONLY TO SPECIFIED GROUPS 146 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 147 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 148 AVAILABLE ONLY TO SPECIFIED DOMAINS 149 AVAILABLE ONLY TO SPECIFIED COUNTRIES 150 AVAILABLE ONLY TO SPECIFIED GROUPS 151 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 152 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 153 AVAILABLE ONLY TO SPECIFIED DOMAINS 154 AVAILABLE ONLY TO SPECIFIED COUNTRIES 155 AVAILABLE ONLY TO SPECIFIED GROUPS 156 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 157 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 158 AVAILABLE ONLY TO SPECIFIED DOMAINS 159 AVAILABLE ONLY TO SPECIFIED COUNTRIES 160 AVAILABLE ONLY TO SPECIFIED GROUPS 161 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 162 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 163 AVAILABLE ONLY TO SPECIFIED DOMAINS 164 AVAILABLE ONLY TO SPECIFIED COUNTRIES 165 AVAILABLE ONLY TO SPECIFIED GROUPS 166 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 167 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 168 AVAILABLE ONLY TO SPECIFIED DOMAINS 169 AVAILABLE ONLY TO SPECIFIED COUNTRIES 170 AVAILABLE ONLY TO SPECIFIED GROUPS 171 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 172 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 173 AVAILABLE ONLY TO SPECIFIED DOMAINS 174 AVAILABLE ONLY TO SPECIFIED COUNTRIES 175 AVAILABLE ONLY TO SPECIFIED GROUPS 176 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 177 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 178 AVAILABLE ONLY TO SPECIFIED DOMAINS 179 AVAILABLE ONLY TO SPECIFIED COUNTRIES 180 AVAILABLE ONLY TO SPECIFIED GROUPS 181 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 182 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 183 AVAILABLE ONLY TO SPECIFIED DOMAINS 184 AVAILABLE ONLY TO SPECIFIED COUNTRIES 185 AVAILABLE ONLY TO SPECIFIED GROUPS 186 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 187 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 188 AVAILABLE ONLY TO SPECIFIED DOMAINS 189 AVAILABLE ONLY TO SPECIFIED COUNTRIES 190 AVAILABLE ONLY TO SPECIFIED GROUPS 191 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 192 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 193 AVAILABLE ONLY TO SPECIFIED DOMAINS 194 AVAILABLE ONLY TO SPECIFIED COUNTRIES 195 AVAILABLE ONLY TO SPECIFIED GROUPS 196 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 197 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 198 AVAILABLE ONLY TO SPECIFIED DOMAINS 199 AVAILABLE ONLY TO SPECIFIED COUNTRIES 200 AVAILABLE ONLY TO SPECIFIED GROUPS 201 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 202 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 203 AVAILABLE ONLY TO SPECIFIED DOMAINS 204 AVAILABLE ONLY TO SPECIFIED COUNTRIES 205 AVAILABLE ONLY TO SPECIFIED GROUPS 206 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 207 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 208 AVAILABLE ONLY TO SPECIFIED DOMAINS 209 AVAILABLE ONLY TO SPECIFIED COUNTRIES 210 AVAILABLE ONLY TO SPECIFIED GROUPS 211 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 212 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 213 AVAILABLE ONLY TO SPECIFIED DOMAINS 214 AVAILABLE ONLY TO SPECIFIED COUNTRIES 215 AVAILABLE ONLY TO SPECIFIED GROUPS 216 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 217 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 218 AVAILABLE ONLY TO SPECIFIED DOMAINS 219 AVAILABLE ONLY TO SPECIFIED COUNTRIES 220 AVAILABLE ONLY TO SPECIFIED GROUPS 221 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 222 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 223 AVAILABLE ONLY TO SPECIFIED DOMAINS 224 AVAILABLE ONLY TO SPECIFIED COUNTRIES 225 AVAILABLE ONLY TO SPECIFIED GROUPS 226 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 227 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 228 AVAILABLE ONLY TO SPECIFIED DOMAINS 229 AVAILABLE ONLY TO SPECIFIED COUNTRIES 230 AVAILABLE ONLY TO SPECIFIED GROUPS 231 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 232 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 233 AVAILABLE ONLY TO SPECIFIED DOMAINS 234 AVAILABLE ONLY TO SPECIFIED COUNTRIES 235 AVAILABLE ONLY TO SPECIFIED GROUPS 236 AVAILABLE ONLY TO SPECIFIED INDIVIDUALS 237 AVAILABLE ONLY TO SPECIFIED ORGANIZATIONS 238 AVAILABLE ONLY TO SPECIFIED DOMAINS 239 AVAILABLE ONLY TO SPECIFIED COUN			

INTRODUCTION

The *NRL Plasma Formulary* is a collection of formulas and reference data in a handy pocket-sized booklet. Over thirty thousand copies of previous editions of the *Formulary* have been produced and distributed during the past decade and a half. Each revision before the present one precipitated a costly exercise in composition and layout. Finally the master became too dilapidated to use. We decided to retypeset the whole book using T_EX, the computerized typesetting system developed by Donald Knuth of Stanford University. Although this was an arduous task, we believe that the difficulty of preparing future editions will thereby be greatly reduced.

In recent years T_EX has been widely adopted by scientists, engineers, and others to produce documents containing technical information. The typical user, like us, is an author or scientific collaborator of the author, technically trained, computer-literate, and a clumsy typist. This user knows exactly how the final document should look and is strongly motivated to mine from *The T_EXbook* the requisite nuggets of wisdom, and to grind away as long as necessary, in order to achieve it. Consequently, highly trained (and highly paid) people are spending a significant fraction of their time doing what used to be regarded as clerical work. It follows that any tool that can make the operation of using T_EX more efficient is potentially valuable.

Knuth's comprehensive and very readable (if idiosyncratic) introductory manual, *The T_EXbook*, devotes about the same amount of space to the production of ruled tables as to diacritical marks in Central European languages (two and a half pages). For most users, the former is by far the more important application. Although the two examples of tables Knuth provides are illuminating and the diligent student can learn a great deal from them, the process is time-consuming. Experience shows that most users want a portfolio of examples that they can use with a minimum of modification as templates for their own applications.

This report is intended to partially fill that need. It contains some sixty pages of text, tables, and displayed equations (the output from an IMAGEN 8/300 laser printer), together with the associated T_EX source. An enlarged copy of each page of the *Formulary*, identical with that which was photographically reduced for the actual book, is reprinted facing the T_EX source code that generated it. Users can browse through the report until they come to a table or displayed equation similar to the one they want to produce, then extract from the facing page the control sequences they need. The code is largely self-contained; however, many macros are taken from the file `PROLOG.TEX`. Whenever the user encounters an (apparently) undefined macro, its definition should be sought there. Additionally, the file `POINTSIZES.TEX` sets up all of the necessary font definitions (magnified seven-point fonts are used in place of the normal ten-point fonts in the *Formulary*, to improve readability after reduction). These files are listed immediately following this introduction.

No claim is made that the code reprinted here is optimum. We warrant only that the T_EX input produces the output you see. Wizards and other supernatural beings could possibly find more flexible, transparent, and elegant ways of printing these tables and equations. Moreover, since three different individuals participated in the project, the programming style is highly nonuniform. Worse still, we were learning as we went along, so the sections typed at the beginning of the effort are rougher than those typed after we became more proficient. We attempted to go back and clean up after ourselves; but demands of space precluded extensive commenting, especially in the complex ruled tables which most needed it. To partly make up for this, we selected two tables as examples and explained them in some detail. These are found on pages 10 and 14. In spite of this, the user will probably have to discover by experimentation why we did much of what we did. This report *docs*, however, contain a bonus that many older scientists (like the present senior author) will appreciate: a version of the *Formulary* with print big enough to read.

The *Formulary* used many T_EX tricks and shortcuts in the interests of compactness and speed of implementation. Particularly irritating, but unavoidable, are the many "hard-wired" measure-

ments. We adjusted spacing, line widths, and positions until the output looked close to our mental picture. While this kind of built-in hack is not very elegant, it is quick and easy to implement, as opposed to trying to find exactly the right macro to handle every possible case. Also, we abbreviated commands wherever possible to conserve space. Originally, each page was printed from a separate file, as shown here, with separate calls to `PROLOG.TEX`. They have since been merged into one file, but the older version was used for better readability. In a number of places, we have made characters 'active': for example, the vertical bar '|' was made to stand for two ampersands '&&' in tables. As mentioned above, most of these abbreviations are in `PROLOG.TEX`.

For convenience, users with DECNET access who can reach `LCP::` can copy the TeX file for any given page. These files, the names of which have the form `PAGExx.TEX`, where `xx` is the page number, as well as `PROLOG.TEX` and `POINTSIZE.TEX`, are currently located in the directory `LCP::SYS2:[GUEST.FORMULARY]`. A file containing the entire set of instructions used to compose the *Formulary*, `FORMULARY.TEX`, is also located in this directory. Users with Internet access can access the files through anonymous FTP (your default will be the `SYS2:[GUEST]` directory, so you need to get `[GUEST.FORMULARY]<filename>.TEX`). The host name is `NRL-LCP.ARPA`, soon to be a domain name `LCP.NRL.MIL`.

It is a pleasure to acknowledge the TeXnical assistance of Dr. Gopal Patnaik and Ken Laskey, two valuable suggestions made by Prof. Knuth, and the encouragement of Dr. Jay Boris.

```

%
% THIS FILE DEFINES THE MACROS USED THROUGHOUT THE FORMULARY.
%
% First, set up the default page sizes and magnification.
%
\magnification=1728
\hoffset=1.25truein
\voffset=1.0truein
\hsize=6.0truein
\vsize=9.0truein
\parindent=0pt
%
% Get the font definitions, and set the font to magnified sevenpoint.
%
\input pointsize
\sevenpoint
\font\headfont=cmbx5 scaled \magstep2
\font\tensorfont=cmssi8
\font\cs=cmsy7
%
% Now, set up all the commonly used macros for the various formulary pages.
%
\catcode'\!=\active
\def{&&}
%
% Special characters.
%
\def\A{{\bf A}}
\def\AOB{{\alpha/\beta}}
\def\app{\displaystyle \approx}
\def\approxlt{\kern 0.35em \raise 0.6ex \hbox{$<$} \kern -0.77em \lower 0.6ex
\hbox{$\sim$} \kern 0.35em}
\def\approxgt{\kern 0.35em \raise 0.6ex \hbox{$>$} \kern -0.77em \lower 0.6ex
\hbox{$\sim$} \kern 0.35em}
\def\B{{\bf B}}
\def\C{{\bf C}}
\def\curl{\del\times}
\def\D{{\bf D}}
\def\del{\nabla}
\def\dint{\displaystyle\int}
\def\diver{\del\cdot}
\def\doint{\displaystyle\oint}
\def\E{{\bf E}}
\def\hbar{{\mathchar'26\mskip-9mu\h}}
\def\lambdabar{{\mathchar'26\mskip-10mu\lambda}}
\def\longrightarrow{\relbar\kern-0.5pt\joinrel\rightarrow}
\def\lraf{\displaystyle \longrightarrow}
\def\part{\partial}
\def\NR{{\bf R}}
\def\T{{\tensorfont T}}
\def\tfsigma{\sigma \kern-0.55em \sigma}
\def\ts{\thinspace}
%
% Special formats.
%
\def\leftdisplay#1{\medskip\leftline{\inndent$\displaystyle#1$}\bigskip}
\def\ital#1{{\it #1/}}

```



Accession For	
NTIS GRA&I	<input checked="" type="checkbox"/>
DTIC TAB	<input type="checkbox"/>
Unannounced	<input type="checkbox"/>
Justification	
By	
Distribution/	
Availability Codes	
Dist	Avail and/or Special
A-1	

```

\def\undertext#1{${\underline{\smash{\hbox{#1}}}}$}
%
% Contractions and commonly used macros.
%
\def\bs#1{\noalign{\vskip-#1}}
\def\bsk{\bigskip}
\def\dfil{\dotfill\ }
\def\H{\hang}
\def\hang{\hangindent \oldparindent}
\def\indent{\hskip \oldparindent \spacefactor=1000}
\def\innocent{\hskip 20pt}
\def\m{${\vphantom{\big{}}}$}
\def\msk{\medskip}
\def\N{\noindent}
\def\nocorr{\kern 0pt}
\def\oldparindent{20pt}
\def\om{\omit}
\def\ov{\bar} %{\overline}
\def\ph{\phantom}
\def\s{\strut}
\def\sk{\noalign{\smallskip}}
\def\ssk{\smallskip}
\def\ablerule{\noalign{\hrule}}
\def\tf{\tensorfont}
\def\trule{\noalign{\hrule}}
\newcount\tscount
\def\tspart{||\om}
\def\tsend{&\cr}
\def\tska#1#2{\om&height#2&\om \tscount=#1 \ifcase\tscount \or\tsend
\or\tspart\tsend
\or\tspart\tspart\tsend
\or\tspart\tspart\tspart\tsend
\or\tspart\tspart\tspart\tspart\tsend
\or\tspart\tspart\tspart\tspart\tspart\tsend
\or\tspart\tspart\tspart\tspart\tspart\tspart\tsend
\or\tspart\tspart\tspart\tspart\tspart\tspart\tspart\tsend\else\bad\fi}
\def\tskb#1{height#1&\om\om\om&\om\om\om&\cr}
\def\tskc#1#2{height#2&\om \tscount=#1 \ifcase\tscount \or\tsend
\or\tspart\tsend
\or\tspart\tspart\tsend
\or\tspart\tspart\tspart\tsend
\or\tspart\tspart\tspart\tspart\tsend
\or\tspart\tspart\tspart\tspart\tspart\tspart\tsend
\or\tspart\tspart\tspart\tspart\tspart\tspart\tspart\tspart\tsend\else\bad\fi}
\def\un{\hat} %{\underline}

```



```

%
%   THIS FILE DEFINES MACROS TO SET THE POINT SIZE FOR THE FILE:
%
% \eightpoint ==> eight point type
% \ninepoint  ==> nine point type
% \tenpoint   ==> ten point type (TeX default)
% \twelvepoint ==> twelve point type
%
\font\ninerm=cmr9   \font\eightrm=cmr8   \font\sixrm=cmr6
\font\ninei=cmmi9   \font\eighti=cmmi8   \font\sixi=cmmi6
\font\ninesy=cmsy9   \font\eightsy=cmsy8   \font\sixsy=cmsy6
\font\ninebf=cmbx9   \font\eightbf=cmbx8   \font\sixbf=cmbx6
\font\ninett=cmtt9   \font\eighttt=cmtt8   \font\seventt=cmtt10
\font\nineit=cmti9   \font\eightit=cmti8   \font\sevenit=cmti7
\font\ninesl=cmsl9   \font\eightsl=cmsl8   \font\sevensl=cmsl8
\font\sevenrm=cmr7   \font\seveni=cmmi7   \font\sevenbf=cmbx7
\font\twelverm=cmr12
\font\twelvei=cmmi12
\font\twelvesy=cmsy10 scaled 1200
\font\twelvebf=cmbx12
\font\tenex=cmex10
\font\twelvett=cmtt12
\font\twelveit=cmti12
\font\twelvesl=cmsl12

\skewchar\twelvei='177
\skewchar\twelvesy='60
\hyphenchar\twelvett=-1
\skewchar\ninei='177 \skewchar\eighti='177 \skewchar\sixi='177
\skewchar\ninesy='60 \skewchar\eightsy='60 \skewchar\sixsy='60
\hyphenchar\ninett=-1 \hyphenchar\eighttt=-1 \hyphenchar\tentt=-1
\catcode'\@=11
\newskip\ttglue

\def\twelvepoint{\def\rm{\fam0\twelverm}% switch to 12 pt. type
\textfont0=\twelverm \scriptfont0=\ninerm \scriptscriptfont0=\sevenrm
\textfont1=\twelvei \scriptfont1=\ninei \scriptscriptfont1=\seveni
\textfont2=\twelvesy \scriptfont2=\ninesy \scriptscriptfont2=\sevensy
\textfont3=\tenex \scriptfont3=\tenex \scriptscriptfont3=\tenex
\textfont\itfam=\twelveit \def\it{\fam\itfam\twelveit}
\textfont\slfam\twelvesl \def\sl{\fam\slfam\twelvesl}
\textfont\ttfam\twelvett \def\tt{\fam\ttfam\twelvett}
\textfont\bffam\twelvebf \scriptfont\bffam=\ninebf
\scriptscriptfont\bffam=\sevenbf \def\bfi{\fam\bffam\twelvebf}
\tt \ttglue=.5em plus .25em minus .15em
\normalbaselineskip=15pt
\setbox\strutbox=\hbox{\vrule height9.5pt depth4.5pt width0pt}
\let\sc=\tenrm \let\big=\twelvebig \normalbaselines\rm }

\def\tenpoint{\def\rm{\fam0\tenrm}% switch to 10-point type
\textfont0=\tenrm \scriptfont0=\sevenrm \scriptscriptfont0=\fiverm
\textfont1=\teni \scriptfont1=\seveni \scriptscriptfont1=\fivei
\textfont2=\tensy \scriptfont2=\sevensy \scriptscriptfont2=\fivesy
\textfont3=\tenex \scriptfont3=\tenex \scriptscriptfont3=\tenex
\textfont\itfam=\tenit \def\it{\fam\itfam\tenit}%
\textfont\slfam=\tensl \def\sl{\fam\slfam\tensl}%
\textfont\ttfam=\tentt \def\tt{\fam\ttfam\tentt}%

```

```

\textfont\bffam=\tenbf \scriptfont\bffam=\sevenbf
\scriptscriptfont\bffam=\fivebf \def\bf{\fam\bffam\tenbf}%
\tt \ttglue=.5em plus.25em minus.15em
\normalbaselineskip=12pt
\setbox\strutbox=\hbox{\vrule height8.5pt depth3.5pt width0pt}%
\normalbaselines\rm}

```

```

\def\ninepoint{\def\rm{\fam0\ninerm}% switch to 9-point type
\textfont0=\ninerm \scriptfont0=\sixrm \scriptscriptfont0=\fiverm
\textfont1=\ninei \scriptfont1=\sixi \scriptscriptfont1=\fivei
\textfont2=\ninesy \scriptfont2=\sixsy \scriptscriptfont2=\fivesy
\textfont3=\tenex \scriptfont3=\tenex \scriptscriptfont3=\tenex
\textfont\itfam=\nineit \def\it{\fam\itfam\nineit}%
\textfont\slfam=\ninesl \def\sl{\fam\slfam\ninesl}%
\textfont\ttfam=\ninett \def\tt{\fam\ttfam\ninett}%
\textfont\bffam=\ninebf \scriptfont\bffam=\sixbf
\scriptscriptfont\bffam=\fivebf \def\bf{\fam\bffam\ninebf}%
\tt \ttglue=.5em plus.25em minus.15em
\normalbaselineskip=11pt
\setbox\strutbox=\hbox{\vrule height8pt depth3pt width0pt}%
\normalbaselines\rm}

```

```

\def\eightpoint{\def\rm{\fam0\eightrm}% switch to 8-point type
\textfont0=\eightrm \scriptfont0=\sixrm \scriptscriptfont0=\fiverm
\textfont1=\eighti \scriptfont1=\sixi \scriptscriptfont1=\fivei
\textfont2=\eightsy \scriptfont2=\sixsy \scriptscriptfont2=\fivesy
\textfont3=\tenex \scriptfont3=\tenex \scriptscriptfont3=\tenex
\textfont\itfam=\eightit \def\it{\fam\itfam\eightit}%
\textfont\slfam=\eightsl \def\sl{\fam\slfam\eightsl}%
\textfont\ttfam=\eighttt \def\tt{\fam\ttfam\eighttt}%
\textfont\bffam=\eightbf \scriptfont\bffam=\sixbf
\scriptscriptfont\bffam=\fivebf \def\bf{\fam\bffam\eightbf}%
\tt \ttglue=.5em plus.25em minus.15em
\normalbaselineskip=9pt
\setbox\strutbox=\hbox{\vrule height7pt depth2pt width0pt}%
\normalbaselines\rm}

```

```

\def\sevenpoint{\def\rm{\fam0\sevenrm}% switch to 7-point type
\textfont0=\sevenrm \scriptfont0=\fiverm \scriptscriptfont0=\fiverm
\textfont1=\seveni \scriptfont1=\fivei \scriptscriptfont1=\fivei
\textfont2=\sevensy \scriptfont2=\fivesy \scriptscriptfont2=\fivesy
\textfont3=\tenex \scriptfont3=\tenex \scriptscriptfont3=\tenex
\textfont\itfam=\sevenit \def\it{\fam\itfam\sevenit}%
\textfont\slfam=\sevensl \def\sl{\fam\slfam\sevensl}%
\textfont\ttfam=\seventt \def\tt{\fam\ttfam\seventt}%
\textfont\bffam=\sevenbf \scriptfont\bffam=\fivebf
\scriptscriptfont\bffam=\fivebf \def\bf{\fam\bffam\sevenbf}%
\tt \ttglue=.5em plus.25em minus.15em
\normalbaselineskip=8pt
\setbox\strutbox=\hbox{\vrule height7pt depth2pt width0pt}%
\normalbaselines\rm}

```

1987
REVISED

NRL PLASMA FORMULARY

DAVID L. BOOK
Laboratory for Computational Physics
Naval Research Laboratory
Washington, DC 20375

Supported by
The Office of Naval Research

```

\input prolog
\pageno=2
\centerline{\headfont CONTENTS}
\bigskip
\baselineskip=12pt
\parindent=0pt
Numerical and Algebraic \dfil \ 3 \par
Vector Identities \dfil \ 4 \par
Differential Operators in Curvilinear Coordinates \dfil \ 6 \par
Dimensions and Units \dfil 10 \par
International System (SI) Nomenclature \dfil 13 \par
Metric Prefixes \dfil 13 \par
Physical Constants (SI) \dfil 14 \par
Physical Constants (cgs) \dfil 16 \par
Formula Conversion \dfil 18 \par
Maxwell's Equations \dfil 19 \par
Electricity and Magnetism \dfil 20 \par
Electromagnetic Frequency/Wavelength Bands \dfil 21 \par
AC Circuits \dfil 22 \par
Dimensionless Numbers of Fluid Mechanics \dfil 23 \par
Shocks \dfil 26 \par
Fundamental Plasma Parameters \dfil 28 \par
Plasma Dispersion Function \dfil 30 \par
Collisions and Transport \dfil 31 \par
Approximate Magnitudes in Some Typical Plasmas \dfil 40 \par
Ionospheric Parameters \dfil 42 \par
Solar Physics Parameters \dfil 43 \par
Thermonuclear Fusion \dfil 44 \par
Relativistic Electron Beams \dfil 46 \par
Beam Instabilities \dfil 48 \par
Lasers \dfil 50 \par
Atomic Physics and Radiation \dfil 52 \par
References \dfil 58
\vfil\eject\end

```

CONTENTS

Numerical and Algebraic	3
Vector Identities	4
Differential Operators in Curvilinear Coordinates,	6
Dimensions and Units	10
International System (SI) Nomenclature	13
Metric Prefixes	13
Physical Constants (SI)	14
Physical Constants (cgs)	16
Formula Conversion	18
Maxwell's Equations,	19
Electricity and Magnetism	20
Electromagnetic Frequency/Wavelength Bands	21
AC Circuits,	22
Dimensionless Numbers of Fluid Mechanics,	23
Shocks,	26
Fundamental Plasma Parameters	28
Plasma Dispersion Function	30
Collisions and Transport	31
Approximate Magnitudes in Some Typical Plasmas	40
Ionospheric Parameters	42
Solar Physics Parameters,	43
Thermonuclear Fusion,	44
Relativistic Electron Beams,	46
Beam Instabilities	48
Lasers,	50
Atomic Physics and Radiation	52
References	58

```

\input prolog
\hoffset=1.125truein
\voffset=1.0truein
\hsize=6.25truein
\vsize=9.0truein
\pageno=3
\centerline{\headfont NUMERICAL AND ALGEBRAIC}
\bigskip
\N Gain in decibels of {\$P_2\$} relative to {\$P_1\$}

$$G = 10 \log_{10} (P_2/P_1)$$

\medskip
\N To within two percent

$$(2\pi)^{1/2} \approx 2.5; \pi^2 \approx 10; e^3 \approx 20, \\ 2^{10} \approx 10^3$$

\medskip
\N Euler-Mascheroni constant  $\gamma = 0.57722$ 
\bigskip
\N Gamma Function  $\Gamma(x+1) = x\Gamma(x)$ :

$$\Gamma(1/6) \approx 5.5663 \quad \Gamma(3/5) \approx 1.4892 \cr
\Gamma(1/5) \approx 4.5908 \quad \Gamma(2/3) \approx 1.3541 \cr
\Gamma(1/4) \approx 3.6256 \quad \Gamma(3/4) \approx 1.2254 \cr
\Gamma(1/3) \approx 2.6789 \quad \Gamma(4/5) \approx 1.1642 \cr
\Gamma(2/5) \approx 2.2182 \quad \Gamma(5/6) \approx 1.1287 \cr
\Gamma(1/2) \approx 1.7725 = \sqrt{\pi} \quad \Gamma(1) = 1$$

\medskip
\N Binomial Theorem (good for  $|x| < 1$  or  $\alpha$  positive integer):

$$(1+x)^\alpha = \sum_{k=0}^{\infty} \binom{\alpha}{k} x^k \equiv 1 + \alpha x + \frac{\alpha(\alpha-1)}{2!} x^2 + \frac{\alpha(\alpha-1)(\alpha-2)}{3!} x^3 + \dots$$

\medskip
\N Rothe-Hagen identity (good for all complex  $x, y, z$  except when singular):

$$\sum_{k=0}^{\infty} \frac{x^k}{x+kz} \binom{\alpha}{k} (x+kz)^{-\alpha} = \frac{y^{\alpha-1}}{x+y+nz} \sum_{k=0}^{\infty} \frac{y^k}{x+y+nz+kz} \binom{\alpha-1}{k} (x+y+nz+kz)^{-\alpha+1}$$

\medskip
\N Gegenbauer's summation formula (good for  $\alpha, \mu$  nonintegral,  $|z| < 1$ ):

$$\sum_{n=0}^{\infty} (-1)^n J_{\alpha-\gamma}(z) J_{\alpha+\gamma}(z) = \frac{\pi^{1/2} \Gamma(\alpha-\mu)}{\Gamma(\alpha+\mu)} J_{\alpha-\mu}(z) J_{\alpha+\mu}(z)$$

\wfil\reject\end

```

NUMERICAL AND ALGEBRAIC

Gain in decibels of P_2 relative to P_1

$$G = 10 \log_{10}(P_2/P_1).$$

To within two percent

$$(2\pi)^{1/2} \approx 2.5; \pi^2 \approx 10; e^3 \approx 20; 2^{10} \approx 10^3.$$

Euler-Mascheroni constant¹ $\gamma = 0.57722$

Gamma Function $\Gamma(x+1) = x\Gamma(x)$:

$\Gamma(1/6) = 5.5663$	$\Gamma(3/5) = 1.4892$
$\Gamma(1/5) = 4.5908$	$\Gamma(2/3) = 1.3541$
$\Gamma(1/4) = 3.6256$	$\Gamma(3/4) = 1.2254$
$\Gamma(1/3) = 2.6789$	$\Gamma(4/5) = 1.1642$
$\Gamma(2/5) = 2.2182$	$\Gamma(5/6) = 1.1288$
$\Gamma(1/2) = 1.7725 = \sqrt{\pi}$	$\Gamma(1) = 1.0$

Binomial Theorem (good for $|x| < 1$ or $\alpha =$ positive integer):

$$(1+x)^\alpha = \sum_{k=0}^{\infty} \binom{\alpha}{k} x^k \equiv 1 + \alpha x + \frac{\alpha(\alpha-1)}{2!} x^2 + \frac{\alpha(\alpha-1)(\alpha-2)}{3!} x^3 + \dots$$

Rothe-Hagen identity² (good for all complex x, y, z except when singular):

$$\sum_{k=0}^n \frac{x}{x+kz} \binom{x+kz}{k} \frac{y}{y+(n-k)z} \binom{y+(n-k)z}{n-k} = \frac{x+y}{x+y+nz} \binom{x+y+nz}{n}.$$

Newberger's summation formula³ [good for μ nonintegral, $\operatorname{Re}(\alpha+\beta) > -1$]:

$$\sum_{n=-\infty}^{\infty} \frac{(-1)^n J_{\alpha-\gamma n}(z) J_{\beta+\gamma n}(z)}{n+\mu} = \frac{\pi}{\sin \mu \pi} J_{\alpha+\gamma \mu}(z) J_{\beta-\gamma \mu}(z).$$

```

\input prolog
\hoffset=1.125truein\voffset=1truein\hsize=6.0truein\vsize=9.0truein
\pageno=4
\centerline{\headfont VECTOR IDENTITIES$^4$} \bsk
\N Notation: $f, \ g, $ are scalars; $\ A, \ B, $ etc., are vectors; $\ T $ is a
tensor; $\ {\bf I} $ is the unit dyad.
% '\T' IS SET A MAGNIFIED EIGHT-POINT CHARACTER, SINCE WE LACKED A
% SEVEN-POINT BOLD SANS-SERIF FONT. IT IS FOLLOWED BY '\ ' TO LEAVE
% A SPACE FOLLOWING.
\msk\ssk
\N\quad $\ {\bf I}^2 $ (1)
$\ A \cdot B \times C = A \times B \cdot C = B \cdot C \times A = B
\times C \cdot A = C \cdot A \times B = C \times A \cdot B $ \msk
\N\quad $\ {\bf I}^2 $ (2)
$\ A \times (B \times C) = (C \times B) \times A = (A \cdot C) B - (A \cdot B) C $ \msk
\N\quad $\ {\bf I}^2 $ (3)
$\ A \times (B \times C) + B \times (C \times A) + C \times (A \times B) = 0 $ \msk
\N\quad $\ {\bf I}^2 $ (4)
$\ (A \times B) \cdot C = (A \cdot C) (B \cdot D) -
(\ A \cdot D) (B \cdot C) $ \msk
\N\quad $\ {\bf I}^2 $ (5)
$\ (A \times B) \times (C \times D) = (A \times B \cdot C) D -
(A \times B \cdot D) C $ \msk
\N\quad $\ {\bf I}^2 $ (6) $\ {\bf I}^2 $ (7) $\ {\bf I}^2 $ (8) $\ {\bf I}^2 $ (9) $\ {\bf I}^2 $ (10)
$\ {\bf I}^2 $ (11) $\ {\bf I}^2 $ (12) $\ {\bf I}^2 $ (13) $\ {\bf I}^2 $ (14) $\ {\bf I}^2 $ (15) $\ {\bf I}^2 $ (16)
\N If $\ {\bf e}_1, \ {\bf e}_2, \ {\bf e}_3 $ are orthonormal unit vectors, a second-order
tensor $\ T $ can be written in the dyadic form
\msk
\N\quad (17) $\ T = \sum_{i,j} T_{ij} {\bf e}_i {\bf e}_j $ \msk
\N In cartesian coordinates the divergence of a tensor is a vector with
components
\msk
\N\quad (18) $\ (\text{div} T)_i = \sum_j (\partial T_{ji} / \partial x_j) $ \msk
\N [This definition is required for consistency with Eq. (29)]. In general
\msk
\N\quad (19) $\ \text{div} (A \cdot B) = (\text{div} A) \cdot B + (A \cdot \text{del}) \cdot B $ \msk
\N\quad (20) $\ \text{div} (f T) = \text{del} f \cdot T + f \text{div} T $ \msk
\vfilleject\end

```


VECTOR IDENTITIES⁴

Notation: f, g , are scalars; \mathbf{A}, \mathbf{B} , etc., are vectors; \mathbf{T} is a tensor; \mathbf{l} is the unit dyad.

- (1) $\mathbf{A} \cdot \mathbf{B} \times \mathbf{C} = \mathbf{A} \times \mathbf{B} \cdot \mathbf{C} = \mathbf{B} \cdot \mathbf{C} \times \mathbf{A} = \mathbf{B} \times \mathbf{C} \cdot \mathbf{A} = \mathbf{C} \cdot \mathbf{A} \times \mathbf{B} = \mathbf{C} \times \mathbf{A} \cdot \mathbf{B}$
- (2) $\mathbf{A} \times (\mathbf{B} \times \mathbf{C}) = (\mathbf{C} \times \mathbf{B}) \times \mathbf{A} = (\mathbf{A} \cdot \mathbf{C})\mathbf{B} - (\mathbf{A} \cdot \mathbf{B})\mathbf{C}$
- (3) $\mathbf{A} \times (\mathbf{B} \times \mathbf{C}) + \mathbf{B} \times (\mathbf{C} \times \mathbf{A}) + \mathbf{C} \times (\mathbf{A} \times \mathbf{B}) = 0$
- (4) $(\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{C} \times \mathbf{D}) = (\mathbf{A} \cdot \mathbf{C})(\mathbf{B} \cdot \mathbf{D}) - (\mathbf{A} \cdot \mathbf{D})(\mathbf{B} \cdot \mathbf{C})$
- (5) $(\mathbf{A} \times \mathbf{B}) \times (\mathbf{C} \times \mathbf{D}) = (\mathbf{A} \times \mathbf{B} \cdot \mathbf{D})\mathbf{C} - (\mathbf{A} \times \mathbf{B} \cdot \mathbf{C})\mathbf{D}$
- (6) $\nabla(fg) = \nabla(gf) = f\nabla g + g\nabla f$
- (7) $\nabla \cdot (f\mathbf{A}) = f\nabla \cdot \mathbf{A} + \mathbf{A} \cdot \nabla f$
- (8) $\nabla \times (f\mathbf{A}) = f\nabla \times \mathbf{A} + \nabla f \times \mathbf{A}$
- (9) $\nabla \cdot (\mathbf{A} \times \mathbf{B}) = \mathbf{B} \cdot \nabla \times \mathbf{A} - \mathbf{A} \cdot \nabla \times \mathbf{B}$
- (10) $\nabla \times (\mathbf{A} \times \mathbf{B}) = \mathbf{A}(\nabla \cdot \mathbf{B}) - \mathbf{B}(\nabla \cdot \mathbf{A}) + (\mathbf{B} \cdot \nabla)\mathbf{A} - (\mathbf{A} \cdot \nabla)\mathbf{B}$
- (11) $\mathbf{A} \times (\nabla \times \mathbf{B}) = (\nabla \mathbf{B}) \cdot \mathbf{A} - (\mathbf{A} \cdot \nabla)\mathbf{B}$
- (12) $\nabla(\mathbf{A} \cdot \mathbf{B}) = \mathbf{A} \times (\nabla \times \mathbf{B}) + \mathbf{B} \times (\nabla \times \mathbf{A}) + (\mathbf{A} \cdot \nabla)\mathbf{B} + (\mathbf{B} \cdot \nabla)\mathbf{A}$
- (13) $\nabla^2 f = \nabla \cdot \nabla f$
- (14) $\nabla^2 \mathbf{A} = \nabla(\nabla \cdot \mathbf{A}) - \nabla \times \nabla \times \mathbf{A}$
- (15) $\nabla \times \nabla f = 0$
- (16) $\nabla \cdot \nabla \times \mathbf{A} = 0$

If $\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3$ are orthonormal unit vectors, a second-order tensor \mathbf{T} can be written in the dyadic form

$$(17) \quad \mathbf{T} = \sum_{i,j} T_{ij} \mathbf{e}_i \mathbf{e}_j$$

In cartesian coordinates the divergence of a tensor is a vector with components

$$(18) \quad (\nabla \cdot \mathbf{T})_i = \sum_j (\partial T_{ji} / \partial x_j)$$

[This definition is required for consistency with Eq. (29)]. In general

$$(19) \quad \nabla \cdot (\mathbf{A}\mathbf{B}) = (\nabla \cdot \mathbf{A})\mathbf{B} + (\mathbf{A} \cdot \nabla)\mathbf{B}$$

$$(20) \quad \nabla \cdot (f\mathbf{T}) = \nabla f \cdot \mathbf{T} + f\nabla \cdot \mathbf{T}$$

```

\input prolog
\hoffset=1.125truein
\voffset=1truein
\hsize=6.0truein
\vsize=9.0truein
\pageno=5
\N Let  $\{\mathbf{r}\}=\{\mathbf{i}\}x+\{\mathbf{j}\}y+\{\mathbf{k}\}z$  be the radius vector of magnitude
 $r$ , from the origin to the point  $x,y,z$ . Then
\medskip
\N\quad(21)  $\operatorname{div}\{\mathbf{r}\}=3$ 
\medskip
\N\quad(22)  $\operatorname{curl}\{\mathbf{r}\}=0$ 
\medskip
\N\quad(23)  $\operatorname{del} r=\{\mathbf{r}\}/r$ 
\medskip
\N\quad(24)  $\operatorname{del}(1/r)=-\{\mathbf{r}\}/r^3$ 
\medskip
\N\quad(25)  $\operatorname{div}(\{\mathbf{r}\}/r^3)=4\pi\delta(\{\mathbf{r}\})$ 
\medskip
\N\quad(26)  $\operatorname{del}\{\mathbf{r}\} = \nabla I$ 
\medskip
\N If  $V$  is a volume enclosed by a surface  $S$  and  $d\mathbf{S}=\mathbf{n}dS$ ,
where  $\mathbf{n}$  is the unit normal outward from  $V$ ,
\medskip
\N\quad(27)  $\int_V dV \operatorname{div} f = \int_S d\mathbf{S} \cdot \mathbf{f}$ 
\medskip
\N\quad(28)  $\int_V dV \operatorname{div} \mathbf{A} = \int_S d\mathbf{S} \cdot \mathbf{A}$ 
\medskip
\N\quad(29)  $\int_V dV \operatorname{div} \nabla \phi = \int_S d\mathbf{S} \cdot \nabla \phi$ 
\medskip
\N\quad(30)  $\int_V dV \operatorname{curl} \mathbf{A} = \int_S d\mathbf{S} \times \mathbf{A}$ 
\medskip
\N\quad(31)  $\int_V dV (\nabla^2 \phi - \nabla \cdot \nabla^2 \phi) = \int_S d\mathbf{S} \cdot (\nabla \phi - \nabla \cdot \nabla \phi)$ 
\medskip
\N\quad(32)  $\int_V dV (\mathbf{A} \cdot \operatorname{curl} \operatorname{curl} \mathbf{B} - \mathbf{B} \cdot \operatorname{curl} \operatorname{curl} \mathbf{A}) =$ 
 $\int_V dV (\nabla \cdot (\mathbf{A} \times \mathbf{B}) - \nabla \cdot (\mathbf{B} \times \mathbf{A}))$ 
\medskip
\N If  $S$  is an open surface bounded by the contour  $C$ , of which the line
element is  $d\mathbf{l}$ ,
\medskip
\N\quad(33)  $\int_S d\mathbf{S} \cdot \nabla \times \mathbf{f} = \int_C d\mathbf{l} \cdot \mathbf{f}$ 
\vfill\eject\end

```

Let $\mathbf{r} = ix + jy + kz$ be the radius vector of magnitude r , from the origin to the point x, y, z . Then

$$(21) \quad \nabla \cdot \mathbf{r} = 3$$

$$(22) \quad \nabla \times \mathbf{r} = 0$$

$$(23) \quad \nabla r = \mathbf{r}/r$$

$$(24) \quad \nabla(1/r) = -\mathbf{r}/r^3$$

$$(25) \quad \nabla \cdot (\mathbf{r}/r^3) = 4\pi\delta(\mathbf{r})$$

$$(26) \quad \nabla \mathbf{r} = I$$

If V is a volume enclosed by a surface S and $d\mathbf{S} = \mathbf{n}dS$, where \mathbf{n} is the unit normal outward from V ,

$$(27) \quad \int_V dV \nabla f = \int_S d\mathbf{S} f$$

$$(28) \quad \int_V dV \nabla \cdot \mathbf{A} = \int_S d\mathbf{S} \cdot \mathbf{A}$$

$$(29) \quad \int_V dV \nabla \cdot \mathbf{T} = \int_S d\mathbf{S} \cdot \mathbf{T}$$

$$(30) \quad \int_V dV \nabla \times \mathbf{A} = \int_S d\mathbf{S} \times \mathbf{A}$$

$$(31) \quad \int_V dV (f \nabla^2 g - g \nabla^2 f) = \int_S d\mathbf{S} \cdot (f \nabla g - g \nabla f)$$

$$(32) \quad \int_V dV (\mathbf{A} \cdot \nabla \times \nabla \times \mathbf{B} - \mathbf{B} \cdot \nabla \times \nabla \times \mathbf{A}) \\ = \int_S d\mathbf{S} \cdot (\mathbf{B} \times \nabla \times \mathbf{A} - \mathbf{A} \times \nabla \times \mathbf{B})$$

If S is an open surface bounded by the contour C , of which the line element is $d\mathbf{l}$,

$$(33) \quad \int_S d\mathbf{S} \times \nabla f = \oint_C d\mathbf{l} f$$

```

\input prolog
\hoffset=1.125truein
\voffset=1truein
\hsize=6.0truein
\vsize=9.0truein
\pageno=6
\N\quad(34) $\dint_S d\{\bf S\}\cdot\curl A=\doint_C d\{\bf l\}\cdot A$
\medskip
\N\quad(35) $\dint_S(d\{\bf S\}\times\del)\times A=\doint_C d\{\bf l\}\times A$
\medskip
\N\quad(36) $\dint_S d\{\bf S\}\cdot(\del f\times\del g)=\doint_C fdg=-\doint_C
gdf$
\bigskip\bigskip
\centerline {\headfont DIFFERENTIAL OPERATORS IN}
\vskip 1pt
\centerline {\headfont CURVILINEAR COORDINATES$^5$}
\bigskip
\N{\headfont Cylindrical Coordinates}
\medskip
\N Divergence
%
% Note that the \leftdisplay macro from PROLOG.TEX is used to produce
% left-justified displayed equations.
%
\leftdisplay{\diver A={1\over r} {\partial\over{\partial r}} (rA_r) + {1\over r}
{{\partial A_\phi}\over{\partial\phi}} + {{\partial A_z}\over{\partial z}}}
\N Gradient
\leftdisplay{(\del f)_r = {{\partial f}\over{\partial r}};\quad (\del f)_\phi =
{1\over r}{{\partial f}\over{\partial\phi}};\quad (\del f)_z = {{\partial f}
\over{\partial z}}}
\N Curl
\leftdisplay{(\curl A)_1={1\over r}{{\partial A_z}\over{\partial\phi}} -
{{\partial A_\phi}\over{\partial z}}}
\leftdisplay{(\curl A)_\phi={{\partial A_r}\over{\partial z}}-{{\partial A_z}
\over{\partial r}}}
\leftdisplay{(\curl A)_z={1\over r}{{\partial\over{\partial r}} (rA_\phi) -
{1\over r} {{\partial A_r}\over{\partial\phi}}}}
\smallskip
\N Laplacian
\leftdisplay{\del^2 f = {1\over r}{{\partial\over{\partial r}}\left(r{{\partial}
f}\over{\partial r}\right)} + {1\over{r^2}}{ {\partial^2 f}\over{\partial\phi^2}
} + {{\partial^2 f}\over{\partial z^2}}}
\fil\ject\end

```

$$(34) \int_S d\mathbf{S} \cdot \nabla \times \mathbf{A} = \oint_C d\mathbf{l} \cdot \mathbf{A}$$

$$(35) \int_S (d\mathbf{S} \times \nabla) \times \mathbf{A} = \oint_C d\mathbf{l} \times \mathbf{A}$$

$$(36) \int_S d\mathbf{S} \cdot (\nabla f \times \nabla g) = \oint_C f dg = - \oint_C g df$$

DIFFERENTIAL OPERATORS IN CURVILINEAR COORDINATES⁵

Cylindrical Coordinates

Divergence

$$\nabla \cdot \mathbf{A} = \frac{1}{r} \frac{\partial}{\partial r} (r A_r) + \frac{1}{r} \frac{\partial A_\phi}{\partial \phi} + \frac{\partial A_z}{\partial z}$$

Gradient

$$(\nabla f)_r = \frac{\partial f}{\partial r}; \quad (\nabla f)_\phi = \frac{1}{r} \frac{\partial f}{\partial \phi}; \quad (\nabla f)_z = \frac{\partial f}{\partial z}$$

Curl

$$(\nabla \times \mathbf{A})_r = \frac{1}{r} \frac{\partial A_z}{\partial \phi} - \frac{\partial A_\phi}{\partial z}$$

$$(\nabla \times \mathbf{A})_\phi = \frac{\partial A_r}{\partial z} - \frac{\partial A_z}{\partial r}$$

$$(\nabla \times \mathbf{A})_z = \frac{1}{r} \frac{\partial}{\partial r} (r A_\phi) - \frac{1}{r} \frac{\partial A_r}{\partial \phi}$$

Laplacian

$$\nabla^2 f = \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial f}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 f}{\partial \phi^2} + \frac{\partial^2 f}{\partial z^2}$$

```

\input prolog
\hoffset=1.125truein
\voffset=1truein
\hsize=6.0truein
\vsize=9.0truein
\pageno=7
\N Laplacian of a vector
%
% Note that the \leftdisplay macro from PROLOG.TEX is used to
% left-justify displayed equations.
%
\leftdisplay{(\del^2 A)_r=\del^2 A_r - {2\over r^2}{(\partial A_\phi)\over
(\partial \phi)}- {A_r\over r^2}}
\leftdisplay{(\del^2 A)_\phi=\del^2 A_\phi + {2\over r^2}
{(\partial A_r)\over(\partial \phi)} - {A_\phi\over r^2}}
\leftdisplay{(\del^2 A)_z=\del^2 A_z}
\medskip
\N Components of $(\A\cdot\del)\B$
\leftdisplay{(\A\cdot\del B)_r=A_r {(\partial B_r)\over(\partial r)}+{A_\phi
\over r}{(\partial B_r)\over(\partial \phi)}+A_z{(\partial B_r)\over(\partial z)}
}-{A_\phi B_\phi\over r}}
\leftdisplay{(\A\cdot\del B)_\phi=A_r {(\partial B_\phi)\over(\partial r)}+
{A_\phi\over r}{(\partial B_\phi)\over(\partial \phi)}+A_z{(\partial B_\phi)\over
(\partial z)}+{A_\phi B_r\over r}}
\leftdisplay{(\A\cdot\del B)_z=A_r {(\partial B_z)\over(\partial r)}+{A_\phi
\over r}{(\partial B_z)\over(\partial \phi)}+A_z{(\partial B_z)\over(\partial
z)}}
\smallskip
\N Divergence of a tensor
\leftdisplay{(\diver\hbox{\tf T\!/})_r={1\over r}{(\partial\over(\partial r))
(rT_{rr})}+{1\over r}{(\partial\over(\partial \phi))T_{\phi r}}+{(\partial\over(\partial z))
T_{zr}}-{T_{\phi\phi}\over r}}
\leftdisplay{(\diver\hbox{\tf T\!/})_\phi={1\over r}{(\partial\over(\partial r))
(rT_{r\phi})}+{1\over r}{(\partial\over(\partial \phi))T_{\phi\phi}}+{(\partial\over(\partial z))
T_{z\phi}}+{(\partial\over(\partial \phi))T_{\phi r}}-{T_{\phi r}\over r}}
\leftdisplay{(\diver\hbox{\tf T\!/})_z={1\over r}{(\partial\over(\partial r))
(rT_{rz})}+{1\over r}{(\partial\over(\partial \phi))T_{\phi z}}+{(\partial\over(\partial z))
T_{zz}}-{T_{zz}\over r}}
\end{document}

```

Laplacian of a vector

$$(\nabla^2 \mathbf{A})_r = \nabla^2 A_r - \frac{2}{r^2} \frac{\partial A_\phi}{\partial \phi} - \frac{A_r}{r^2}$$

$$(\nabla^2 \mathbf{A})_\phi = \nabla^2 A_\phi + \frac{2}{r^2} \frac{\partial A_r}{\partial \phi} - \frac{A_\phi}{r^2}$$

$$(\nabla^2 \mathbf{A})_z = \nabla^2 A_z$$

Components of $(\mathbf{A} \cdot \nabla) \mathbf{B}$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_r = A_r \frac{\partial B_r}{\partial r} + \frac{A_\phi}{r} \frac{\partial B_r}{\partial \phi} + A_z \frac{\partial B_r}{\partial z} - \frac{A_\phi B_\phi}{r}$$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_\phi = A_r \frac{\partial B_\phi}{\partial r} + \frac{A_\phi}{r} \frac{\partial B_\phi}{\partial \phi} + A_z \frac{\partial B_\phi}{\partial z} + \frac{A_\phi B_r}{r}$$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_z = A_r \frac{\partial B_z}{\partial r} + \frac{A_\phi}{r} \frac{\partial B_z}{\partial \phi} + A_z \frac{\partial B_z}{\partial z}$$

Divergence of a tensor

$$(\nabla \cdot \mathbf{T})_r = \frac{1}{r} \frac{\partial}{\partial r} (r T_{rr}) + \frac{1}{r} \frac{\partial T_{\phi r}}{\partial \phi} + \frac{\partial T_{zr}}{\partial z} - \frac{T_{\phi\phi}}{r}$$

$$(\nabla \cdot \mathbf{T})_\phi = \frac{1}{r} \frac{\partial}{\partial r} (r T_{r\phi}) + \frac{1}{r} \frac{\partial T_{\phi\phi}}{\partial \phi} + \frac{\partial T_{z\phi}}{\partial z} + \frac{T_{\phi r}}{r}$$

$$(\nabla \cdot \mathbf{T})_z = \frac{1}{r} \frac{\partial}{\partial r} (r T_{rz}) + \frac{1}{r} \frac{\partial T_{\phi z}}{\partial \phi} + \frac{\partial T_{zz}}{\partial z}$$

```

\input prolog
\hoffset=1.125truein\voffset=1truein\hsize=6.0truein\vsize=9.0truein
\pageno=8
\N{\headfont Spherical Coordinates} \msk
\N Divergence
%
% Note that the \leftdisplay macro from PROLOG.TEX is used to
% left-justify displayed equations.
%
\leftdisplay{\diver A={1\over r^2}}{\part\over{\part r}}(r^2 A_r) +
{1\over r\sin\theta}}{\part\over{\part\theta}}(\sin\theta A_\theta) +
{1\over r\sin\theta}}{\part A_\phi}\over{\part\phi}}}

\N Gradient
\leftdisplay{(\del f)_r={\part f}\over{\part r}};\quad
(\del f)_\theta={1\over r} {\part f}\over{\part\theta}};\quad
(\del f)_\phi={1\over r\sin\theta} {\part f}\over{\part\phi}}}

\N Curl
\leftdisplay{(\curl A)_r={1\over r\sin\theta}}{\part\over{\part\theta}}
(\sin\theta A_\phi)-{1\over r\sin\theta}}{\part
A_\theta}\over{\part\phi}}}
\leftdisplay{(\curl A)_\theta={1\over r\sin\theta}}{\part
A_r}\over{\part\phi}}-{1\over r}{\part\over{\part r}}(rA_\phi))
\leftdisplay{(\curl A)_\phi={1\over r}{\part\over{\part r}}
(rA_\theta)-{1\over r}{\part A_r}\over{\part\theta}}} \msk

\N Laplacian
\leftdisplay{\del^2 f={1\over r^2}}{\part\over{\part r}}\left(r^2
{\part f}\over{\part r}\right)+{1\over r^2\sin\theta}}{\part
\over{\part\theta}}\left(\sin\theta{\part f}\over{\part\theta}}
\right)+{1\over r^2\sin^2\theta}}{\part^2 f}\over{\part\phi^2}}}

\N Laplacian of a vector
\leftdisplay{(\del^2 A)_r=\del^2 A_r-{2A_r}\over{r^2}}-{2\over{r^2}}{\part
A_\theta}\over{\part\theta}}-{2\cot\theta A_\theta}\over{r^2}}-
{2\over{r^2\sin\theta}}{\part A_\phi}\over{\part\phi}}}
\leftdisplay{(\del^2 A)_\theta=\del^2 A_\theta+{2\over{r^2}}{\part A_r}\over
{\part\theta}}-{A_\theta}\over{r^2\sin^2\theta}}-{2\cos\theta}\over
{r^2\sin^2\theta}}{\part A_\phi}\over{\part\phi}}}
\leftdisplay{(\del^2 A)_\phi=\del^2 A_\phi-{A_\phi}\over{r^2\sin^2\theta}}+
{2\over{r^2\sin\theta}}{\part A_r}\over{\part\phi}}+{2\cos\theta}\over
{r^2\sin^2\theta}}{\part A_\theta}\over{\part\phi}}}
\vf{1}\eject\end

```


Spherical Coordinates

Divergence

$$\nabla \cdot \mathbf{A} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 A_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta A_\theta) + \frac{1}{r \sin \theta} \frac{\partial A_\phi}{\partial \phi}$$

Gradient

$$(\nabla f)_r = \frac{\partial f}{\partial r}; \quad (\nabla f)_\theta = \frac{1}{r} \frac{\partial f}{\partial \theta}; \quad (\nabla f)_\phi = \frac{1}{r \sin \theta} \frac{\partial f}{\partial \phi}$$

Curl

$$(\nabla \times \mathbf{A})_r = \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta A_\phi) - \frac{1}{r \sin \theta} \frac{\partial A_\theta}{\partial \phi}$$

$$(\nabla \times \mathbf{A})_\theta = \frac{1}{r \sin \theta} \frac{\partial A_r}{\partial \phi} - \frac{1}{r} \frac{\partial}{\partial r} (r A_\phi)$$

$$(\nabla \times \mathbf{A})_\phi = \frac{1}{r} \frac{\partial}{\partial r} (r A_\theta) - \frac{1}{r} \frac{\partial A_r}{\partial \theta}$$

Laplacian

$$\nabla^2 f = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial f}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial f}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 f}{\partial \phi^2}$$

Laplacian of a vector

$$(\nabla^2 \mathbf{A})_r = \nabla^2 A_r - \frac{2A_r}{r^2} - \frac{2}{r^2} \frac{\partial A_\theta}{\partial \theta} - \frac{2 \cot \theta A_\theta}{r^2} - \frac{2}{r^2 \sin \theta} \frac{\partial A_\phi}{\partial \phi}$$

$$(\nabla^2 \mathbf{A})_\theta = \nabla^2 A_\theta + \frac{2}{r^2} \frac{\partial A_r}{\partial \theta} - \frac{A_\theta}{r^2 \sin^2 \theta} - \frac{2 \cos \theta}{r^2 \sin^2 \theta} \frac{\partial A_\phi}{\partial \phi}$$

$$(\nabla^2 \mathbf{A})_\phi = \nabla^2 A_\phi - \frac{A_\phi}{r^2 \sin^2 \theta} + \frac{2}{r^2 \sin \theta} \frac{\partial A_r}{\partial \phi} + \frac{2 \cos \theta}{r^2 \sin^2 \theta} \frac{\partial A_\theta}{\partial \phi}$$

Components of $(\mathbf{A} \cdot \nabla)\mathbf{B}$

$$(\mathbf{A} \cdot \nabla)\mathbf{B})_r = A_r \frac{\partial B_r}{\partial r} + \frac{A_\theta}{r} \frac{\partial B_r}{\partial \theta} + \frac{A_\phi}{r \sin \theta} \frac{\partial B_r}{\partial \phi} - \frac{A_\theta B_\theta + A_\phi B_\phi}{r}$$

$$(\mathbf{A} \cdot \nabla)\mathbf{B})_\theta = A_r \frac{\partial B_\theta}{\partial r} + \frac{A_\theta}{r} \frac{\partial B_\theta}{\partial \theta} + \frac{A_\phi}{r \sin \theta} \frac{\partial B_\theta}{\partial \phi} + \frac{A_\theta B_r}{r} - \frac{\cot \theta A_\phi B_\phi}{r}$$

$$(\mathbf{A} \cdot \nabla)\mathbf{B})_\phi = A_r \frac{\partial B_\phi}{\partial r} + \frac{A_\theta}{r} \frac{\partial B_\phi}{\partial \theta} + \frac{A_\phi}{r \sin \theta} \frac{\partial B_\phi}{\partial \phi} + \frac{A_\phi B_r}{r} + \frac{\cot \theta A_\phi B_\theta}{r}$$

Divergence of a tensor

$$(\nabla \cdot \mathbf{T})_r = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 T_{rr}) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta T_{\theta r}) \\ + \frac{1}{r \sin \theta} \frac{\partial T_{\phi r}}{\partial \phi} - \frac{T_{\theta\theta} + T_{\phi\phi}}{r}$$

$$(\nabla \cdot \mathbf{T})_\theta = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 T_{r\theta}) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta T_{\theta\theta}) \\ + \frac{1}{r \sin \theta} \frac{\partial T_{\phi\theta}}{\partial \phi} + \frac{T_{\theta r}}{r} - \frac{\cot \theta T_{\phi\phi}}{r}$$

$$(\nabla \cdot \mathbf{T})_\phi = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 T_{r\phi}) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta T_{\theta\phi}) \\ + \frac{1}{r \sin \theta} \frac{\partial T_{\phi\phi}}{\partial \phi} + \frac{T_{\phi r}}{r} + \frac{\cot \theta T_{\phi\theta}}{r}$$


```

\tska{7}{1ex}
\s|\hfil Quantity |\hfil bol |\om\hfil SI \hfil |\om\hfil Gaussian
%
% \tska is a macro used as a tableskip. It skips a variable distance
% vertically in a table, continuing all rules, in order to improve the
% spacing, and is adjusted so that the spacing is pleasing to the eye.
% At the beginning of each line is the macro \m or \s. These are struts,
% invisible hboxes of fixed height, which determine the height of each
% line. \m is slightly taller than \s. Again, they were chosen by eye.
% \bs is defined as \noalign{\vskip-#1} in prolog.tex. It enables us
% to "back up" in tables.
%
\hfil | \hfil Units | \hfil Factor | \hfil Units &\cr
%
% The preamble automatically puts an \hfil to the right of each entry,
% so we need add \hfil only on the left, to center them.
%
\tska{7}{7pt} \trule \tska{7}{1pt} \trule \tska{7}{2pt}
%
% We skip down, leaving a double rule beneath the table header. The
% body of the table follows:
%
\m| Capacitance \hidewidth |  $\text{\$C\$ } \{t^{-2}q^{-2} \over ml^{-2}\} | 1 |$ 
farad |  $\text{\$9\texttimes10^{-11}\$} | \text{cm} & \backslash\text{cr}$ 
%
% Note the use of \hidewidth. TeX allocated more space than necessary
% if we allowed it to position the text itself, so we use \hidewidth to
% conceal the length of the longer entries. This prevents "overfull
% hbox" errors.
%
\m| Charge |  $\text{\$q\$ } | q | \{m^{-1/2}l^{-3/2} \over t\} | \text{coulomb}$ 
\hidewidth |  $\text{\$3\texttimes10^{-9}\$} | \text{statcoulomb} \backslash\text{hidewidth} & \backslash\text{cr} \backslash\text{tska}{7}{2pt}$ 
\s| Charge |  $\text{\$}\rho\text{\$ } | \{q \over l^{-3}\} | \{m^{-1/2} \over l^{-3/2}t\} |$ 
coulomb \hidewidth |  $\text{\$3\texttimes10^{-3}\$} | \text{statcoulomb} \backslash\text{hidewidth} & \backslash\text{cr} \backslash\text{bs}{1.5ex}$ 
\s| \quad density | | | \quad /m $\text{\$^{-3}\$} | | \quad /\text{cm}\text{\$^{-3}\$} & \backslash\text{cr}$ 
%
% Charge density is spread out over two line . \bs{1.5ex} is used
% to position the two lines more closely.
%
\m| Conductance \hidewidth | |  $\{tq^{-2} \over ml^{-2}\} | \{1 \over t\} |$ 
siemens |  $\text{\$9\texttimes10^{-11}\$} | \text{cm/sec} & \backslash\text{cr} \backslash\text{tska}{7}{2pt}$ 
\m| Conductivity |  $\text{\$}\sigma\text{\$ } | \{tq^{-2}\over ml^{-3}\} | 1\over t |$ 
siemens |  $\text{\$9\texttimes10^{-9}\$} | \text{sec}\text{\$^{-1}\$} & \backslash\text{cr} \backslash\text{bs}{2.0ex}$ 
\s| | | | \quad /m | | & \backslash\text{cr}
\m| Current |  $\text{\$I,i\$ } | q\over t | \{m^{-1/2}l^{-3/2}\}\over t^{-2}|$ 
ampere |  $\text{\$3\texttimes10^{-9}\$} | \text{statampere} & \backslash\text{cr} \backslash\text{tska}{7}{2pt}$ 
\s| Current |  $\text{\$}\{\bf J\},\{\bf j\}\text{\$ } | q\over l^{-2}t\} | \{m^{-1/2}\} \over$ 
 $\{l^{-1/2}t^{-2}\} | \text{ampere} | \text{\$3\texttimes10^{-5}\$} | \text{statampere} & \backslash\text{cr} \backslash\text{bs}{1.5ex}$ 

```

```

\s| \quad density | | | |\quad /m\${}^2\$ | | \quad /cm\${}^2\$ & \cr
\m| Density | $\rho\$ | m\over{1^3} | m\over{1^3} | kg/m\${}^3\$ |
    10^{-3}\$ | g/cm\${}^3\$ & \cr
\s| Displacement \hidewidth | {\bf D} | q\over{1^2} | {m^{-1/2}}
    \over {1^{-1/2}t} | coulomb \hidewidth | $12\pi\backslashtimes10^{-5}\$ \hidewidth |
    statcoulomb \hidewidth & \cr \bs{1.75ex}
\s| | | |\quad /m\${}^2\$ | | \quad /cm\${}^2\$ & \cr
\m| Electric field | {\bf E} | {m}\over{t^2q} | {m^{-1/2}}\over{1^{-1/2}t} |
    volt/m | $\displaystyle{1\over3}\times10^{-4}\$ | statvolt/cm & \cr \tska{7}{2pt}
\m| Electro- | {\cs E}, | {m}^2\over{t^2q} | {m^{-1/2}}1^{-1/2}}
    \over t | volt | $\displaystyle{1\over3}\times10^{-2}\$ | statvolt & \cr
    \bs{2.0ex}
\s| \quad motance |Emf | | | | & \cr
\m| Energy | $U,W\$ \hidewidth | {m}^2\over{t^2} | {m}^2 \over
    {t^2} | joule | $10^{-7}\$ | erg & \cr
\m| Energy | $w,\epsilon\$ | m\over{1t^2} | m\over{1 t^2} |
    joule/m\${}^3\$ \hidewidth | $10\$ | erg/cm\${}^3\$ & \cr \bs{2.0ex}
\s| \quad density | | | | | & \cr \tska{7}{2pt} \cr \trule}}$$
%
% End of Table.
%
\vfill\ject\end

```

DIMENSIONS AND UNITS

To get the value of a quantity in Gaussian units, multiply the value expressed in SI units by the conversion factor. Multiples of 3 in the conversion factors result from approximating the speed of light $c = 2.9979 \times 10^{10}$ cm/sec $\approx 3 \times 10^{10}$ cm/sec.

Physical Quantity	Sym- bol	Dimensions		SI Units	Conversion Factor	Gaussian Units
		SI	Gaussian			
Capacitance	C	$\frac{t^2 q^2}{ml^2}$	l	farad	9×10^{11}	cm
Charge	q	q	$\frac{m^{1/2} l^{3/2}}{t}$	coulomb	3×10^9	statcoulomb
Charge density	ρ	$\frac{q}{l^3}$	$\frac{m^{1/2}}{l^{3/2} t}$	coulomb /m ³	3×10^3	statcoulomb /cm ³
Conductance		$\frac{t q^2}{ml^2}$	$\frac{l}{t}$	siemens	9×10^{11}	cm/sec
Conductivity	σ	$\frac{t q^2}{ml^3}$	$\frac{1}{t}$	siemens /m	9×10^9	sec ⁻¹
Current	I, i	$\frac{q}{t}$	$\frac{m^{1/2} l^{3/2}}{t^2}$	ampere	3×10^9	statampere
Current density	\mathbf{J}, \mathbf{j}	$\frac{q}{l^2 t}$	$\frac{m^{1/2}}{l^{1/2} t^2}$	ampere /m ²	3×10^5	statampere /cm ²
Density	ρ	$\frac{m}{l^3}$	$\frac{m}{l^3}$	kg/m ³	10^{-3}	g/cm ³
Displacement	\mathbf{D}	$\frac{q}{l^2}$	$\frac{m^{1/2}}{l^{1/2} t}$	coulomb /m ²	$12\pi \times 10^5$	statcoulomb /cm ²
Electric field	\mathbf{E}	$\frac{ml}{t^2 q}$	$\frac{m^{1/2}}{l^{1/2} t}$	volt/m	$\frac{1}{3} \times 10^{-4}$	statvolt/cm
Electro- motive force	$\mathcal{E},$ Emf	$\frac{ml^2}{t^2 q}$	$\frac{m^{1/2} l^{1/2}}{t}$	volt	$\frac{1}{3} \times 10^{-2}$	statvolt
Energy	U, W	$\frac{ml^2}{t^2}$	$\frac{ml^2}{t^2}$	joule	10^7	erg
Energy density	w, ϵ	$\frac{m}{lt^2}$	$\frac{m}{lt^2}$	joule/m ³	10	erg/cm ³

```

\input prolog \pageno=11
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\centerline{\ } \nointerlineskip
$$\vbox{\tabskip=0pt \offinterlineskip
\halign to\hsize{&\vrule#\tabskip=0.25em plus1em&\hfil&\vrule#&
#\hfil&\vrule#&\$ \displaystyle{#} \$\hfil&\vrule#&\$ \displaystyle{#} \$\hfil&\vrule#
&\hfil&\vrule#&\hfil&\vrule#&\hfil&\vrule#\tabskip=0pt\cr\bs{0.2truein}\trule
\om & height2pt &\om | \om | \multispan3 | \om | \om | \om & \cr
\s|\om | \om | \multispan3\hfil Dimensions \hfil | \om | \om | \om & \cr
\s | \hfil Physical | \hfil Sym- \hidewidth | \multispan3 | \hfil SI |
\hfil Conversion \hidewidth | \hfil Gaussian & \cr \bs{1.5ex}
\noalign{\moveright1.91truein\vbox{\hrule width1.45truein}} \tska{7}{1.0ex}
\s | \hfil Quantity | \hfil bol | \om \hfil SI \hfil | \om \hfil
Gaussian \hfil | \hfil Units | \hfil Factor | \hfil Units & \cr
\tska{7}{7pt} \trule \tska{7}{1.0pt} \trule \tska{7}{2pt}
\m| Force | {\bf F} | {m}\over{t^2} | {m}\over{t^2} |
newton | $10^{-5}$ | dyne & \cr
\m| Frequency | $f$, \nu$ | 1\over t | 1\over t | hertz | 1 | hertz & \cr
\m| Impedance | $Z$ | {m}^2\over{tq^2} | t\over l | ohm|
$ \displaystyle{1\over 9}\times 10^{-11}$ \hidewidth | sec/cm & \cr \tska{7}{2pt}
\m| Inductance | $L$ | {m}^2\over{q^2} | {t^2}\over l | henry |
$ \displaystyle{1\over 9}\times 10^{-11}$ \hidewidth | sec^2/cm & \cr \tska{7}{2pt}
\s| Length | $l$ | 1 | 1 | meter (m) | $10^{-2}$ |
centimeter \hidewidth & \cr \tska{7}{2pt} \bs{1.0ex}
\s| | | | | \quad (cm) & \cr
\s| Magnetic | {\bf H} | q\over{lt} | {m^{1/2}}\over{l^{1/2}t} |
ampere-- | $4\pi\times 10^{-3}$ \hidewidth | oersted & \cr \bs{1.5ex}
\s| \quad intensity | | | | \quad turn/m | | & \cr
\m| Magnetic flux | $\Phi$ | {m}^2\over{tq} | {m^{1/2}}l^{3/2}}
\over t | weber | $10^{-8}$ | maxwell & \cr \tska{7}{2pt}
\m| Magnetic | {\bf B} | m\over{tq} | {m^{1/2}}\over{l^{1/2}t} |
tesla | $10^{-4}$ | gauss & \cr \bs{1.75ex}
\s| \quad induction | | | | | & \cr
\m| Magnetic | $m$, \mu$ | {l^2 q}\over t | {m^{1/2}}l^{5/2}}\over t |
ampere--m$^2$ \hidewidth | $10^{-3}$ | oersted-- & \cr \bs{1.75ex}
\s| \quad moment | | | | | \quad cm$^3$ & \cr
\s| Magnetization \hidewidth | {\bf M} | q\over{lt} | {m^{1/2}}\over
{l^{1/2}t} | ampere-- \hidewidth | $10^{-3}$ | oersted & \cr \bs{1.75ex}
\s| | | | | \quad turn/m | | & \cr
\m| Magneto- | {\bf M}, | q\over t | {m^{1/2}}l^{1/2}}\over{t^2} |
ampere-- | $ \displaystyle{4\pi\over 10}$ | gilbert & \cr \bs{2.0ex}
\s| \quad motance | Hmf | | | \quad turn | | & \cr
\m| Mass | $m$, M$ \hidewidth | m | m | kilogram | $10^{-3}$ | gram (g)& \cr \bs{1.75ex}
\s| | | | | \quad (kg) | | & \cr
\m| Momentum | {\bf p}, {\bf P} | {m}\over t | {m}\over t |
kg--m/s | $10^{-5}$ | g--cm/sec & \cr
\m| Momentum | | m\over{l^2 t} | m\over{l^2 t} | kg/m$^2$--s |
$10^{-1}$ | g/cm$^2$--sec & \cr \bs{1.75ex}
\s| \quad density | | | | | & \cr
\m| Permeability | $\mu$ | {m}\over{q^2} | 1 | henry/m |
$ \displaystyle{1\over 4\pi}\times 10^{-7}$ \ph{C}$ | \quad --- & \cr
\tska{7}{2pt} \trule}} $\vfil\egroup\end

```


Physical Quantity	Sym- bol	Dimensions		SI Units	Conversion Factor	Gaussian Units
		SI	Gaussian			
Force	F	$\frac{ml}{t^2}$	$\frac{ml}{t^2}$	newton	10^5	dyne
Frequency	f, ν	$\frac{1}{t}$	$\frac{1}{t}$	hertz	1	hertz
Impedance	Z	$\frac{ml^2}{tq^2}$	$\frac{t}{l}$	ohm	$\frac{1}{9} \times 10^{-11}$	sec/cm
Inductance	L	$\frac{ml^2}{q^2}$	$\frac{t^2}{l}$	henry	$\frac{1}{9} \times 10^{-11}$	sec ² /cm
Length	l	l	l	meter (m)	10^2	centimeter (cm)
Magnetic intensity	H	$\frac{q}{lt}$	$\frac{m^{1/2}}{l^{1/2}t}$	ampere- turn/m	$4\pi \times 10^{-3}$	oersted
Magnetic flux	Φ	$\frac{ml^2}{tq}$	$\frac{m^{1/2}l^{3/2}}{t}$	weber	10^8	maxwell
Magnetic induction	B	$\frac{m}{tq}$	$\frac{m^{1/2}}{l^{1/2}t}$	tesla	10^4	gauss
Magnetic moment	m, μ	$\frac{l^2q}{t}$	$\frac{m^{1/2}l^{5/2}}{t}$	ampere·m ²	10^3	oersted· cm ³
Magnetization	M	$\frac{q}{lt}$	$\frac{m^{1/2}}{l^{1/2}t}$	ampere- turn/m	10^{-3}	oersted
Magneto- motance	$\mathcal{M},$ Mmf	$\frac{q}{t}$	$\frac{m^{1/2}l^{1/2}}{t^2}$	ampere turn	$\frac{4\pi}{10}$	gilbert
Mass	m, M	m	m	kilogram (kg)	10^3	gram (g)
Momentum	p, P	$\frac{ml}{t}$	$\frac{ml}{t}$	kg m/s	10^5	g cm/sec
Momentum density		$\frac{m}{l^2t}$	$\frac{m}{l^2t}$	kg/m ² s	10^{-1}	g/cm ² sec
Permeability	μ	$\frac{ml}{q^2}$	1	henry/m	$\frac{1}{4\pi} \times 10^7$	

```

\input prolog \pageno=12
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\centerline{\ } \nointerlineskip
$$\vbox{\tabskip=0pt \offinterlineskip
\halign to\hsize{#\vrule#\tabskip=0.25em plus1em&#\hfil&\vrule#&
#\hfil&\vrule#&#\displaystyle{#}$\hfil&\vrule#&#\displaystyle{#}$\hfil&\vrule#
&#\hfil&\vrule#&#\hfil&\vrule#&#\hfil&\vrule#\tabskip=0pt\cr\bs{0.2truein}\trule
\om & height2pt & \om | \cm | \multispan3 | \om | \om | \om & \cr
\s| \om | \om | \multispan3 \hfil Dimensions \hfil | \om | \om | \om & \cr
\s | \hfil Physical | \hfil Sym- \hidewidth | \multispan3 | \hfil SI |
\hfil Conversion \hidewidth | \hfil Gaussian & \cr
\bs{1.5ex}\noalign{\moveright1.69truein\vbox{\hrule width1.45truein}}\tska{7}{1.0ex}
\s | \hfil Quantity | \hfil bol | \om \hfil SI \hfil | \om \hfil
Gaussian \hfil | \hfil Units | \hfil Factor | \hfil Units & \cr
\tska{7}{7pt} \trule \tska{7}{1.0pt} \trule \tska{7}{2pt}
\m| Permittivity | $\epsilon$ | {t^2q^2}\over{ml^3} | 1 | farad/m |
$36\pi\times 10^{-9}$ | \quad --- & \cr
\m| Polarization | {\bf P} | q\over{l^2} | {m^{-1/2}}\over{l^{-1/2}t} |
coulomb/m$^2$ \hidewidth | $3\times 10^{-5}$ | statcoulomb\hidewidth & \cr \bs{1.5ex}
\s| | | | | \quad /cm$^2$ & \cr
\m| Potential | $V, \phi$ \hidewidth | {ml^2}\over{t^2q} | {m^{-1/2}l^{-1/2}}
\over t | volt | $\displaystyle{1\over 3}\times 10^{-2}$ | statvolt & \cr \tska{7}{3pt}
\m| Power | $P$ | {ml^2}\over{t^3} | {ml^2}\over{t^3} | watt | $10^{-7}$ |
erg/sec & \cr
\m| Power | | m\over{l t^3} | m\over{l t^3} | watt/m$^3$ | 10 |
erg/cm$^3$--sec \hidewidth & \cr \bs{1.75ex}
\s| \quad density | | | | | & \cr
\m| Pressure | $p, P$ \hidewidth | m\over{l t^2} | m\over{l t^2} |
pascal | 10 | dyne/cm$^2$ & \cr \tska{7}{2pt}
\s| Reluctance | {\cs R} | {q^2}\over{ml^2} | 1\over l |
ampere--turn \hidewidth | $4\pi\times 10^{-9}$ \hidewidth | cm$^{-1}$ & \cr \bs{1.5ex}
\s| | | | | \quad /weber | | & \cr
\m| Resistance | $R$ | {ml^2}\over{tq^2} | t\over l | ohm |
$\displaystyle{1\over 9}\times 10^{-11}$ \hidewidth | sec/cm & \cr \tska{7}{3pt}
\m| Resistivity | $\eta, \rho$ | {ml^3}\over{tq^2} | t | ohm--m |
$\displaystyle{1\over 9}\times 10^{-9}$ | sec & \cr \tska{7}{3pt}
\s| Thermal con- \hidewidth | $\kappa, k$ | {ml}\over{t^3} |
{ml}\over{t^3} | watt/m-- | $10^{-5}$ | erg/cm--sec-- \hidewidth & \cr \bs{1.5ex}
\s| \quad ductivity | | | | \quad deg (K) $\phi$ | | \quad
deg (K) $\phi$ & \cr \tska{7}{2pt}
\s| Time | $t$ | t | t | second (s) | 1 | second (sec) & \cr
\m| Vector | {\bf A} | {ml}\over{tq} | {m^{-1/2}l^{-1/2}}\over t |
weber/m | $10^{-6}$ | gauss--cm & \cr \bs{2.0ex}
\s| \quad potential | | | | | & \cr
\m| Velocity | {\bf v} | 1\over t | 1\over t | m/s | $10^{-2}$ | cm/sec & \cr
\m| Viscosity | $\eta, \mu$ | m\over{l t} | m\over{l t} | kg/m--s | 10 |
poise & \cr \tska{7}{2pt}
\m| Vorticity | $\zeta$ | 1\over t | 1\over t | s$^{-1}$ | 1 | sec$^{-1}$ & \cr
\m| Work | $W$ | {ml^2}\over{t^2} | {ml^2}\over{t^2} | joule | $10^{-7}$ |
erg & \cr \tska{7}{2pt} \trule}}$ \vfil\ eject\end

```

Physical Quantity	Symbol	Dimensions		SI Units	Conversion Factor	Gaussian Units
		SI	Gaussian			
Permittivity	ϵ	$\frac{t^2 q^2}{ml^3}$	1	farad/m	$36\pi \times 10^9$	
Polarization	\mathbf{P}	$\frac{q}{l^2}$	$\frac{m^{1/2}}{l^{1/2}t}$	coulomb/m ²	3×10^5	statcoulomb/cm ²
Potential	V, ϕ	$\frac{ml^2}{t^2 q}$	$\frac{m^{1/2} l^{1/2}}{t}$	volt	$\frac{1}{3} \times 10^{-2}$	statvolt
Power	P	$\frac{ml^2}{t^3}$	$\frac{ml^2}{t^3}$	watt	10^7	erg/sec
Power density		$\frac{m}{lt^3}$	$\frac{m}{lt^3}$	watt/m ³	10	erg/cm ³ sec
Pressure	p, P	$\frac{m}{lt^2}$	$\frac{m}{lt^2}$	pascal	10	dyne/cm ²
Reluctance	\mathcal{R}	$\frac{q^2}{ml^2}$	$\frac{1}{l}$	ampere turn/weber	$4\pi \times 10^{-9}$	cm ⁻¹
Resistance	R	$\frac{ml^2}{tq^2}$	$\frac{t}{l}$	ohm	$\frac{1}{9} \times 10^{-11}$	sec/cm
Resistivity	η, ρ	$\frac{ml^3}{tq^2}$	t	ohm m	$\frac{1}{9} \times 10^{-9}$	sec
Thermal conductivity	κ, k	$\frac{ml}{t^3}$	$\frac{ml}{t^3}$	watt/m deg (K)	10^5	erg/cm sec deg (K)
Time	t	t	t	second (s)	1	second (sec)
Vector potential	\mathbf{A}	$\frac{ml}{tq}$	$\frac{m^{1/2} l^{1/2}}{t}$	weber/m	10^6	gauss cm
Velocity	\mathbf{v}	$\frac{l}{t}$	$\frac{l}{t}$	m/s	10^2	cm/sec
Viscosity	η, μ	$\frac{m}{lt}$	$\frac{m}{lt}$	kg/m s	10	poise
Vorticity	ζ	$\frac{1}{t}$	$\frac{1}{t}$	s ⁻¹	1	sec ⁻¹
Work	W	$\frac{ml^2}{t^2}$	$\frac{ml^2}{t^2}$	joule	10^7	erg

The listing for page 13 begins on the next page.

```

\input prolog
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein
\pageno=13
\centerline{\headfont INTERNATIONAL SYSTEM (SI) NOMENCLATURE$^6$}
\nointerlineskip
$$\vbox{\tabskip=Opt \offinterlineskip \def\S{({$ph{*}$})}
  % \S is used to leave the space of an asterisk, so that text lines up.
\align to \hsize{\vrule# \tabskip=0.5em plus2em minus0.5em&\hfil\strut&\vrule#
&\hfil&\vrule##&\hfil#\hfil&\vrule#\tabskip=Opt\hskip1.0pt&\vrule#
\tabskip=0.5em plus2em minus1em&\hfil\strut&\vrule##&\hfil&\vrule#
&\hfil#\hfil&\vrule#\tabskip=Opt\cr \trule \tskb{2pt} \bs{1pt}
&\hfil Physical\hfil Name\Symbol\
&\hfil Physical\hfil Name\Symbol&\cr \bs{1pt}
&\hfil Quantity\hfil of Unit\hfil for Unit\
&\hfil Quantity\hfil of Unit\hfil for Unit&\cr \tskb{2pt} \bs{1pt} \trule
\tskb{1.0pt} \trule \tskb{2pt} \bs{1pt}
&*length\meter|m|&electric\volt|V&\cr \bs{2pt}
&\om\om\om|&potential||&\cr \bs{2pt}
&*mass\kilogram|kg|&\om\om\om&\cr \bs{2pt}
&\om\om\om|&electric|ohm|$\Omega$&\cr \bs{2pt}
&*time\second|s|&resistance|\om\om&\cr \tskb{2pt}
&*current\ampere|A|&electric|siemens|S&\cr \bs{2pt}
&\om\om\om|&conductance|\om\om&\cr \bs{2pt}
&*temperature\kelvin|K|&\om\om\om&\cr \bs{2pt}
&\om\om\om|&electric|farad|F&\cr \bs{2pt}
&*amount of\mole|mol|&capacitance|\om\om&\cr \bs{2pt}
&\S substance|\om\om\om|&\om\om\om&\cr \bs{2pt} \bs{2pt}
&\om\om\om|&magnetic flux|weber|Wb&\cr \bs{2pt}
&*luminous\candela|cd|&\om\om\om&\cr \bs{2pt}
&\S intensity|\om\om\om|&magnetic|henry|H&\cr \bs{2pt}
&\om\om\om|&inductance|\om\om&\cr \bs{2pt}
&\dag plane angle\radian|rad|&\om\om\om&\cr \bs{2pt}
&\om\om\om|&magnetic|tesla|T&\cr \bs{2pt}
&\dag solid angle\steradian|sr|&intensity|\om\om&\cr \tskb{2pt}
&\S frequency|hertz|Hz|&luminous flux|lumen|lm&\cr \tskb{2pt}
&\S energy|joule|J|&illuminance|lux|lx&\cr \tskb{2pt}
&\S force|newton|N|&activity (of a|becquerel|Bq&\cr \bs{2pt}
&\om\om\om|&radioactive|\om\om&\cr \bs{2pt}
&\S pressure|pascal|Pa|&source)|\om\om&\cr \tskb{2pt}
&\S power|watt|W|&absorbed dose|gray|Gy&\cr \bs{2pt}
&\om\om\om|&(of ionizing)|\om\om&\cr \bs{2pt}
&\S electric charge|coulomb|C|&radiation)|\om\om&\cr \bs{1pt}
height{2pt}\om\om\om|&\om\om\om&\cr \trule}}
\skip=4pt
*SI base unit \quad \dag Supplementary Unit
%
% Second table on page 13
%
\skip\centerline{\headfont METRIC PREFIXES}
\nointerlineskip
$$\vbox{\tabskip=Opt\offinterlineskip \def\hw{\hidewidth}
  % THE MACRO \hw IS JUST USED TO SAVE SPACE AND TYPING.
\align to \hsize{\vrule#\tabskip=1em plus2em minus1em&\quad&\hfil&\vrule#
&\hfil&\vrule##&\hfil#\hfil&\strut&\vrule#\tabskip=Opt\hskip1.0pt&\vrule#
\tabskip=1em plus2em minus1em&\quad&\hfil&\vrule##&\hfil&\vrule#
&\hfil#\hfil&\vrule#\tabskip=Opt\cr \trule \tskb{2pt}
&\om\hfil\hw Multiple\hw\hfil\hfil\hw Prefix\hw\hw Symbol\hw\

```

```

&\om\hfil\hw Multiple\hw\hfil|\hfil\hw Prefix\hw|\hw Symbol\hw&\cr
\tskb{2pt} \trule \tskb{1.0pt} \trule \tskb{2pt}
&$10^{-1}$|deci|d|&$10$|deca|da&\cr \tskb{1pt}
&$10^{-2}$|centi|c|&$10^{-2}$|hecto|h&\cr \tskb{1pt}
&$10^{-3}$|milli|m|&$10^{-3}$|kilo|k&\cr \tskb{1pt}
&$10^{-6}$|micro|$\mu$|&$10^{-6}$|mega|M&\cr \tskb{1pt}
&$10^{-9}$|nano|n|&$10^{-9}$|giga|G&\cr \tskb{1pt}
&$10^{-12}$|pico|p|&$10^{-12}$|tera|T&\cr \tskb{1pt}
&$10^{-15}$|femto|f|&$10^{-15}$|peta|P&\cr \tskb{1pt}
&$10^{-18}$|atto|a|&$10^{-18}$|exa|E&\cr \tskb{2pt} \trule}}$$
\vfill\eject\end

```

INTERNATIONAL SYSTEM (SI) NOMENCLATURE⁶

Physical Quantity	Name of Unit	Symbol for Unit	Physical Quantity	Name of Unit	Symbol for Unit
*length	meter	m	electric potential	volt	V
*mass	kilogram	kg	electric resistance	ohm	Ω
*time	second	s	electric conductance	siemens	S
*current	ampere	A	electric capacitance	farad	F
*temperature	kelvin	K	magnetic flux	weber	Wb
*amount of substance	mole	mol	magnetic inductance	henry	H
*luminous intensity	candela	cd	magnetic intensity	tesla	T
†plane angle	radian	rad	luminous flux	lumen	lm
†solid angle	steradian	sr	illuminance	lux	lx
frequency	hertz	Hz	activity (of a radioactive source)	becquerel	Bq
energy	joule	J	absorbed dose (of ionizing radiation)	gray	Gy
force	newton	N			
pressure	pascal	Pa			
power	watt	W			
electric charge	coulomb	C			

*SI base unit

†Supplementary Unit

METRIC PREFIXES

Multiple	Prefix	Symbol	Multiple	Prefix	Symbol
10^{-1}	deci	d	10	deca	da
10^{-2}	centi	c	10^2	hecto	h
10^{-3}	milli	m	10^3	kilo	k
10^{-6}	micro	μ	10^6	mega	M
10^{-9}	nano	n	10^9	giga	G
10^{-12}	pico	p	10^{12}	tera	T
10^{-15}	femto	f	10^{15}	peta	P
10^{-18}	atto	a	10^{18}	exa	E

The listing for page 14 begins on the next page.


```

& \quad\quad free space | | | & \cr \tskc{4}{2pt}
& Proton/electron mass |  $\{m_p\}/\{m_e\}$  |  $1.8362\times 10^{-3}$  | & \cr
& \quad\quad ratio | | | & \cr \tskc{4}{2pt}
& Electron charge/mass |  $e/\{m_e\}$  |  $1.7588\times 10^{11}$  | C$,kg$^{-1}$ & \cr
& \quad\quad ratio | | | & \cr \tskc{4}{2pt}
& Rydberg constant |  $R_\infty=\displaystyle\frac{me^4}{8\epsilon_0^2 h^3}$  |  $1.0974\times 10^{-7}$  | m$^{-1}$ & \cr \tskc{4}{2pt}
& Bohr radius |  $a_0=\epsilon_0 h^2/\pi me^2$  |  $5.2918\times 10^{-11}$  | m & \cr
& \tskc{4}{2pt}
& Atomic cross section |  $\pi a_0^2$  |  $8.7974\times 10^{-21}$  | m$^2$ & \cr
& \tskc{4}{2pt}
& Classical electron radius |  $r_e=e^2/4\pi\epsilon_0 mc^2$  |  $2.8179\times 10^{-15}$  | m & \cr \tskc{4}{2pt}
& Thomson cross section |  $(8\pi/3)r_e^2$  |  $6.6524\times 10^{-29}$  | m$^2$ & \cr
& \tskc{4}{2pt}
& Compton wavelength of |  $h/\{m_{ec}\}$  |  $2.4263\times 10^{-12}$  | m & \cr
& \quad\quad electron |  $\hbar/\{m_{ec}\}$  |  $3.8616\times 10^{-13}$  | m
& \cr \tskc{4}{2pt}
& Fine-structure constant |  $\alpha=e^2/2\epsilon_0 h c$  |
&  $7.2974\times 10^{-3}$  | & \cr
& !  $\alpha^{-1}$  |  $137.04$  | & \cr \tskc{4}{2pt}
& First radiation constant |  $c_1=2\pi hc^2$  |  $3.7418\times 10^{-2}$  |
& W$,m$^2$ & \cr \tskc{4}{2pt}
& Second radiation |  $c_2=hc/k$  |  $1.4388\times 10^{-2}$  | m$,K & \cr
& \quad\quad constant | | | & \cr \tskc{4}{2pt}
& Stefan-Boltzmann |  $\sigma$  |  $5.6703\times 10^{-8}$  |
& W$,m$^{-2}$K$^{-4}$ & \cr
& \quad\quad constant | | | & \cr \tskc{4}{2pt} \trule}}$$
% END OF RULED TABLE.
\vfil\eject\end

```

PHYSICAL CONSTANTS (SI)⁷

Physical Quantity	Symbol	Value	Units
Boltzmann constant	k	1.3807×10^{-23}	J K^{-1}
Elementary charge	e	1.6022×10^{-19}	C
Electron mass	m_e	9.1095×10^{-31}	kg
Proton mass	m_p	1.6726×10^{-27}	kg
Gravitational constant	G	6.6726×10^{-11}	$\text{m}^3 \text{s}^{-2} \text{kg}^{-1}$
Planck constant	h	6.6262×10^{-34}	J s
	$\hbar = h/2\pi$	1.0546×10^{-34}	J s
Speed of light in vacuum	c	2.9979×10^8	m s^{-1}
Permittivity of free space	ϵ_0	8.8542×10^{-12}	F m^{-1}
Permeability of free space	μ_0	$4\pi \times 10^{-7}$	H m^{-1}
Proton/electron mass ratio	m_p/m_e	1.8362×10^3	
Electron charge/mass ratio	e/m_e	1.7588×10^{11}	C kg^{-1}
Rydberg constant	$R_\infty = \frac{me^4}{8\epsilon_0^2 ch^3}$	1.0974×10^7	m^{-1}
Bohr radius	$a_0 = \epsilon_0 h^2 / \pi m e^2$	5.2918×10^{-11}	m
Atomic cross section	πa_0^2	8.7974×10^{-21}	m^2
Classical electron radius	$r_e = e^2 / 4\pi \epsilon_0 m c^2$	2.8179×10^{-15}	m
Thomson cross section	$(8\pi/3)r_e^2$	6.6524×10^{-29}	m^2
Compton wavelength of electron	$h/m_e c$	2.4263×10^{-12}	m
	$\hbar/m_e c$	3.8616×10^{-13}	m
Fine-structure constant	$\alpha = e^2 / 2\epsilon_0 h c$	7.2974×10^{-3}	
	α^{-1}	137.04	
First radiation constant	$c_1 = 2\pi h c^2$	3.7418×10^{-2}	W m^2
Second radiation constant	$c_2 = hc/k$	1.4388×10^{-2}	m K
Stefan-Boltzmann constant	σ	5.6703×10^{-8}	$\text{W m}^{-2} \text{K}^{-4}$

```

\input prolog
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\pageno=15
$$\vbox{\offinterlineskip \def\quad{\hskip 4pt} \def\quid{\hskip 1pt} \hrule
%
% The definitions of \quad and \quid change from table to table in order
% to keep overall table widths the same.
%
\halign {\&\vrule# &\quad #\hfil \quid &\vrule# &\strut \quad #\hfil \quid
&\vrule# &\quad #\hfil \quid &\vrule# &\quad #\hfil \quid &\vrule# \cr \tskc{4}{2pt}
&\hfil Physical Quantity | \hfil Symbol | \hfil Value | \hfil Units & \cr
\tskc{4}{2pt} \trule \tskc{4}{1.0pt} \trule \tskc{4}{2pt}
& Wavelength associated |  $\lambda_0 = hc/e$  |  $1.2399 \times 10^{-6}$  | m &
\cr & \quad \quad with 1 eV | | | \cr \tskc{4}{2pt}
& Frequency associated |  $\nu_0 = e/h$  |  $2.4180 \times 10^{14}$  | Hz & \cr
& \quad \quad with 1 eV | | | \cr \tskc{4}{2pt}
& Wave number associated |  $k_0 = e/hc$  |  $8.0655 \times 10^5$  |  $m^{-1}$  & \cr
& \quad \quad with 1 eV | | | \cr \tskc{4}{2pt}
& Energy associated with |  $h\nu_0$  |  $1.6022 \times 10^{-19}$  | J & \cr
& \quad \quad 1 eV | | | \cr \tskc{4}{2pt}
& Energy associated with |  $hc$  |  $1.9865 \times 10^{-25}$  | J & \cr
& \quad \quad 1  $m^{-1}$  | | | \cr \tskc{4}{2pt}
& Energy associated with |  $\frac{3}{8}(\epsilon_0)^{-2} h^2$  | 13.606 | eV & \cr
& \quad \quad 1 Rydberg | | | \cr \tskc{4}{2pt}
& Energy associated with |  $k/e$  |  $8.6173 \times 10^{-5}$  | eV & \cr
& \quad \quad 1 Kelvin | | | \cr \tskc{4}{2pt}
& Temperature associated |  $e/k$  |  $1.1605 \times 10^4$  | K & \cr
& \quad \quad with 1 eV | | | \cr \tskc{4}{2pt}
& Avogadro number |  $N_A$  |  $6.0220 \times 10^{23}$  |  $mol^{-1}$  & \cr \tskc{4}{2pt}
& Faraday constant |  $F=N_Ae$  |  $9.6485 \times 10^4$  | C\$,  $mol^{-1}$  & \cr \tskc{4}{2pt}
& Gas constant |  $R=N_Ak$  | \hfil 8.3144 | J\$,  $K^{-1}mol^{-1}$  & \cr
\tskc{4}{2pt}
& Loschmidt's number |  $n_0$  |  $2.6868 \times 10^{25}$  |  $m^{-3}$  & \cr
& \quad \quad (no. density at STP) | | | \cr \tskc{4}{2pt}
& Atomic mass unit |  $m_u$  |  $1.6606 \times 10^{-27}$  | kg & \cr \tskc{4}{2pt}
& Standard temperature |  $T_0$  | \hfil 273.16 | K & \cr \tskc{4}{2pt}
& Atmospheric pressure |  $p_0=n_0kT_0$  |  $1.0133 \times 10^5$  | Pa & \cr \tskc{4}{2pt}
& Pressure of 1 mm Hg | | |  $1.3332 \times 10^2$  | Pa & \cr
& \quad \quad (1 torr) | | | \cr \tskc{4}{2pt}
& Molar volume at STP |  $V_0=RT_0/p_0$  |  $2.2415 \times 10^{-2}$  |  $m^3$  & \cr
\tskc{4}{2pt}
& Molar weight of air |  $M_{(air)}$  |  $2.8971 \times 10^{-2}$  | kg & \cr \tskc{4}{2pt}
& calorie (cal) | | \hfil 4.1868 | J & \cr \tskc{4}{2pt}
& Gravitational |  $g$  | \hfil 9.8067 |  $m\$, s^{-2}$  & \cr
& \quad \quad acceleration | | | \cr \tskc{4}{2pt}} \hrule}$$
\vfil\ject\end

```

Physical Quantity	Symbol	Value	Units
Wavelength associated with 1 eV	$\lambda_0 = hc/e$	1.2399×10^{-6}	m
Frequency associated with 1 eV	$\nu_0 = e/h$	2.4180×10^{14}	Hz
Wave number associated with 1 eV	$k_0 = e/hc$	8.0655×10^5	m^{-1}
Energy associated with 1 eV	$h\nu_0$	1.6022×10^{-19}	J
Energy associated with 1 m^{-1}	hc	1.9865×10^{-25}	J
Energy associated with 1 Rydberg	$me^3/8\epsilon_0^2 h^2$	13.606	eV
Energy associated with 1 Kelvin	k/e	8.6173×10^{-5}	eV
Temperature associated with 1 eV	e/k	1.1605×10^4	K
Avogadro number	N_A	6.0220×10^{23}	mol^{-1}
Faraday constant	$F = N_A e$	9.6485×10^4	C mol^{-1}
Gas constant	$R = N_A k$	8.3144	$\text{J K}^{-1} \text{mol}^{-1}$
Loschmidt's number (no. density at STP)	n_0	2.6868×10^{25}	m^{-3}
Atomic mass unit	m_u	1.6606×10^{-27}	kg
Standard temperature	T_0	273.16	K
Atmospheric pressure	$p_0 = n_0 k T_0$	1.0133×10^5	Pa
Pressure of 1 mm Hg (1 torr)		1.3332×10^2	Pa
Molar volume at STP	$V_0 = RT_0/p_0$	2.2415×10^{-2}	m^3
Molar weight of air	M_{air}	2.8971×10^{-2}	kg
calorie (cal)		4.1868	J
Gravitational acceleration	g	9.8067	m s^{-2}

```

\input prolog
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\pageno=16
\centerline{\headfont PHYSICAL CONSTANTS (cgs)$^7$}
$$\vbox{\offinterlineskip \def\quad{\hskip 3pt} \def\quid{\hskip 0pt} \hrule
%
% The definitions of \quad and \quid change from table to table in order
% to keep overall table widths the same.
%
\halign {\&\vrule# &\quad #\hfil \quid &\vrule# &\strut \quad #\hfil \quid
&\vrule# &\quad #\hfil \quid &\vrule# &\quad #\hfil \quid &\vrule# \cr \tskc{4}{2pt}
&\hfil Physical Quantity | \hfil Symbol | \hfil Value | \hfil Units &\cr \tskc{4}{2pt}
\trule
height1.0pt &\om | \om | \om | \cm &\cr
\trule \tskc{4}{2pt}
& Boltzmann constant | $k$ | $1.3807\times 10^{-16}$ | erg/deg$, $(K)$ &\cr \tskc{4}{2pt}
& Elementary charge | $e$ | $4.8032\times 10^{-10}$ | statcoulomb &\cr
& | | \quad (statcoul) &\cr \tskc{4}{2pt}
& Electron mass | $m_e$ | $9.1095\times 10^{-28}$ | g &\cr \tskc{4}{2pt}
& Proton mass | $m_p$ | $1.6726\times 10^{-24}$ | g &\cr \tskc{4}{2pt}
& Gravitational constant | $G$ | $6.6720\times 10^{-8}$ |
dyne-cm$^2$/g$^2$ &\cr \tskc{4}{2pt}
& Planck constant | $h$ | $6.6262\times 10^{-27}$ | erg-sec &\cr
& | $\hbar=h/2\pi$ | $1.0546\times 10^{-27}$ | erg-sec &\cr \tskc{4}{2pt}
& Speed of light in vacuum | $c$ | $2.9979\times 10^{10}$ | cm/sec &\cr \tskc{4}{2pt}
& Proton/electron mass | $\{m_p\}/\{m_e\}$ | $1.8362\times 10^3$ | &\cr
& \quad\quad ratio | | &\cr \tskc{4}{2pt}
& Electron charge/mass | $e/\{m_e\}$ | $5.2728\times 10^{17}$ |
statcoul/g &\cr
& \quad\quad ratio | | &\cr \tskc{4}{2pt}
& Rydberg constant | $R_\infty=\displaystyle{\{2\pi^2me^4\}\over\{ch^3\}}$ |
$1.0974\times 10^5$ | cm$^{-1}$ &\cr \tskc{4}{2pt}
& Bohr radius | $a_0=\hbar^2/\{me^2\}$ | $5.2918\times 10^{-9}$ | cm &\cr \tskc{4}{2pt}
& Atomic cross section | $\pi \{a_0\}^2$ | $8.7974\times 10^{-17}$ |
cm$^2$ &\cr \tskc{4}{2pt}
& Classical electron radius | $r_e=e^2/\{mc^2\}$ | $2.8179\times 10^{-13}$ |
cm &\cr \tskc{4}{2pt}
& Thomson cross section | $(8\pi/3)\{r_e\}^2$ | $6.6524\times 10^{-25}$ |
cm$^2$ &\cr \tskc{4}{2pt}
& Compton wavelength of | $h/\{m_{ec}\}$ | $2.4263\times 10^{-10}$ | cm &\cr
& \quad\quad electron | $\hbar/\{m_{ec}\}$ | $3.8616\times 10^{-11}$ | cm &\cr \tskc{4}{2pt}
& Fine-structure constant | $\alpha=e^2/\hbar c$ | $7.2974\times 10^{-3}$ | &\cr
& | $\alpha^{-1}$ | \hfil $137.04$ | &\cr \tskc{4}{2pt}
& First radiation constant | $c_1=2\pi hc^2$ | $3.7418\times 10^{-5}$ |
erg-cm$^2$/sec &\cr \tskc{4}{2pt}
& Second radiation | $c_2=hc/k$ | \hfil $1.4388$ | cm-deg$, $(K)$ &\cr
& \quad\quad constant | | &\cr \tskc{4}{2pt}
& Stefan-Boltzmann | $\sigma$ | $5.6703\times 10^{-5}$ | erg/cm$^2$- &\cr
& \quad\quad constant | | \quad sec-deg$^4$ &\cr \tskc{4}{2pt}
& Wavelength associated | $\lambda_{0\$} | $1.2399\times 10^{-4}$ | cm &\cr
& \quad\quad with 1 eV | | &\cr \tskc{4}{2pt} \trule}}$$
\vfil\ejct\end

```

PHYSICAL CONSTANTS (cgs)⁷

Physical Quantity	Symbol	Value	Units
Boltzmann constant	k	1.3807×10^{-16}	erg/deg (K)
Elementary charge	e	4.8032×10^{-10}	statcoulomb (statcoul)
Electron mass	m_e	9.1095×10^{-28}	g
Proton mass	m_p	1.6726×10^{-24}	g
Gravitational constant	G	6.6720×10^{-8}	dyne-cm ² /g ²
Planck constant	h	6.6262×10^{-27}	erg-sec
	$\hbar = h/2\pi$	1.0546×10^{-27}	erg-sec
Speed of light in vacuum	c	2.9979×10^{10}	cm/sec
Proton/electron mass ratio	m_p/m_e	1.8362×10^3	
Electron charge/mass ratio	e/m_e	5.2728×10^{17}	statcoul/g
Rydberg constant	$R_\infty = \frac{2\pi^2 m e^4}{ch^3}$	1.0974×10^5	cm ⁻¹
Bohr radius	$a_0 = \hbar^2/m_e e^2$	5.2918×10^{-9}	cm
Atomic cross section	πa_0^2	8.7974×10^{-17}	cm ²
Classical electron radius	$r_e = e^2/m_e c^2$	2.8179×10^{-13}	cm
Thomson cross section	$(8\pi/3)r_e^2$	6.6524×10^{-25}	cm ²
Compton wavelength of electron	$h/m_e c$	2.4263×10^{-10}	cm
	$\hbar/m_e c$	3.8616×10^{-11}	cm
Fine-structure constant	$\alpha = e^2/\hbar c$	7.2974×10^{-3}	
	α^{-1}	137.04	
First radiation constant	$c_1 = 2\pi\hbar c^2$	3.7418×10^{-5}	erg-cm ² /sec
Second radiation constant	$c_2 = \hbar c/k$	1.4388	cm-deg (K)
Stefan-Boltzmann constant	σ	5.6703×10^{-5}	erg/cm ² - sec-deg ⁴
Wavelength associated with 1 eV	λ_0	1.2399×10^{-4}	cm

```

\input prolog
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\pageno=17
$$\vbox{\offinterlineskip \def\quad{\hskip 4pt} \def\quid{\hskip 1pt} \hrule
%
% The definitions of \quad and \quid change from table to table in order to
% keep overall table widths the same.
%
\halign {\&\vrule# &\quad #\hfil \quid &\vrule# &\strut \quad #\hfil \quid
&\vrule# &\quad #\hfil \quid &\vrule# &\quad #\hfil \quid &\vrule# \cr \tskc{4}{2pt}
&\hfil Physical Quantity | \hfil Symbol | \hfil Value | \hfil Units &\cr \tskc{4}{2pt}
\trule
height1.0pt &\om | \om | \om | \om &\cr
\trule \tskc{4}{2pt}
& Frequency associated | $\nu_0$ | $2.4180\times 10^{-14}$ | Hz &\cr
&\quad\quad with 1 eV | | | &\cr \tskc{4}{2pt}
& Wave number associated | $k_0$ | $8.0655\times 10^{-3}$ | cm$^{-1}$ &\cr
&\quad\quad with 1 eV | | | &\cr \tskc{4}{2pt}
& Energy associated with | | $1.6022\times 10^{-12}$ | erg &\cr
&\quad\quad 1 eV | | | &\cr \tskc{4}{2pt}
& Energy associated with | | $1.9865\times 10^{-16}$ | erg &\cr
&\quad\quad 1 cm$^{-1}$ | | | &\cr \tskc{4}{2pt}
& Energy associated with | | \hfil 13.606 | eV &\cr
&\quad\quad 1 Rydberg | | | &\cr \tskc{4}{2pt}
& Energy associated with | | $8.6173\times 10^{-5}$ | eV &\cr
&\quad\quad 1 deg Kelvin | | | &\cr \tskc{4}{2pt}
& Temperature associated | | $1.1605\times 10^{-4}$ | deg$, $(K)$ &\cr
&\quad\quad with 1 eV | | | &\cr \tskc{4}{2pt}
& Avogadro number | $N_A$ | $6.0220\times 10^{23}$ | mol$^{-1}$ &\cr \tskc{4}{2pt}
& Faraday constant | $F=N_Ae$ | $2.8925\times 10^{14}$ |
statcoul/mol &\cr \tskc{4}{2pt}
& Gas constant | $R=N_Ak$ | $8.3144\times 10^{-7}$ | erg/deg-mol &\cr \tskc{4}{2pt}
& Loschmidt's number | $n_0$ | $2.6868\times 10^{19}$ | cm$^{-3}$ &\cr
&\quad\quad (no. density at STP) | | | &\cr \tskc{4}{2pt}
& Atomic mass unit | $m_u$ | $1.6606\times 10^{-24}$ | g &\cr \tskc{4}{2pt}
& Standard temperature | $T_0$ | \hfil 273.16 | deg$, $(K)$ &\cr \tskc{4}{2pt}
& Atmospheric pressure | $p_0=n_0kT_0$ | $1.0133\times 10^6$ |
dyne/cm$^2$ &\cr \tskc{4}{2pt}
& Pressure of 1 mm Hg | | $1.3332\times 10^{-3}$ | dyne/cm$^2$ &\cr
&\quad\quad (1 torr) | | | &\cr \tskc{4}{2pt}
& Molar volume at STP | $V_0=RT_0/p_0$ | $2.2415\times 10^{-4}$ | cm$^3$ &\cr \tskc{4}{2pt}
& Molar weight of air | $M_{\rm air}$ | \hfil 28.971 | g &\cr \tskc{4}{2pt}
& calorie (cal) | | $4.1868\times 10^{-7}$ | erg &\cr \tskc{4}{2pt}
& Gravitational | $g$ | \hfil 980.67 | cm/sec$^2$ &\cr
&\quad\quad acceleration | | | &\cr \tskc{4}{2pt}} \hrule}$$
\wfil\select\end

```


Physical Quantity	Symbol	Value	Units
Frequency associated with 1 eV	ν_0	2.4180×10^{14}	Hz
Wave number associated with 1 eV	k_0	8.0655×10^3	cm^{-1}
Energy associated with 1 eV		1.6022×10^{-12}	erg
Energy associated with 1 cm^{-1}		1.9865×10^{-16}	erg
Energy associated with 1 Rydberg		13.606	eV
Energy associated with 1 deg Kelvin		8.6173×10^{-5}	eV
Temperature associated with 1 eV		1.1605×10^4	deg (K)
Avogadro number	N_A	6.0220×10^{23}	mol^{-1}
Faraday constant	$F = N_A e$	2.8925×10^{14}	statcoul/mol
Gas constant	$R = N_A k$	8.3144×10^7	erg/deg-mol
Loschmidt's number (no. density at STP)	n_0	2.6868×10^{19}	cm^{-3}
Atomic mass unit	m_u	1.6606×10^{-24}	g
Standard temperature	T_0	273.16	deg (K)
Atmospheric pressure	$p_0 = n_0 k T_0$	1.0133×10^6	dyne/cm ²
Pressure of 1 mm Hg (1 torr)		1.3332×10^3	dyne/cm ²
Molar volume at STP	$V_0 = RT_0/p_0$	2.2415×10^4	cm ³
Molar weight of air	M_{air}	28.971	g
calorie (cal)		4.1868×10^7	erg
Gravitational acceleration	g	980.67	cm/sec ²

FORMULA CONVERSION⁸

Here $\alpha = 10^2 \text{ cm m}^{-1}$, $\beta = 10^7 \text{ erg J}^{-1}$, $\epsilon_0 = 8.8542 \times 10^{-12} \text{ F m}^{-1}$, $\mu_0 = 4\pi \times 10^{-7} \text{ H m}^{-1}$, $c = (\epsilon_0 \mu_0)^{-1/2} = 2.9979 \times 10^8 \text{ m s}^{-1}$, and $\hbar = 1.0546 \times 10^{-34} \text{ J s}$. To derive a dimensionally correct SI formula from one expressed in Gaussian units, substitute for each quantity according to $\bar{Q} = \bar{k}Q$, where \bar{k} is the coefficient in the second column of the table corresponding to Q (overbars denote variables expressed in Gaussian units). Thus, the formula $\bar{a}_0 = \bar{\hbar}^2 / \bar{m} \bar{e}^2$ for the Bohr radius becomes $\alpha a_0 = (\hbar \beta)^2 / [(m \beta / \alpha^2)(e^2 \alpha \beta / 4\pi \epsilon_0)]$, or $a_0 = \epsilon_0 \hbar^2 / \pi m e^2$. To go from SI to natural units in which $\hbar = c = 1$ (distinguished by a circumflex), use $Q = \hat{k}^{-1} \hat{Q}$, where \hat{k} is the coefficient corresponding to Q in the third column. Thus $\hat{a}_0 = 4\pi \epsilon_0 \hbar^2 / [(\hat{m} \hbar / c)(\hat{e}^2 \epsilon_0 \hbar c)] = 4\pi / \hat{m} \hat{e}^2$. (In transforming from SI units, do not substitute for ϵ_0 , μ_0 , or c .)

Physical Quantity	Gaussian Units to SI	Natural Units to SI
Capacitance	$\alpha / 4\pi \epsilon_0$	ϵ_0^{-1}
Charge	$(\alpha \beta / 4\pi \epsilon_0)^{1/2}$	$(\epsilon_0 \hbar c)^{-1/2}$
Charge density	$(\beta / 4\pi \alpha^5 \epsilon_0)^{1/2}$	$(\epsilon_0 \hbar c)^{-1/2}$
Current	$(\alpha \beta / 4\pi \epsilon_0)^{1/2}$	$(\mu_0 / \hbar c)^{1/2}$
Current density	$(\beta / 4\pi \alpha^3 \epsilon_0)^{1/2}$	$(\mu_0 / \hbar c)^{1/2}$
Electric field	$(4\pi \beta \epsilon_0 / \alpha^3)^{1/2}$	$(\epsilon_0 / \hbar c)^{1/2}$
Electric potential	$(4\pi \beta \epsilon_0 / \alpha)^{1/2}$	$(\epsilon_0 / \hbar c)^{1/2}$
Electric conductivity	$(4\pi \epsilon_0)^{-1}$	ϵ_0^{-1}
Energy	β	$(\hbar c)^{-1}$
Energy density	β / α^3	$(\hbar c)^{-1}$
Force	β / α	$(\hbar c)^{-1}$
Frequency	1	c^{-1}
Inductance	$4\pi \epsilon_0 / \alpha$	μ_0^{-1}
Length	α	1
Magnetic induction	$(4\pi \beta / \alpha^3 \mu_0)^{1/2}$	$(\mu_0 \hbar c)^{-1/2}$
Magnetic intensity	$(4\pi \mu_0 \beta / \alpha^3)^{1/2}$	$(\mu_0 / \hbar c)^{1/2}$
Mass	β / α^2	c / \hbar
Momentum	β / α	\hbar^{-1}
Power	β	$(\hbar c^2)^{-1}$
Pressure	β / α^3	$(\hbar c)^{-1}$
Resistance	$4\pi \epsilon_0 / \alpha$	$(\epsilon_0 / \mu_0)^{1/2}$
Time	1	c
Velocity	α	c^{-1}

```

\input prolog
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein
\pageno=19
\centerline{\headfont MAXWELL'S EQUATIONS}

$$\frac{d}{dt} \oint \mathbf{B} \cdot d\mathbf{l} = - \frac{d}{dt} \int \mathbf{E} \cdot d\mathbf{A}$$

\trule \tskc{3}{2pt}
&\hfil Name or Description\om\hfil SI\hfil\om\hfil Gaussian\hfil&\cr
\tskc{3}{2pt} \trule \tskc{3}{1pt} \trule \tskc{3}{2.5pt}
&Faraday's law\del\times E=-{\partial\B\over\partial t}\del\times E
=-{1\over c}{\partial\B\over\partial t}\cr \tskc{3}{2pt} \tskc{3}{2pt}
&Ampere's law\del\times {\bf H}={\partial\B\over\partial t}+{\bf J}\del\times
{\bf H}={1\over c}{\partial\B\over\partial t}+{4\pi\over c}{\bf J}\cr \tskc{3}{4pt}
&Poisson equation\del\cdot D=\rho\del\cdot D=4\pi\rho\cr \tskc{3}{2pt}
&[Absence of magnetic\del\cdot B=0\del\cdot B=0\cr \bs{2pt}
&\quad monopoles]\om\om\cr \tskc{3}{2pt} \bs{2pt}
&Lorentz force onq\left(E+{\bf v}\times\B\right)q\left(E+{1\over c}{\bf v}
\times\B\right)\cr \bs{2pt} \bs{2pt} \bs{2pt}
&\quad charge q\om\om\cr \tskc{3}{2pt}
&Constitutive\B=\epsilon E\B=\epsilon E\cr \bs{1pt}
&\quad relations\B=\mu{\bf H}\B=\mu{\bf H}\cr \tskc{3}{2pt} \trule}}$$
\vskip-6pt
In a plasma,  $\mu \approx \mu_0 = 4\pi \times 10^{-7}$ ,  $\{\rm H\}$ ,  $\{\rm m\}^{-1}$ 
(Gaussian units:  $\mu \approx 1$ ). The permittivity satisfies
 $\epsilon \approx \epsilon_0 = 8.8542 \times 10^{-12}$ ,  $\{\rm F\}$ ,  $\{\rm m\}^{-1}$ 
(Gaussian:  $\epsilon \approx 1$ )$ provided that all charge is regarded as free.
Using the drift approximation  $\mathbf{v}_{\perp} = \mathbf{E} \times \mathbf{B} / B^2$  to calculate
polarization charge density gives rise to a dielectric constant
 $K \equiv \epsilon / \epsilon_0 = 1 + 36\pi \times 10^{-9} \rho / B^2$  (SI)  $= 1 + 4\pi \rho / c^2 B^2$  (Gaussian), where  $\rho$  is the mass density.

\indent
The electromagnetic energy in volume  $V$  is given by

$$W = \frac{1}{2} \int_V dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D})$$

(rm SI)\ph{\rm Gaussian.}\cr

$$W = \frac{1}{8\pi} \int_V dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D})$$
 (rm Gaussian).
\ph{\rm SI}\cr}}

Poynting's theorem is

$$\frac{\partial W}{\partial t} + \oint_S \mathbf{H} \cdot d\mathbf{S} = - \int_V dV \mathbf{J} \cdot \mathbf{E}$$

where  $S$  is the closed surface bounding  $V$  and the Poynting vector (energy
flux across  $S$ ) is given by  $\mathbf{N} = \mathbf{E} \times \mathbf{H}$  (SI) or  $\mathbf{N} = c \mathbf{E} \times \mathbf{H} / 4\pi$  (Gaussian).
\vfill\ejct\end

```

MAXWELL'S EQUATIONS

Name or Description	SI	Gaussian
Faraday's law	$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$	$\nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}$
Ampere's law	$\nabla \times \mathbf{H} = \frac{\partial \mathbf{D}}{\partial t} + \mathbf{J}$	$\nabla \times \mathbf{H} = \frac{1}{c} \frac{\partial \mathbf{D}}{\partial t} + \frac{4\pi}{c} \mathbf{J}$
Poisson equation	$\nabla \cdot \mathbf{D} = \rho$	$\nabla \cdot \mathbf{D} = 4\pi\rho$
[Absence of magnetic monopoles]	$\nabla \cdot \mathbf{B} = 0$	$\nabla \cdot \mathbf{B} = 0$
Lorentz force on charge q	$q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$	$q\left(\mathbf{E} + \frac{1}{c} \mathbf{v} \times \mathbf{B}\right)$
Constitutive relations	$\mathbf{D} = \epsilon \mathbf{E}$ $\mathbf{B} = \mu \mathbf{H}$	$\mathbf{D} = \epsilon \mathbf{E}$ $\mathbf{B} = \mu \mathbf{H}$

In a plasma, $\mu \approx \mu_0 = 4\pi \times 10^{-7} \text{ H m}^{-1}$ (Gaussian units: $\mu \approx 1$). The permittivity satisfies $\epsilon \approx \epsilon_0 = 8.8542 \times 10^{-12} \text{ F m}^{-1}$ (Gaussian: $\epsilon \approx 1$) provided that all charge is regarded as free. Using the drift approximation $\mathbf{v}_\perp = \mathbf{E} \times \mathbf{B}/B^2$ to calculate polarization charge density gives rise to a dielectric constant $K \equiv \epsilon/\epsilon_0 = 1 + 36\pi \times 10^9 \rho/B^2$ (SI) $= 1 + 4\pi \rho c^2/B^2$ (Gaussian), where ρ is the mass density.

The electromagnetic energy in volume V is given by

$$W = \frac{1}{2} \int_V dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D}) \quad (\text{SI})$$

$$= \frac{1}{8\pi} \int_V dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D}) \quad (\text{Gaussian}).$$

Poynting's theorem is

$$\frac{\partial W}{\partial t} + \int \mathbf{N} \cdot d\mathbf{S} = - \int_V dV \mathbf{J} \cdot \mathbf{E},$$

where S is the closed surface bounding V and the Poynting vector (energy flux across S) is given by $\mathbf{N} = \mathbf{E} \times \mathbf{H}$ (SI) or $\mathbf{N} = c\mathbf{E} \times \mathbf{H}/4\pi$ (Gaussian).

```

\input prolog.tex
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein
\pageno=20
\centerline{\headfont ELECTRICITY AND MAGNETISM}
\medskip\indent
In the following,  $\epsilon$ =dielectric permittivity,  $\mu$ =permeability
of conductor,  $\mu'$ =permeability of surrounding medium,
 $\sigma$ =conductivity,  $f=\omega/2\pi$ =radiation frequency,
 $\kappa_m=\mu/\mu_0$  and  $\kappa_e=\epsilon/\epsilon_0$ . Where subscripts are
used, '1' denotes a conducting medium and '2' a propagating (lossless
dielectric) medium. All units are SI unless otherwise specified.

\msk \halign{\hfil\quad\&=\ \hfil\cr
Permittivity of free space &  $\epsilon_0$  &  $8.8542$   

 $\times 10^{-12}$  \, F/m \cr \sk
Permeability of free space &  $\mu_0$  &  $4\pi \times 10^{-7}$  \, H/m \cr \sk
 $1.2566 \times 10^{-6}$  \, H/m \cr \sk
Resistance of free space &  $R_0$  &  $(\mu_0/\epsilon_0)^{1/2}=376.73$  \,  

 $\Omega$  \cr \sk
Capacity of parallel plates of area &  $C$  &  $\epsilon A/a$  \cr
\halign{\quad\&, separated by distance  $d$ } \sk
Capacity of concentric cylinders &  $C$  &  $2\pi\epsilon \ln(b/a)$  \cr
\halign{\quad\& of length  $l$ , radii  $a, b$ } \sk
Capacity of concentric spheres of &  $C$  &  $4\pi\epsilon ab/(b-a)$  \cr
\halign{\quad\& radii  $a, b$ } \sk
Self-inductance of wire of length &  $L$  &  $\mu l$  \cr
\halign{\quad\&  $l$ , carrying uniform current} \sk
Mutual inductance of parallel wires &  $L$  &  $(\mu' l/4\pi)$   

 $\ln(d/a)$  \cr
\halign{\quad\& of length  $l$ , radius  $a$ , separated} \sk
\halign{\quad\& by distance  $d$ } \sk
Inductance of circular loop of radius &  $L$  &  $b \ln(8b/a)$   

 $-2$  \cr
 $\ln(0.65ex)$   

\halign{\quad\&  $a$ , made of wire of radius  $a$ ,} \sk
\halign{\quad\& carrying uniform current} \sk
Relaxation time in a lossy medium &  $\tau$  &  $\epsilon/\sigma$  \cr \sk
Skin depth in a lossy medium &  $\delta$  &  $(2/\omega\mu\sigma)^{1/2}$   

 $= (\pi f \mu \sigma)^{-1/2}$  \cr \sk
Wave impedance in a lossy medium &  $Z$  &  $(\mu/(\epsilon+i\sigma/\omega))^{1/2}$   

 $\ln(0.65ex)$  \cr \sk
Transmission coefficient at &  $T$  &  $4.22 \times 10^{-4} (f \kappa_m)$   

 $\kappa_e/\sigma)^{1/2}$  \cr
\halign{\quad\& conducting surface} \sk
\halign{\quad\& good only for  $T \ll 1$ } \sk
Field at distance  $r$  from straight wire &  $E_\theta$  &  $(2I/r)$  \cr
 $(2\pi \times 10^{-7} I/r)$  \cr
\halign{\quad\& carrying current  $I$  (amperes) } \cr
 $(0.2I/r)$  \cr
Field at distance  $z$  along axis from &  $B_z$  &  $(\mu_0 I^2)/(2\pi^2 z^3)$  \cr
\halign{\quad\& circular loop of radius  $a$ } \sk
\halign{\quad\& carrying current  $I$ } \sk
\fil\hfill\end

```

ELECTRICITY AND MAGNETISM

In the following, ϵ = dielectric permittivity, μ = permeability of conductor, μ' = permeability of surrounding medium, σ = conductivity, $f = \omega/2\pi$ = radiation frequency, $\kappa_m = \mu/\mu_0$ and $\kappa_e = \epsilon/\epsilon_0$. Where subscripts are used, '1' denotes a conducting medium and '2' a propagating (lossless dielectric) medium. All units are SI unless otherwise specified.

Permittivity of free space	$\epsilon_0 = 8.8542 \times 10^{-12} \text{ F m}^{-1}$
Permeability of free space	$\mu_0 = 4\pi \times 10^{-7} \text{ H m}^{-1}$ $= 1.2566 \times 10^{-6} \text{ H m}^{-1}$
Resistance of free space	$R_0 = (\mu_0/\epsilon_0)^{1/2} = 376.73 \Omega$
Capacity of parallel plates of area A , separated by distance d	$C = \epsilon A/d$
Capacity of concentric cylinders of length l , radii a, b	$C = 2\pi\epsilon l \ln(b/a)$
Capacity of concentric spheres of radii a, b	$C = 4\pi\epsilon ab/(b-a)$
Self-inductance of wire of length l , carrying uniform current	$L = \mu l$
Mutual inductance of parallel wires of length l , radius a , separated by distance d	$L = (\mu' l/4\pi) [1 + 4 \ln(d/a)]$
Inductance of circular loop of radius b , made of wire of radius a , carrying uniform current	$L = b \left\{ \mu' [\ln(8b/a) - 2] + \mu/4 \right\}$
Relaxation time in a lossy medium	$\tau = \epsilon/\sigma$
Skin depth in a lossy medium	$\delta = (2/\omega\mu\sigma)^{1/2} = (\pi f\mu\sigma)^{-1/2}$
Wave impedance in a lossy medium	$Z = [\mu/(\epsilon + i\sigma/\omega)]^{1/2}$
Transmission coefficient at conducting surface (good only for $T \ll 1$) ⁹	$T = 4.22 \times 10^{-4} (f\kappa_{m1}\kappa_{e2}/\sigma)^{1/2}$
Field at distance r from straight wire carrying current I (amperes)	$B_\mu = \mu I/2\pi r$ tesla $= 0.2I/r$ gauss (r in cm)
Field at distance z along axis from circular loop of radius a carrying current I	$B_z = \mu a^2 I/[2(a^2 + z^2)^{3/2}]$

ELECTROMAGNETIC FREQUENCY/ WAVELENGTH BANDS¹⁰

Designation	Frequency Range		Wavelength Range	
	Lower	Upper	Lower	Upper
ULF*		10 Hz	3 Mm	
ELF*	10 Hz	3 kHz	100 km	3 Mm
VLF	3 kHz	30 kHz	10 km	100 km
LF	30 kHz	300 kHz	1 km	10 km
MF	300 kHz	3 MHz	100 m	1 km
HF	3 MHz	30 MHz	10 m	100 m
VHF	30 MHz	300 MHz	1 m	10 m
UHF	300 MHz	3 GHz	10 cm	1 m
SHF†	3 GHz	30 GHz	1 cm	10 cm
S	2.6	3.95	7.6	11.5
G	3.95	5.85	5.1	7.6
J	5.3	8.2	3.7	5.7
H	7.05	10.0	3.0	4.25
X	8.2	12.4	2.4	3.7
M	10.0	15.0	2.0	3.0
P	12.4	18.0	1.67	2.4
K	18.0	26.5	1.1	1.67
R	26.5	40.0	0.75	1.1
EHF	30 GHz	300 GHz	1 mm	1 cm
Submillimeter	300 GHz	3 THz	100 μ m	1 mm
Infrared	3 THz	430 THz	700 nm	100 μ m
Visible	430 THz	750 THz	400 nm	700 nm
Ultraviolet	750 THz	30 PHz	10 nm	400 nm
X Ray	30 PHz	3 EHz	100 pm	10 nm
Gamma Ray	3 EHz			100 pm

Note: In spectroscopy the angstrom (\AA) is sometimes used ($1 \text{\AA} = 10^{-8} \text{ cm} = 0.1 \text{ nm}$).

*The boundary between ULF and ELF is variously defined.

†The SHF (microwave) band is further subdivided approximately as shown.¹¹

```

\input prolog
\hoffset=1.25truein
\voffset=1.0truein
\hsize=6.0truein
\vsize=9.0truein
\pageno=22
\centerline{\headfont AC CIRCUITS}
\medskip\indent
For a resistance  $R$ , inductance  $L$ , and capacitance  $C$  in series with a
voltage source  $V=V_0\exp(i\omega t)$  (here  $i=\sqrt{-1}$ ), the current is given
by  $I=dq/dt$ , where  $q$  satisfies

$$L\{d^2q/dt^2\}+R\{dq/dt\}+q/C = V.$$

Solutions are  $q(t)=q_s+q_t$ ,  $I(t)=I_s+I_t$ , where the steady state is
 $I_s = i\omega q_s = V/Z$  in terms of the impedance  $Z = R + i(\omega L - 1/\omega C)$ 
and  $I_t = dq_t/dt$ . For initial conditions  $q(0)\equiv q_0 = \bar{q}_0+q_s$ ,
 $I(0)\equiv I_0$ , the transients can be of three types,
depending on  $\Delta=R^2-4L/C$ :
\medskip\noindent
(a) Overdamped,  $\Delta>0$ 

$$q_t = \frac{I_0+\gamma_+ \bar{q}_0}{\gamma_+-\gamma_-} \exp(-\gamma_+ t) - \frac{I_0+\gamma_- \bar{q}_0}{\gamma_+-\gamma_-} \exp(-\gamma_- t),$$


$$I_t = \frac{\gamma_+(I_0+\gamma_- \bar{q}_0)}{\gamma_+-\gamma_-} \exp(-\gamma_+ t) - \frac{\gamma_-(I_0+\gamma_+ \bar{q}_0)}{\gamma_+-\gamma_-} \exp(-\gamma_- t),$$

where  $\gamma_{\pm}=(R\pm\Delta^{1/2})/2L$ ;
\medskip\noindent
(b) Critically damped,  $\Delta=0$ 

$$q_t = \left[ \bar{q}_0 + (I_0+\gamma_R \bar{q}_0)t \right] \exp(-\gamma_R t),$$


$$I_t = \left[ I_0 - (I_0+\gamma_R \bar{q}_0)\gamma_R t \right] \exp(-\gamma_R t),$$

where  $\gamma_R=R/2L$ ;
\medskip\noindent
(c) Underdamped,  $\Delta<0$ 

$$q_t = \left[ \frac{\gamma_R \bar{q}_0 + I_0}{\omega_1} \sin\omega_1 t + \bar{q}_0 \cos\omega_1 t \right] \exp(-\gamma_R t),$$


$$I_t = \left[ I_0 \cos\omega_1 t - \left( \omega_1^2 + \gamma_R^2 \right) \bar{q}_0 + \gamma_R I_0 \right] \frac{\sin\omega_1 t}{\omega_1} \exp(-\gamma_R t),$$

where  $\omega_1=\omega_0(1-R^2C/4L)^{1/2}$  and  $\omega_0=(LC)^{-1/2}$  is
the resonant frequency. At  $\omega=\omega_0$ ,  $Z=R$ . The quality of the
circuit is  $Q=\omega_0 L/R$ . In stability results when  $L$ ,  $R$ ,  $C$  are
not all of the same sign.
\vskip0.5truein
\vfilleject\end

```

AC CIRCUITS

For a resistance R , inductance L , and capacitance C in series with a voltage source $V = V_0 \exp(i\omega t)$ (here $i = \sqrt{-1}$), the current is given by $I = dq/dt$, where q satisfies

$$L \frac{d^2 q}{dt^2} + R \frac{dq}{dt} + \frac{q}{C} = V.$$

Solutions are $q(t) = q_s + q_t$, $I(t) = I_s + I_t$, where the steady state is $I_s = i\omega q_s = V/Z$ in terms of the impedance $Z = R + i(\omega L - 1/\omega C)$ and $I_t = dq_t/dt$. For initial conditions $q(0) \equiv q_0 = \bar{q}_0 + q_s$, $I(0) \equiv I_0$, the transients can be of three types, depending on $\Delta = R^2 - 4L/C$:

(a) Overdamped, $\Delta > 0$

$$\begin{aligned} q_t &= \frac{I_0 + \gamma_+ \bar{q}_0}{\gamma_+ - \gamma_-} \exp(-\gamma_- t) - \frac{I_0 + \gamma_- \bar{q}_0}{\gamma_+ - \gamma_-} \exp(-\gamma_+ t), \\ I_t &= \frac{\gamma_+ (I_0 + \gamma_- \bar{q}_0)}{\gamma_+ - \gamma_-} \exp(-\gamma_+ t) - \frac{\gamma_- (I_0 + \gamma_+ \bar{q}_0)}{\gamma_+ - \gamma_-} \exp(-\gamma_- t), \end{aligned}$$

where $\gamma_{\pm} = (R \pm \Delta^{1/2})/2L$;

(b) Critically damped, $\Delta = 0$

$$\begin{aligned} q_t &= [\bar{q}_0 + (I_0 + \gamma_R \bar{q}_0)t] \exp(-\gamma_R t), \\ I_t &= [I_0 - (I_0 + \gamma_R \bar{q}_0)\gamma_R t] \exp(-\gamma_R t), \end{aligned}$$

where $\gamma_R = R/2L$;

(c) Underdamped, $\Delta < 0$

$$\begin{aligned} q_t &= \left[\frac{\gamma_R q_0 + I_0}{\omega_1} \sin \omega_1 t + \bar{q}_0 \cos \omega_1 t \right] \exp(-\gamma_R t), \\ I_t &= \left[I_0 \cos \omega_1 t - \frac{(\omega_1^2 + \gamma_R^2) \bar{q}_0 + \gamma_R I_0}{\omega_1} \sin(\omega_1 t) \right] \exp(-\gamma_R t), \end{aligned}$$

where $\omega_1 = \omega_0(1 - R^2 C/4L)^{1/2}$ and $\omega_0 = (LC)^{-1/2}$ is the resonant frequency. At $\omega = \omega_0$, $Z = R$. The quality of the circuit is $Q = \omega_0 L/R$. Instability results when L , R , C are not all of the same sign.

```

\input prolog
\hoffset=1truein\voffset=1truein\hsize=6.5truein\vsize=9truein
\pageno=23
\centerline{\headfont DIMENSIONLESS NUMBERS OF FLUID MECHANICS$^{-12}$}
\bsk % BEGINNING OF TABLE.
\vbox{\tabskip=0pt \offinterlineskip \def\quid{\hskip0.5em\relax}
\halign to \hsize{\vrule#\tabskip=0.5em plus2em&\strut#\hfil&\vrule##&
\hfil&\vrule##&\hfil&\vrule##&\hfil&\vrule#\tabskip=0pt\cr \trule
&\hfil Name(s)\om\hidewidth Symbol\hidewidth\hfil Definition\hfil
Significance&\cr \tskc{4}{2pt} \trule \tskc{4}{1.0pt} \trule \tskc{4}{2pt}
&Alfvén, |A|, Ka|V_A/V|{\kern-0.5em*(Magnetic force/&\cr
&\om\quid K'arm'an\hfil\om\om\quad inertial force)$^{-1/2}$&\cr \tskc{4}{2pt}
&Bond|Bd|$(\rho'\text{-}\rho)L^2g/\Sigma$|Gravitational force/&\cr
&\om\om\om\quad surface tension&\cr \tskc{4}{2pt}
&Boussinesq|B|$(V/(2gR))^{1/2}$|Inertial force/&\cr
&\om\om\om\quad gravitational force)$^{-1/2}$&\cr \tskc{4}{2pt}
&Brinkman|Br|$\mu V^2/k\Delta T$|Viscous heat/conducted heat&\cr \tskc{4}{2pt}
&Capillary|Cp|$\mu V/\Sigma$|Viscous force/surface tension&\cr \tskc{4}{2pt}
&Carnot|Ca|$(T_2-T_1)/T_2$|Theoretical Carnot cycle&\cr
&\om\om\om\quad efficiency&\cr \tskc{4}{2pt}
&Cauchy, |Cy, Hk|$\rho V^2/\Gamma=M^2$|Inertial force/&\cr
&\om\quid Hooke\hfil\om\om\quad compressibility force&\cr \tskc{4}{2pt}
&Clausius|Cl|$(LV^3\rho/k\Delta T)$|Kinetic energy flow rate/heat&\cr
&\om\om\om\quad conduction rate&\cr \tskc{4}{2pt}
&Cowling|C|$(V_A/V)^2=M^2$|Magnetic force/inertial force&\cr \tskc{4}{2pt}
&Crispation|Cr|$\mu \kappa/\Sigma L$|Effect of diffusion/effect of&\cr
&\om\om\om\quad surface tension&\cr \tskc{4}{2pt}
&Dean|D|$(D^{3/2}V/\nu(2r))^{1/2}$|Transverse flow due to&\cr
&\om\om\om\quad curvature/longitudinal flow&\cr \tskc{4}{2pt}
&[Drag|C_D|$(\rho'\text{-}\rho)Lg/$|Drag force/inertial force&\cr
&\om\quid coefficient]\hfil\om\quad $\rho'\text{-}\rho V^2$|&\cr \tskc{4}{2pt}
&Eckert|E|$(V^2/c_p\Delta T)$|Kinetic energy/change in&\cr
&\om\om\om\quad thermal energy&\cr \tskc{4}{2pt}
&Ekman|Ek|$(\nu/2\Omega L^2)^{1/2}=$|Viscous force/Coriolis force)$^{-1/2}$
&\hidewidth&\cr
&\om\om\quad $(Ro/Re)^{1/2}$|&\cr \tskc{4}{2pt}
&Euler|Eu|$(\Delta p/\rho V^2)$|Pressure drop due to friction/&\cr
&\om\om\om\quad dynamic pressure&\cr \tskc{4}{2pt}
&Froude|Fr|$(V/(gL))^{1/2}$|{\kern-0.5em\dag}Inertial force/gravitational or
&\kern0.25em&\cr
&\om\om$V/NL$|&\quad buoyancy force)$^{-1/2}$&\cr \tskc{4}{2pt}
&Gay-Lussac|Ga|$1/\beta\Delta T$|Inverse of relative change in&\cr
&\om\om\om\quad volume during heating&\cr \tskc{4}{2pt}
&Grashof|Gr|$(gL^3\beta\Delta T/\nu^2)$|Buoyancy force/viscous force&\cr \tskc{4}{2pt}
&[Hall|C_H|$ \lambda/r_L$|Gyrofrequency/&\cr
&\om\quid coefficient]\hfil\om\om\quad collision frequency&\cr \tskc{4}{2pt}
&\tskc{4}{2pt} \trule}}
•(dag) Also defined as the inverse (square) of the quantity shown.
\fil -ject&end

```

DIMENSIONLESS NUMBERS OF FLUID MECHANICS¹²

Name(s)	Symbol	Definition	Significance
Alfvén, Kármán	Al, Ka	V_A/V	*(Magnetic force/ inertial force) ^{1/2}
Bond	Bd	$(\rho' - \rho)L^2g/\Sigma$	Gravitational force/ surface tension
Boussinesq	B	$V/(2gR)^{1/2}$	(Inertial force/ gravitational force) ^{1/2}
Brinkman	Br	$\mu V^2/k\Delta T$	Viscous heat/conducted heat
Capillary	Cp	$\mu V/\Sigma$	Viscous force/surface tension
Carnot	Ca	$(T_2 - T_1)/T_2$	Theoretical Carnot cycle efficiency
Cauchy, Hooke	Cy, Hk	$\rho V^2/\Gamma = M^2$	Inertial force/ compressibility force
Clausius	Cl	$LV^3\rho/k\Delta T$	Kinetic energy flow rate/heat conduction rate
Cowling	C	$(V_A/V)^2 = Al^2$	Magnetic force/inertial force
Crispation	Cr	$\mu\kappa/\Sigma L$	Effect of diffusion/effect of surface tension
Dean	D	$D^{3/2}V/\nu(2r)^{1/2}$	Transverse flow due to curvature/longitudinal flow
[Drag coefficient]	C_D	$(\rho' - \rho)Lg/\rho'V^2$	Drag force/inertial force
Eckert	E	$V^2/c_p\Delta T$	Kinetic energy/change in thermal energy
Ekman	Ek	$(\nu/2\Omega L^2)^{1/2} =(Ro/Re)^{1/2}$	(Viscous force/Coriolis force) ^{1/2}
Euler	Eu	$\Delta p/\rho V^2$	Pressure drop due to friction/ dynamic pressure
Froude	Fr	$V/(gL)^{1/2}$ V/NL	†(Inertial force/gravitational or buoyancy force) ^{1/2}
Gay-Lussac	Ga	$1/\beta\Delta T$	Inverse of relative change in volume during heating
Grashof	Gr	$gL^3\beta\Delta T/\nu^2$	Buoyancy force/viscous force
[Hall coefficient]	C_H	λ/r_L	Gyrofrequency/ collision frequency

*(†) Also defined as the inverse (square) of the quantity shown.

```

\input prolog \pageno=24 % BEGINNING OF TABLE.
\hoffset=1truein\voffset=1truein\hsize=6.5truein\vsizer=9truein
\ vbox{\tabskip=0pt \offinterlineskip \def\quid{\hskip0.5em\relax}
\halign to \hsize{\vrule# \tabskip=0.5em plus2em\strut#\hfil\vrule##
\hfil\vrule##\hfil\vrule##\hfil\vrule#\tabskip=0pt\cr \trule
&\hfil Name(s)\om\hidewidth Symbol\hidewidth\hfil Definition\hfil
Significance\cr
% DRAW DOUBLE RULE BENEATH HEADING.
\tskc{4}{2pt} \trule \tskc{4}{1.0pt} \trule \tskc{4}{2pt}
&Hartmann|H| $\$BL/(\mu\eta)$ | $\$1/2=\$|$ Magnetic force/ $\$ \cr
&\om\om\quad (Rm\ts Re\ts C)\$^{-1/2}\$ \hidewidth\quad$ dissipative
force $\$ \cr \tskc{4}{2pt}$ 
&Knudsen|Kn| $\$ \lambda/L$ |Hydrodynamic time/ $\$ \cr
&\om\om\om\quad$ collision time $\$ \cr \tskc{4}{2pt}$ 
&Lorentz|Lo| $\$V/c$ |Magnitude of relativistic effects $\$ \cr \tskc{4}{2pt}$ 
&Lundquist|Lu| $\mu_0LV/A\eta$ | $\$ \cr
&\{\bf J\}\times\{\bf B\}$ |force/resistive magnetic $\$ \cr
&\om\om\quad A\ts Rm\quad$ diffusion force  $\$ \cr \tskc{4}{2pt}$ 
&Mach|M| $\$V/C_s$ |Magnitude of compressibility $\$ \cr
&\om\om\om\quad$ effects $\$ \cr \tskc{4}{2pt}$ 
&Magnetic|Mm| $\$V/V_A=\$ A\ts^{-1}\$ |($ Inertial force/magnetic force) $\$^{-1/2}\$ \cr
\hidewidth\quad$ 
&\om\quid Mach\hfil\om\om\om $\$ \cr \tskc{4}{2pt}$ 
&Magnetic|Rm| $\mu_0LV/\eta$ |Flow velocity/magnetic diffusion $\$ \cr
&\om\quid$ Reynolds\hfil\om\om\quad velocity $\$ \cr \tskc{4}{2pt}$ 
&Newton|Nt| $\$F/\rho L^2V^2$ |Imposed force/inertial force $\$ \cr \tskc{4}{2pt}$ 
&Nusselt|Nt| $\alpha L/k$ |Total heat transfer/thermal $\$ \cr
&\om\om\om\quad$ conduction $\$ \cr \tskc{4}{2pt}$ 
&P'eclet|Pe| $\$LV/\kappa$ |Heat convection/heat conduction\hidewidth $\$ \cr \tskc{4}{2pt}$ 
&Poisueille|Po| $\$D^2\Delta p/\mu LV$ |Pressure force/viscous force $\$ \cr \tskc{4}{2pt}$ 
&Prandtl|Pr| $\$ \nu/\kappa$ |Momentum diffusion/ $\$ \cr
&\om\quid$ Schmidt\hfil\om\om\quad heat diffusion $\$ \cr \tskc{4}{2pt}$ 
&Rayleigh|Ra| $\$gH^3\beta\Delta T/\nu\kappa$ |Buoyancy force/diffusion force $\$ \cr
\tskc{4}{2pt}$ 
&Reynolds|Re| $\$LV/\mu$ |Inertial force/viscous force $\$ \cr \tskc{4}{2pt}$ 
&Richardson|Ri| $\$(NH/\Delta V)^2$ |Buoyancy effects/ $\$ \cr
&\om\om\om\quad$ vertical shear effects $\$ \cr \tskc{4}{2pt}$ 
&Rossby|Ro| $\$V/2\Omega a$ |Inertial force/Coriolis force $\$ \cr \tskc{4}{2pt}$ 
&Stanton|St| $\alpha/\rho c_p V$ |Thermal conduction loss/ $\$ \cr
&\om\om\om\quad$ heat capacity $\$ \cr \tskc{4}{2pt}$ 
&Stefan|Sf| $\sigma L^2/k$ |Radiated heat/conducted heat $\$ \cr \tskc{4}{2pt}$ 
&Stokes|S| $\nu/L^2f$ |Viscous damping rate/ $\$ \cr
&\om\om\om\quad$ vibration frequency $\$ \cr \tskc{4}{2pt}$ 
&Strouhal|Stf| $L/V$ |Vibration speed/flow velocity $\$ \cr \tskc{4}{2pt}$ 
&Taylor|Ta| $\$(\rho\omega a^2L)/\mu$ |Centrifugal force/viscous force $\$ \cr
&\rho\omega^2L^3/(12\mu)\quad$ Centrifugal force/ $\$ \cr
&\om\om\quad$ viscous force $\$^{1/2}\$ \cr \tskc{4}{2pt}$ 
&Thring|Th| $\$T/\rho c_p$ |Convective heat transport/ $\$ \cr
&\om\quad$ radiative heat transport $\$ \cr \tskc{4}{2pt}$ 
& Weber|We| $\rho V^2L/\sigma$ |Inertial force/surface tension $\$ \cr
\tskc{4}{2pt}$ 
\end{pre>

```

Name(s)	Symbol	Definition	Significance
Hartmann	H	$BL/(\mu\eta)^{1/2} = (\text{Rm Re C})^{1/2}$	Magnetic force/ dissipative force
Knudsen	Kn	λ/L	Hydrodynamic time/ collision time
Lorentz	Lo	V/c	Magnitude of relativistic effects
Lundquist	Lu	$\mu_0 LV_A/\eta = \text{Al Rm}$	$\mathbf{J} \times \mathbf{B}$ force/resistive magnetic diffusion force
Mach	M	V/C_S	Magnitude of compressibility effects
Magnetic Mach	Mm	$V/V_A = \text{Al}^{-1}$	(Inertial force/magnetic force) ^{1/2}
Magnetic Reynolds	Rm	$\mu_0 LV/\eta$	Flow velocity/magnetic diffusion velocity
Newton	Nt	$F/\rho L^2 V^2$	Imposed force/inertial force
Nusselt	N	$\alpha L/k$	Total heat transfer/thermal conduction
Péclet	Pe	LV/κ	Heat convection/heat conduction
Poiseuille	Po	$D^2 \Delta p/\mu LV$	Pressure force/viscous force
Prandtl, Schmidt	Pr, Sc	ν/κ	Momentum diffusion/ heat diffusion
Rayleigh	Ra	$gH^3 \beta \Delta T/\nu \kappa$	Buoyancy force/diffusion force
Reynolds	Re	LV/ν	Inertial force/viscous force
Richardson	Ri	$(NH/\Delta V)^2$	Buoyancy effects/ vertical shear effects
Rossby	Ro	$V/2\Omega L \sin \Lambda$	Inertial force/Coriolis force
Stanton	St	$\alpha/\rho c_p V$	Thermal conduction loss/ heat capacity
Stefan	Sf	$\sigma LT^3/k$	Radiated heat/conducted heat
Stokes	S	$\nu/L^2 f$	Viscous damping rate/ vibration frequency
Strouhal	Sr	fL/V	Vibration speed/flow velocity
Taylor	Ta	$(2\Omega L^2/\nu)^2$ $R^{1/2}(\Delta R)^{3/2}$ $\cdot (\Omega/\nu)$	Centrifugal force/viscous force (Centrifugal force/ viscous force) ^{1/2}
Thring, Boltzmann	Th, Bo	$\rho c_p V/\epsilon \sigma T^3$	Convective heat transport/ radiative heat transport
Weber	W	$\rho LV^2/\Sigma$	Inertial force/surface tension

```

\input prolog
\hoffset=1truein\voffset=1truein\hsize=6.5truein\vsize=9truein
\pageno=25
{\headfont Nomenclature:}
\msk \def\Cr{\cr \noalign{\vskip 2.5pt}}
% \Cr HAS BEEN DEFINED TO LEAVE AN EXTRA SPACE AFTER EACH ALIGNED LINE.
\halign{#\hfil&\qqquad#\hfil\cr
B$&Magnetic induction\Cr
C_s,c$&Speeds of sound, light\Cr
c_p$&Specific heat at constant pressure (units $\rm m^2\ts s^{-2}\ts
K^{-1}$)\Cr
D=2R$&Pipe diameter\Cr
F$&Imposed force\Cr
f$&Vibration frequency\Cr
g$&Gravitational acceleration\Cr
H, L$&Vertical, horizontal length scales\Cr
k=\rho c_p \kappa$&Thermal conductivity (units $\rm kg\ts m^{-1}\ts s^{-2}$)\Cr
N=(g/H)^{1/2}$&Brunt--V\ais\al"a frequency\Cr
R$&Radius of pipe or channel\Cr
r$&Radius of curvature of pipe or channel\Cr
r_L$&Larmor radius\Cr
T$&Temperature\Cr
V$&Characteristic flow velocity\Cr
V_A=B/(\mu_0\rho)^{1/2}$&Alfv\en speed\cr
\alpha$&Newton's-law heat coefficient, $\displaystyle k\{\part T \over
\part x\}=\alpha\Delta T$\Cr
\beta$&Volumetric expansion coefficient, $dV/V = \beta dT$\Cr
\Gamma$&Bulk modulus (units $\rm kg\ts m^{-1}\ts s^{-2}$)\Cr
\Delta R, \Delta V, \Delta p, \Delta T$&Imposed difference in two radii,
velocities,\cr
&pressures, or temperatures\Cr
\epsilon$&Surface emissivity\Cr
\eta$&Electrical resistivity\Cr
\kappa$&Thermal diffusivity (units $\rm m^2\ts s^{-1}$)\Cr
\Lambda$&Latitude of point on earth's surface\Cr
\lambda$&Collisional mean free path\Cr
\mu=\rho\nu$&Bulk viscosity\Cr
\mu_0$&Permeability of free space\Cr
\nu$&Kinematic viscosity (units $\rm m^2\ts s^{-1}$)\Cr
\rho$&Mass density of fluid medium\Cr
\rho_o$&Mass density of bubble, droplet, or moving object\Cr
\sigma$&Surface tension (units $\rm kg\ts s^{-2}$)\Cr
\sigma_s$&Stefan-Boltzmann constant\Cr
\Omega$&Angular velocity\Cr
\omega$&Angular velocity

```


Nomenclature:

B	Magnetic induction
C_s, c	Speeds of sound, light
c_p	Specific heat at constant pressure (units $\text{m}^2 \text{s}^{-2} \text{K}^{-1}$)
$D = 2R$	Pipe diameter
F	Imposed force
f	Vibration frequency
g	Gravitational acceleration
H, L	Vertical, horizontal length scales
$k = \rho c_p \kappa$	Thermal conductivity (units $\text{kg m}^{-1} \text{s}^{-2}$)
$N = (g/H)^{1/2}$	Brunt-Väisälä frequency
R	Radius of pipe or channel
r	Radius of curvature of pipe or channel
r_L	Larmor radius
T	Temperature
V	Characteristic flow velocity
$V_A = B/(\mu_0 \rho)^{1/2}$	Alfvén speed
α	Newton's-law heat coefficient, $k \frac{\partial T}{\partial x} = \alpha \Delta T$
β	Volumetric expansion coefficient, $dV/V = \beta dT$
Γ	Bulk modulus (units $\text{kg m}^{-1} \text{s}^{-2}$)
$\Delta R, \Delta V, \Delta p, \Delta T$	Imposed difference in two radii, velocities, pressures, or temperatures
ϵ	Surface emissivity
η	Electrical resistivity
κ	Thermal diffusivity (units $\text{m}^2 \text{s}^{-1}$)
Λ	Latitude of point on earth's surface
λ	Collisional mean free path
$\mu = \rho \nu$	Bulk viscosity
μ_0	Permeability of free space
ν	Kinematic viscosity (units $\text{m}^2 \text{s}^{-1}$)
ρ	Mass density of fluid medium
ρ'	Mass density of bubble, droplet, or moving object
Σ	Surface tension (units kg s^{-2})
σ	Stefan Boltzmann constant
Ω	Solid-body rotational angular velocity

```

\input prolog
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein
\pageno=26
{\def\ups{\upsilon} \def\NQ{\medskip\N\quad} \def\PH{${\phantom{1}}$}
\def\sq{\phantom{1}}2}} \def\medskip{\vskip4.0pt}
% \PH SKIPS THE SPACE OF THE DIGIT '1' SO THAT EQUATIONS NUMBERS LINE UP.
% \medskip HAS BEEN REDEFINED TO ADJUST THE SPACING OF THE TABLE.
% \NQ SKIPS SPACE DOWN AND INDENTS ONE \quad TO ALIGN EQUATIONS.
\centerline{\headfont SHOCKS} \bsk\indent
At a shock front propagating in a magnetized fluid at an angle  $\theta$  with
respect to the magnetic induction  $\{B\}$ , the jump conditions are  $\{13, 14\}$ 
\NQ\PH(1)  $\rho U = \rho U \equiv q$ ;
\NQ\PH(2)  $\rho U^2 + p + B_{\perp}^2 / 2\mu = \rho U^2 +$ 
 $\rho p + \rho B_{\perp}^2 / 2\mu$ ;
\NQ\PH(3)  $\rho UV - B_{\parallel} B_{\perp} / \mu =$ 
 $\rho U V - B_{\parallel} B_{\perp} / \mu$ ;
\NQ\PH(4)  $B_{\parallel} = B_{\parallel}$ ;
\NQ\PH(5)  $UB_{\perp} - VB_{\parallel} = U B_{\perp} -$ 
 $V B_{\parallel}$ ;
\NQ\PH(6)  $\{1/2\}(U^2 + V^2) + w + (UB_{\perp}^2 - VB_{\parallel}^2) / \mu \rho U$ 
\smallskip
 $\{1/2\}(\rho U^2 + \rho V^2) + \rho w + (\rho U B_{\perp}^2 -$ 
 $\rho V B_{\parallel}^2) / \mu \rho U$ .
\medskip\N
Here  $U$  and  $V$  are components of the fluid velocity normal and tangential to
the front in the shock frame;  $\rho = 1/\ups$  is the mass density;  $p$  is the
pressure;  $B_{\perp} = B \sin \theta$ ,  $B_{\parallel} = B \cos \theta$ ;  $\mu$  is the
magnetic permeability ( $\mu = 4\pi$  in cgs units); and the specific enthalpy is
 $w = e + p\ups$ , where the specific internal energy  $e$  satisfies  $de = Tds$ 
-  $p d\ups$  in terms of the temperature  $T$  and the specific entropy  $s$ .
Quantities in the region behind (downstream from) the front are distinguished by
a bar. If  $\{B\} = 0$ , then  $\{15\}$ 
\NQ\PH(7)  $U - \rho U = \left[ (\rho p - p)(\ups - \rho \ups) \right]^{1/2}$ ;
\NQ\PH(8)  $(\rho p - p)(\ups - \rho \ups)^{-1} = q^2$ ;
\NQ\PH(9)  $\rho w - w = \{1/2\}(\rho p - p)(\ups + \rho \ups)$ ;
\NQ(10)  $\rho e - e = \{1/2\}(\rho p + p)(\ups - \rho \ups)$ .
\medskip\N
In what follows we assume that the fluid is a perfect gas with adiabatic index
 $\gamma = 1 + 2/n$ , where  $n$  is the number of degrees of freedom. Then  $p =$ 
 $\rho RT/m$ , where  $R$  is the universal gas constant and  $m$  is the molar weight;
the sound speed is given by  $\{C_s\}^2 = (\partial p / \partial \rho)_s = \gamma$ 
 $p\ups$ ; and  $w = \gamma e = \gamma p \ups / (\gamma + 1)$ . For a general oblique
shock in a perfect gas the quantity  $X = r^{-1}(U/V_A)^2$  satisfies  $\{14\}$ 
\NQ(11)  $(X - \beta/\alpha)(X - \cos^2 \theta)^2 =$ 
 $X \sin^2 \theta \left[ 1 + (r-1)/2\alpha \right] X - \cos^2 \theta$ ,
where  $r = \rho / \rho$ ,  $\alpha = \{1/2\} \left[ \gamma + 1 - (\gamma - 1)r \right]$ ,
and  $\beta = \{C_s\}^2 / \{V_A\}^2 = 4\pi \gamma p / B^2$ .
\medskip\N
The density ratio is bounded by
\NQ(12)  $1 < r < (\gamma + 1) / (\gamma - 1)$ . \medskip\N
If the shock is normal to  $\{B\}$  (i.e., if  $\theta = \pi/2$ ), then
\NQ(13)  $U^2 = (r/\alpha) \left[ \{C_s\}^2 + \{V_A\}^2 \left[ 1 + (1 - \gamma/2)(r-1) \right] r \right]$ ;
\vfil\eject\end

```

SHOCKS

At a shock front propagating in a magnetized fluid at an angle θ with respect to the magnetic induction \mathbf{B} , the jump conditions are^{13,14}

$$\begin{aligned}
 (1) \quad & \rho U = \bar{\rho} \bar{U} \equiv q; \\
 (2) \quad & \rho U^2 + p + B_{\perp}^2/2\mu = \bar{\rho} \bar{U}^2 + \bar{p} + \bar{B}_{\perp}^2/2\mu; \\
 (3) \quad & \rho UV - B_{\parallel} B_{\perp}/\mu = \bar{\rho} \bar{U} \bar{V} - \bar{B}_{\parallel} \bar{B}_{\perp}/\mu; \\
 (4) \quad & B_{\parallel} = \bar{B}_{\parallel}; \\
 (5) \quad & UB_{\perp} - VB_{\parallel} = \bar{U} \bar{B}_{\perp} - \bar{V} \bar{B}_{\parallel}; \\
 (6) \quad & \frac{1}{2}(U^2 + V^2) + w + (UB_{\perp}^2 - VB_{\parallel} B_{\perp})/\mu\rho U \\
 & \quad = \frac{1}{2}(\bar{U}^2 + \bar{V}^2) + \bar{w} + (\bar{U} \bar{B}_{\perp}^2 - \bar{V} \bar{B}_{\parallel} \bar{B}_{\perp})/\mu\bar{\rho} \bar{U}.
 \end{aligned}$$

Here U and V are components of the fluid velocity normal and tangential to the front in the shock frame; $\rho = 1/v$ is the mass density; p is the pressure; $B_{\perp} = B \sin \theta$, $B_{\parallel} = B \cos \theta$; μ is the magnetic permeability ($\mu = 4\pi$ in cgs units); and the specific enthalpy is $w = e + pv$, where the specific internal energy e satisfies $de = Tds - pdv$ in terms of the temperature T and the specific entropy s . Quantities in the region behind (downstream from) the front are distinguished by a bar. If $\mathbf{B} = 0$, then¹⁵

$$\begin{aligned}
 (7) \quad & U - \bar{U} = [(\bar{p} - p)(v - \bar{v})]^{1/2}; \\
 (8) \quad & (\bar{p} - p)(v - \bar{v})^{-1} = q^2; \\
 (9) \quad & \bar{w} - w = \frac{1}{2}(\bar{p} - p)(v + \bar{v}); \\
 (10) \quad & \bar{v} - v = \frac{1}{2}(\bar{p} + p)(v - \bar{v}).
 \end{aligned}$$

In what follows we assume that the fluid is a perfect gas with adiabatic index $\gamma = 1 + 2/n$, where n is the number of degrees of freedom. Then $p = \rho RT/m$, where R is the universal gas constant and m is the molar weight; the sound speed is given by $C_s^2 = (\partial p / \partial \rho)_s = \gamma p v$; and $w = \gamma e = \gamma p v / (\gamma + 1)$. For a general oblique shock in a perfect gas the quantity $X = v^{-1}(U/V_A)^2$ satisfies¹⁴

$$(11) \quad (X - \beta/\alpha)(X - \cos^2 \theta)^2 = X \sin^2 \theta \left\{ [1 + (r - 1)/2\alpha] X - \cos^2 \theta \right\},$$

where $r = \rho/\bar{\rho}$, $\alpha = \frac{1}{2}[\gamma + 1 - (\gamma - 1)r]$, and $\beta = C_s^2/V_A^2 = 4\pi\gamma p/B^2$.

The density ratio is bounded by

$$(12) \quad 1 < r < (\gamma + 1)/(\gamma - 1).$$

If the shock is normal to \mathbf{B} (i.e., if $\theta = \pi/2$), then

$$(13) \quad U^2 = (r/\alpha) \left\{ C_s^2 + V_A^2 [1 + (1 - \gamma/2)(r - 1)] \right\};$$

```

\input prolog
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein
\pageno=27
{\def\ups{\upsilon}\def\NQ{\medskip\N\quad}
\def\sqr{\sqrt}\def\medskip{\vskip4.0pt}
% \PH SKIPS THE SPACE OF THE DIGIT '1', SO THAT EQUATION NUMBERS LINE UP.
% \medskip HAS BEEN REDEFINED, TO ADJUST THE SPACING OF THE TABLE.
% \NQ SKIPS SPACE DOWN AND INDENTS ONE \quad, TO ALIGN EQUATIONS.
\NQ(14) 
$$\frac{U}{\rho V} = \frac{B}{B} = r;$$

\NQ(15) 
$$\rho V = V;$$

\NQ(16) 
$$\rho p = p + (1 - r^{-1})\rho U^2 + (1 - r^2)B^2/2\mu.$$
 \medskip\N
If  $\theta = 0$ , there are two possibilities: switch-on shocks, which require
 $\beta < 1$  and for which
\NQ(17) 
$$U^2 = r(V_A)^2;$$

\NQ(18) 
$$\rho U = (V_A)^2/U;$$

\NQ(19) 
$$\rho B_{\perp} = 2B_{\parallel} \sqrt{r-1}(\alpha - \beta);$$

\NQ(20) 
$$\rho V = \rho U \rho B_{\perp} / B_{\parallel};$$

\NQ(21) 
$$\rho p = p + \rho U^2(1 - \alpha + \beta)(1 - r^{-1}),$$
 \medskip\N
and acoustic (hydrodynamic) shocks, for which
\NQ(22) 
$$U^2 = (r/\alpha)C_s^2;$$

\NQ(23) 
$$\rho U = U/r;$$

\NQ(24) 
$$\rho V = \rho B_{\perp} = 0;$$

\NQ(25) 
$$\rho p = p + \rho U^2(1 - r^{-1}).$$
 \medskip\N
For acoustic shocks the specific volume and pressure are related by
\NQ(26)

$$\frac{\rho}{\rho_0} = \frac{1}{1 + (\gamma - 1)p/\rho_0} \left[ \frac{1}{1 + (\gamma - 1)p/\rho_0} + (\gamma - 1)p/\rho_0 \right];$$

\medskip\N
In terms of the upstream Mach number  $M = U/C_s$ ,
\NQ(27) 
$$\frac{\rho}{\rho_0} = \frac{1}{1 + (\gamma - 1)M^2};$$


$$U = \frac{1}{M} \sqrt{\frac{2(\gamma - 1)M^2}{2\gamma - \gamma - 1}};$$

\NQ(28) 
$$\frac{p}{p_0} = \frac{2\gamma M^2 - \gamma - 1}{2\gamma - \gamma - 1};$$

\NQ(29) 
$$\frac{T}{T_0} = \frac{2(\gamma - 1)M^2 + 2}{2\gamma M^2 - \gamma - 1} \frac{1}{2\gamma - \gamma - 1};$$

\NQ(30) 
$$\frac{M^2}{M_0^2} = \frac{2(\gamma - 1)M^2 + 2}{2\gamma M^2 - \gamma - 1}.$$
 \medskip\N
The entropy change across the shock is
\NQ(31) 
$$\Delta s \equiv s - s_0 = c_v \ln \left[ \frac{\rho}{\rho_0} \left( \frac{p}{p_0} \right)^{1/\gamma} \right],$$

\medskip\N
where  $c_v = R/(\gamma - 1)m$  is the specific heat at constant volume; here  $R$ 
is the gas constant. In the weak-shock limit ( $M \rightarrow 1$ ),
\NQ(32) 
$$\Delta s \rightarrow \frac{c_v}{3} \frac{2\gamma(\gamma - 1)}{\gamma + 1} (M^2 - 1)^3 \approx \frac{16}{3} \gamma R \frac{1}{\gamma + 1} m (M - 1)^3.$$

\medskip\N
The radius at time  $t$  of a strong spherical blast wave resulting from the
explosive release of energy  $E$  in a medium with uniform density  $\rho$  is
\NQ(33) 
$$R_S = C_0 (Et^2/\rho)^{1/5},$$

\medskip\N
where  $C_0$  is a constant depending on  $\gamma$ . For  $\gamma = 7/5$ ,  $C_0 = 1.033$ .
\vfil\egroup\end

```

$$(14) \quad U/\bar{U} = \bar{B}/B = r;$$

$$(15) \quad \bar{V} = V;$$

$$(16) \quad \bar{p} = p + (1 - r^{-1})\rho U^2 + (1 - r^2)B^2/2\mu.$$

If $\theta = 0$, there are two possibilities: switch-on shocks, which require $\beta < 1$ and for which

$$(17) \quad U^2 = rV_A^2;$$

$$(18) \quad \bar{U} = V_A^2/U;$$

$$(19) \quad \bar{B}_\perp^2 = 2B_\parallel^2(r - 1)(\alpha - \beta);$$

$$(20) \quad \bar{V} = \bar{U}\bar{B}_\perp/B_\parallel;$$

$$(21) \quad \bar{p} = p + \rho U^2(1 - \alpha + \beta)(1 - r^{-1}),$$

and acoustic (hydrodynamic) shocks, for which

$$(22) \quad U^2 = (r/\alpha)C_s^2;$$

$$(23) \quad \bar{U} = U/r;$$

$$(24) \quad \bar{V} = \bar{B}_\perp = 0;$$

$$(25) \quad \bar{p} = p + \rho U^2(1 - r^{-1}).$$

For acoustic shocks the specific volume and pressure are related by

$$(26) \quad \bar{v}/v = [(\gamma + 1)p + (\gamma - 1)\bar{p}] / [(\gamma - 1)p + (\gamma + 1)\bar{p}].$$

In terms of the upstream Mach number $M = U/C_s$,

$$(27) \quad \bar{\rho}/\rho = v/\bar{v} = U/\bar{U} = (\gamma + 1)M^2/[(\gamma - 1)M^2 + 2];$$

$$(28) \quad \bar{p}/p = (2\gamma M^2 - \gamma + 1)/(\gamma + 1);$$

$$(29) \quad \bar{T}/T = [(\gamma - 1)M^2 + 2](2\gamma M^2 - \gamma + 1)/(\gamma + 1)^2 M^2;$$

$$(30) \quad \bar{M}^2 = [(\gamma - 1)M^2 + 2]/[2\gamma M^2 - \gamma + 1].$$

The entropy change across the shock is

$$(31) \quad \Delta s \equiv \bar{s} - s = c_v \ln[(\bar{p}/p)(\rho/\bar{\rho})^\gamma],$$

where $c_v = R/(\gamma - 1)m$ is the specific heat at constant volume; here R is the gas constant. In the weak-shock limit ($M \rightarrow 1$),

$$(32) \quad \Delta s \rightarrow c_v \frac{2\gamma(\gamma - 1)}{3(\gamma + 1)}(M^2 - 1)^3 \approx \frac{16\gamma R}{3(\gamma + 1)m}(M - 1)^3.$$

The radius at time t of a strong spherical blast wave resulting from the explosive release of energy E in a medium with uniform density ρ is

$$(33) \quad R_S = C_0(Et^2/\rho)^{1/5},$$

where C_0 is a constant depending on γ . For $\gamma = 7/5$, $C_0 = 1.033$.

FUNDAMENTAL PLASMA PARAMETERS

All quantities are in Gaussian cgs units except temperature (T , T_e , T_i) expressed in eV and ion mass (m_i) expressed in units of the proton mass, $\mu = m_i/m_p$; Z is charge state; k is Boltzmann's constant; K is wavelength; γ is the adiabatic index; $\ln \Lambda$ is the Coulomb logarithm.

Frequencies

electron gyrofrequency	$f_{ce} = \omega_{ce}/2\pi = 2.80 \times 10^6 B \text{ Hz}$ $\omega_{ce} = eB/m_e c = 1.76 \times 10^7 B \text{ rad/sec}$
ion gyrofrequency	$f_{ci} = \omega_{ci}/2\pi = 1.52 \times 10^3 Z\mu^{-1} B \text{ Hz}$ $\omega_{ci} = eB/m_i c = 9.58 \times 10^3 Z\mu^{-1} B \text{ rad/sec}$
electron plasma frequency	$f_{pe} = \omega_{pe}/2\pi = 8.98 \times 10^3 n_e^{1/2} \text{ Hz}$ $\omega_{pe} = (4\pi n_e e^2/m_e)^{1/2}$ $= 5.64 \times 10^4 n_e^{1/2} \text{ rad/sec}$
ion plasma frequency	$f_{pi} = \omega_{pi}/2\pi$ $= 2.10 \times 10^2 Z\mu^{-1/2} n_i^{1/2} \text{ Hz}$ $\omega_{pi} = (4\pi n_i Z^2 e^2/m_i)^{1/2}$ $= 1.32 \times 10^3 Z\mu^{-1/2} n_i^{1/2} \text{ rad/sec}$
electron trapping rate	$\nu_{Te} = (eKE/m_e)^{1/2}$ $= 7.26 \times 10^8 K^{1/2} E^{1/2} \text{ sec}^{-1}$
ion trapping rate	$\nu_{Ti} = (eKE/m_i)^{1/2}$ $= 1.69 \times 10^7 K^{1/2} E^{1/2} \mu^{-1/2} \text{ sec}^{-1}$
electron collision rate	$\nu_e = 2.91 \times 10^{-6} n_i \ln \Lambda T_e^{-3/2} \text{ sec}^{-1}$
ion collision rate	$\nu_i = 4.78 \times 10^{-8} n_i Z^2 \ln \Lambda T_i^{-3/2} \text{ sec}^{-1}$

Lengths

electron deBrogie length	$\lambda = h/(m_e kT_e)^{1/2} = 2.76 \times 10^{-8} T_e^{-1/2} \text{ cm}$
classical distance of minimum approach	$e^2/kT = 1.44 \times 10^{-7} T^{-1} \text{ cm}$
electron gyroradius	$r_e = v_{Te}/\omega_{ce} = 2.38 T_e^{1/2} B^{-1} \text{ cm}$
ion gyroradius	$r_i = v_{Ti}/\omega_{ci}$ $= 1.02 \times 10^2 \mu^{1/2} Z^{-1} T_i^{1/2} B^{-1} \text{ cm}$
plasma skin depth	$c/\omega_{pe} = 5.31 \times 10^7 n_e^{-1/2} \text{ cm}$
Debye length	$\lambda_D = (kT/4\pi n e^2)^{1/2}$ $= 7.43 \times 10^7 T^{1/2} n^{-1/2} \text{ cm}$

```

\input prolog
\hoffset=1truein\hoffset=1truein\size=6.5truein\size=9truein
\pageno=29
{\headfont Velocities}
\smallskip{\def\CR{\cr\noalign{\vskip2.5pt}}
% \CR SKIPS A SPACE AFTER EACH LINE.
\halign{\quad#\hfil&\quad$\hfil\cr
electron thermal velocity& $v_{Te}=(kT_e/m_e)^{1/2}$ \CR
&\ph{ $v_{Te}=4.19\times10^{-7}T_e^{1/2}$ },\rm cm/sec\CR
ion thermal velocity& $v_{Ti}=(kT_i/m_i)^{1/2}$ \CR
&\ph{ $v_{Ti}=9.79\times10^{-5}\mu^{-1/2}T_i^{1/2}$ },\rm cm/sec\CR
ion sound velocity& $C_s=(\gamma kT_e/m_i)^{1/2}$ \CR
&\ph{ $C_s=9.79\times10^{-5}(\gamma ZT_e/\mu)^{1/2}$ },\rm cm/sec\CR
Alfvén velocity& $v_A=B/(4\pi n_i m_i)^{1/2}$ \CR
&\ph{ $v_A=2.18\times10^{-11}\mu^{-1/2}n_i^{-1/2}B$ },\rm cm/sec\CR
\noalign{\headfont Dimensionless}
\sk
(electron/proton mass ratio)& $(m_e/m_p)^{1/2}=2.33\times10^{-2}=1/42.9$  CR
number of particles in& $(4\pi/3)n\lambda_D^3=1.72\times10^9T^{3/2}n^{-1/2}$ \cr
\quad Debye sphere\CR
Alfvén velocity/speed of light& $v_A/c=7.28\mu^{-1/2}n_i^{-1/2}B$ \CR
electron plasma/gyrofrequency&
\omega_{pe}/\omega_{ce}=3.21\times10^{-3}n_e^{1/2}B^{-1}\cr
\quad ratio\CR
ion plasma/gyrofrequency ratio&
\omega_{pi}/\omega_{ci}=0.137\mu^{1/2}n_i^{1/2}B^{-1}\CR
thermal/magnetic energy ratio&\beta=8\pi nkT/B^2=4.03\times10^{-11}nTB^{-2} CR
magnetic/ion rest energy ratio& $B^2/8\pi n_i m_i c^2=26.5\mu^{-1}n_i^{-1}B^2$  CR
\sk
\noalign{\headfont Miscellaneous}
\sk
Bohm diffusion coefficient& $D_B=(ckT/16eB)$ \CR
&\ph{ $D_B=6.25\times10^{-6}TB^{-1}$ },\rm cm^2/sec\CR
transverse Spitzer resistivity&
\eta_{\perp}=1.15\times10^{-14}Z\ln\Lambda T^{-3/2},\rm sec\CR
&\ph{\eta_{\perp}=1.03\times10^{-2}Z\ln\Lambda T^{-3/2}},\Omega\rm cm\CR}}
\smallskip\noindent
The anomalous collision rate due to low-frequency ion-sound turbulence is

$$\nu\approx\omega_{pe}\widetilde{W}/kT=5.64\times10^{-4}n_e^{1/2}\widetilde{W}/kT,\rm sec^{-1},$$

where  $\widetilde{W}$  is the total energy of waves with  $\omega/K < v_{Ti}$ .
\smallskip\noindent
Magnetic pressure is given by

$$P_{\rm mag}=B^2/8\pi=3.98\times10^{-6}B^2,\rm dynes/cm^2=3.93(B/B_0)^2,\rm atm,$$

where  $B_0=10^6,\rm G=1,T$ .
\smallskip\noindent
Detonation energy of 1 kiloton of high explosive is

$$W_{\rm kT}=10^{12},\rm cal=4.2\times10^{19},\rm erg.$$

\fill\egjectend

```


Velocities

electron thermal velocity	$v_{Te} = (kT_e/m_e)^{1/2}$ $= 4.19 \times 10^7 T_e^{1/2} \text{ cm/sec}$
ion thermal velocity	$v_{Ti} = (kT_i/m_i)^{1/2}$ $= 9.79 \times 10^5 \mu^{-1/2} T_i^{1/2} \text{ cm/sec}$
ion sound velocity	$C_s = (\gamma Z k T_e / m_i)^{1/2}$ $= 9.79 \times 10^5 (\gamma Z T_e / \mu)^{1/2} \text{ cm/sec}$
Alfvén velocity	$v_A = B / (4\pi n_i m_i)^{1/2}$ $= 2.18 \times 10^{11} \mu^{-1/2} n_i^{-1/2} B \text{ cm/sec}$

Dimensionless

(electron/proton mass ratio) ^{1/2}	$(m_e/m_p)^{1/2} = 2.33 \times 10^{-2} = 1/42.9$
number of particles in Debye sphere	$(4\pi/3)n\lambda_D^3 = 1.72 \times 10^9 T^{3/2} n^{-1/2}$
Alfvén velocity/speed of light	$v_A/c = 7.28 \mu^{-1/2} n_i^{-1/2} B$
electron plasma/gyrofrequency ratio	$\omega_{pe}/\omega_{ce} = 3.21 \times 10^{-3} n_e^{1/2} B^{-1}$
ion plasma/gyrofrequency ratio	$\omega_{pi}/\omega_{ci} = 0.137 \mu^{1/2} n_i^{1/2} B^{-1}$
thermal/magnetic energy ratio	$\beta = 8\pi n k T / B^2 = 4.03 \times 10^{-11} n T B^{-2}$
magnetic/ion rest energy ratio	$B^2 / 8\pi n_i m_i c^2 = 26.5 \mu^{-1} n_i^{-1} B^2$

Miscellaneous

Bohm diffusion coefficient	$D_B = (ckT/16eB)$ $= 6.25 \times 10^6 T B^{-1} \text{ cm}^2/\text{sec}$
transverse Spitzer resistivity	$\eta_{\perp} = 1.15 \times 10^{-14} Z \ln \Lambda T^{-3/2} \text{ sec}$ $= 1.03 \times 10^{-2} Z \ln \Lambda T^{-3/2} \Omega \text{ cm}$

The anomalous collision rate due to low-frequency ion-sound turbulence is

$$\nu^* \approx \omega_{pe} \tilde{W} / kT = 5.64 \times 10^4 n_e^{1/2} \tilde{W} / kT \text{ sec}^{-1},$$

where \tilde{W} is the total energy of waves with $\omega/K < v_{Ti}$.

Magnetic pressure is given by

$$P_{\text{mag}} = B^2 / 8\pi = 3.98 \times 10^6 B^2 \text{ dynes/cm}^2 = 3.93 (B/B_0)^2 \text{ atm.}$$

where $B_0 = 10 \text{ kG} = 1 \text{ T}$.

Detonation energy of 1 kiloton of high explosive is

$$W_{\text{kT}} = 10^{12} \text{ cal} = 4.2 \times 10^{19} \text{ erg.}$$

```

\input prolog
\hoffset=1.25truein
\voffset=1.0truein
\hsize=6.0truein
\vsizer=9.0truein
\pageno=30
\centerline{\headfont PLASMA DISPERSION FUNCTION}
\bigskip\N Definition$\{16\}$ (first form valid only for $\operatorname{Im}\zeta>0$):

$$Z(\zeta)=\pi^{-1/2}\int_{-\infty}^{+\infty}dt\, \exp\left(-t^2\right)\over t-\zeta=2i\exp\left(-\zeta^2\right)\int_{-\infty}^{+\infty}i\zeta dt\, \exp\left(-t^2\right).$$

\N Physically, $\zeta=x+iy$ is the ratio of wave phase velocity to thermal velocity.
\medskip\N Differential equation:

$$\{dZ\over d\zeta\}=-2\left(1+\zeta Z\right),\>Z(0)=i\pi^{-1/2};\quad\{d^2Z\over d\zeta^2\}+2\zeta\{dZ\over d\zeta\}+2Z=0.$$

\N Real argument ($y=0$):

$$Z(x)=\exp\left(-x^2\right)\left[i\pi^{-1/2}-2\int_0^x dt\, \exp\left(t^2\right)\right].$$

\N Imaginary argument ($x=0$):

$$Z(iy)=i\pi^{-1/2}\exp\left(y^2\right)\left[1-\operatorname{erf}(y)\right].$$

\N Power series (small argument):

$$Z(\zeta)=i\pi^{-1/2}\exp\left(-\zeta^2\right)-2\zeta\left(1-2\zeta^2/3+4\zeta^4/15-8\zeta^6/105+\cdots\right).$$

\N Asymptotic series, $\mid\zeta\mid\gg 1$ (Ref. 17):

$$Z(\zeta)=i\pi^{-1/2}\sigma\exp\left(-\zeta^2\right)-\zeta^{-1}\left(1+1/2\zeta^{-2}+3/4\zeta^{-4}+15/8\zeta^{-6}+\cdots\right),$$

where

$$\sigma=\left\{\begin{array}{l} \operatorname{erfc}(\zeta) \\ \operatorname{erfc}(\zeta) \end{array}\right.$$

\N Symmetry properties (the asterisk denotes complex conjugation):

$$Z(\zeta^*)=-\left[Z(-\zeta)\right]^*.$$


$$Z(\zeta)=\left[Z(\zeta)\right]^*+2i\pi^{-1/2}\exp\left[-\zeta^2\right]\quad(\operatorname{Im}\zeta>0).$$

\N Two-pole approximations$\{18\}$ (good for $\zeta$ in upper half plane except when $\operatorname{Im}\zeta<\pi^{1/2}x^2\exp(-x^2)$, $\operatorname{Re}\zeta\gg 1$):

$$Z(\zeta)\approx\frac{0.50+0.81i}{a-\zeta}-\frac{0.50-0.81i}{a+\zeta},\quad a=0.51-0.81i;$$


$$Z(\zeta)\approx\frac{0.50+0.96i}{b-\zeta^2}+\frac{0.50-0.96i}{b+\zeta^2},\quad b=0.48-0.91i.$$

\end

```

PLASMA DISPERSION FUNCTION

Definition¹⁶ (first form valid only for $\text{Im } \zeta > 0$):

$$Z(\zeta) = \pi^{-1/2} \int_{-\infty}^{+\infty} \frac{dt \exp(-t^2)}{t - \zeta} = 2i \exp(-\zeta^2) \int_{-\infty}^{i\zeta} dt \exp(-t^2).$$

Physically, $\zeta = x + iy$ is the ratio of wave phase velocity to thermal velocity.

Differential equation:

$$\frac{dZ}{d\zeta} = -2(1 + \zeta Z), \quad Z(0) = i\pi^{1/2}; \quad \frac{d^2 Z}{d\zeta^2} + 2\zeta \frac{dZ}{d\zeta} + 2Z = 0.$$

Real argument ($y = 0$):

$$Z(x) = \exp(-x^2) \left(i\pi^{1/2} - 2 \int_0^x dt \exp(t^2) \right).$$

Imaginary argument ($x = 0$):

$$Z(iy) = i\pi^{1/2} \exp(y^2) [1 - \text{erf}(y)].$$

Power series (small argument):

$$Z(\zeta) = i\pi^{1/2} \exp(-\zeta^2) - 2\zeta \left(1 - 2\zeta^2/3 + 4\zeta^4/15 - 8\zeta^6/105 + \dots \right).$$

Asymptotic series, $|\zeta| \gg 1$ (Ref. 17):

$$Z(\zeta) = i\pi^{1/2} \sigma \exp(-\zeta^2) - \zeta^{-1} \left(1 + 1/2\zeta^2 + 3/4\zeta^4 + 15/8\zeta^6 + \dots \right),$$

where

$$\sigma = \begin{cases} 0 & y > |x|^{-1} \\ 1 & |y| < |x|^{-1} \\ 2 & y < -|x|^{-1} \end{cases}$$

Symmetry properties (the asterisk denotes complex conjugation):

$$Z(\zeta^*) = -[Z(-\zeta)]^*:$$

$$Z(\zeta^*) = [Z(\zeta)]^* + 2i\pi^{1/2} \exp[-(\zeta^*)^2] \quad (y > 0).$$

Two-pole approximations¹⁸ (good for ζ in upper half plane except when $y < \pi^{1/2} x^2 \exp(-x^2)$, $x \gg 1$):

$$Z(\zeta) \approx \frac{0.50 + 0.81i}{a - \zeta} - \frac{0.50 - 0.81i}{a^* + \zeta}, \quad a = 0.51 - 0.81i;$$

$$Z'(\zeta) \approx \frac{0.50 + 0.96i}{(b - \zeta)^2} + \frac{0.50 - 0.96i}{(b^* + \zeta)^2}, \quad b = 0.48 - 0.91i.$$

```

\input prolog
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\pageno=31
\centerline{\headfont COLLISIONS AND TRANSPORT} \bsk\indent
Temperatures are in eV; the corresponding value of Boltzmann's constant
is  $k = 1.60 \times 10^{-12}$  erg/eV; masses  $\mu$ ,  $\mu'$  are in units of
the proton mass;  $e_\alpha = Z_\alpha e$  is the charge of species  $\alpha$ .
All other units are cgs except where noted. \msk
{\headfont Relaxation Rates} \ssk\indent
Rates are associated with four relaxation processes arising from the
interaction of test particles (labeled  $\alpha$ ) streaming with velocity  $\mathbf{v}_\alpha$ 
through a background of field particles (labeled  $\beta$ ):

$$\begin{aligned} \nu_s &= \text{slowing down} \\ \nu_\perp &= \text{transverse diffusion} \\ \nu_\parallel &= \text{parallel diffusion} \\ \nu_\epsilon &= \text{energy loss} \end{aligned}$$

where the averages are performed over an ensemble of test particles and
a Maxwellian field particle distribution. The exact formulas may be
written

$$\begin{aligned} \nu_s &= (1 + m_\alpha/m_\beta) \psi(x) \\ \nu_\perp &= 2 \left[ (1 - 1/2 x) \psi(x) + \psi'(x) \right] \\ \nu_\parallel &= \left[ \psi(x)/x \right] \nu_0 \\ \nu_\epsilon &= 2 \left[ (m_\alpha/m_\beta) \psi(x) - \psi'(x) \right] \nu_0 \end{aligned}$$

where

$$\begin{aligned} \nu_0 &= 4\pi \{e_\alpha\}^2 \{e_\beta\}^2 \lambda_{\alpha\beta} n_\beta / \\ & \{m_\alpha\}^2 \{v_\alpha\}^3 ; \quad x = m_\beta \{v_\alpha\}^2 / \\ & 2kT_\beta \end{aligned}$$


$$\psi(x) = \frac{2}{\sqrt{\pi}} \int_0^x dt e^{-t^2} ; \quad \psi'(x) = \frac{d\psi}{dx}$$

and  $\lambda_{\alpha\beta} = \ln \Lambda_{\alpha\beta}$  is the Coulomb
logarithm (see below). Limiting forms of  $\nu_s$ ,  $\nu_\perp$  and
 $\nu_\parallel$  are given in the following table. All the expressions shown
\vfilleject\end

```

COLLISIONS AND TRANSPORT

Temperatures are in eV; the corresponding value of Boltzmann's constant is $k = 1.60 \times 10^{-12}$ erg/eV; masses μ, μ' are in units of the proton mass; $e_\alpha = Z_\alpha e$ is the charge of species α . All other units are cgs except where noted.

Relaxation Rates

Rates are associated with four relaxation processes arising from the interaction of test particles (labeled α) streaming with velocity \mathbf{v}_α through a background of field particles (labeled β):

slowing down	$\frac{d\mathbf{v}_\alpha}{dt} = -\nu_s^{\alpha/\beta} \mathbf{v}_\alpha$
transverse diffusion	$\frac{d}{dt}(\mathbf{v}_\alpha - \bar{\mathbf{v}}_\alpha)_\perp^2 = \nu_\perp^{\alpha/\beta} v_\alpha^2$
parallel diffusion	$\frac{d}{dt}(\mathbf{v}_\alpha - \bar{\mathbf{v}}_\alpha)_\parallel^2 = \nu_\parallel^{\alpha/\beta} v_\alpha^2$
energy loss	$\frac{d}{dt} v_\alpha^2 = -\nu_\epsilon^{\alpha/\beta} v_\alpha^2$

where the averages are performed over an ensemble of test particles and a Maxwellian field particle distribution. The exact formulas may be written¹⁹

$$\begin{aligned} \nu_s^{\alpha/\beta} &= (1 + m_\alpha/m_\beta) \psi(x^{\alpha/\beta}) \nu_0^{\alpha/\beta}; \\ \nu_\perp^{\alpha/\beta} &= 2 \left[(1 - 1/2 x^{\alpha/\beta}) \psi(x^{\alpha/\beta}) + \psi'(x^{\alpha/\beta}) \right] \nu_0^{\alpha/\beta}; \\ \nu_\parallel^{\alpha/\beta} &= \left[\psi(x^{\alpha/\beta}) / x^{\alpha/\beta} \right] \nu_0^{\alpha/\beta}; \\ \nu_\epsilon^{\alpha/\beta} &= 2 \left[(m_\alpha/m_\beta) \psi(x^{\alpha/\beta}) - \psi'(x^{\alpha/\beta}) \right] \nu_0^{\alpha/\beta}, \end{aligned}$$

where

$$\nu_0^{\alpha/\beta} = 4\pi e_\alpha^2 e_\beta^2 \lambda_{\alpha\beta} n_\beta / m_\alpha^2 v_\alpha^3; \quad x^{\alpha/\beta} = m_\beta v_\alpha^2 / 2kT_\beta;$$

$$\psi(x) = \frac{2}{\sqrt{\pi}} \int_0^x dt t^{1/2} e^{-t}; \quad \psi'(x) = \frac{d\psi}{dx},$$

and $\Lambda_{\alpha\beta} = \ln \Lambda_{\alpha\beta}$ is the Coulomb logarithm (see below). Limiting forms of ν_s, ν_\perp and ν_\parallel are given in the following table. All the expressions shown

```

\input prolog \pageno=32
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
have units cm3\,$sec-1$. Test particle energy  $\epsilon$  and field
particle temperature  $T$  are both in eV;  $\mu = m_i/m_p$  where  $m_p$  is the
proton mass;  $Z$  is ion charge state; in electron--electron and ion--ion
encounters, field particle quantities are distinguished by a prime. The two
expressions given below for each rate hold for very slow  $(x \ll 1)$  and
very fast  $(x \gg 1)$  test particles, respectively. \msk
% \app AND \lra ARE BOTH DEFINED IN PROLOG.TEX
\halign{\quad#\hfil &#\hfil &#\hfil \cr
&\hfil\undertext{Slow} &\hfil\undertext{Fast} \cr
\kern-1em Electron--electron \hidewidth \cr
\displaystyle \nu_s\{e/e\}/n_e\lambda_{ee}\} \approx 5.8 \times 10^{-6}
T^{-3/2} \lra 7.7 \times 10^{-6} \epsilon^{-3/2} \cr
\displaystyle \nu_{\perp}\{e/e\}/n_e\lambda_{ee}\} \approx 5.8 \times 10^{-6}
T^{-1/2} \epsilon^{-1} \lra 7.7 \times 10^{-6} \epsilon^{-3/2} \cr
\displaystyle \nu_{\parallel}\{e/e\}/n_e\lambda_{ee}\} \approx 2.9 \times
10^{-6} T^{-1/2} \epsilon^{-1} \lra 3.9 \times 10^{-6} T \epsilon^{-5/2} \cr
\halign{\smallskip Electron--ion \smallskip \cr
\displaystyle \nu_s\{e/i\}/n_i Z^2 \lambda_{ei}\} \approx 0.23 \mu^{3/2} T^{-3/2}
\lra 3.9 \times 10^{-6} \epsilon^{-3/2} \cr
\displaystyle \nu_{\perp}\{e/i\}/n_i Z^2 \lambda_{ei}\} \approx 2.5 \times 10^{-4}
\mu^{1/2} T^{-1/2} \epsilon^{-1} \lra 7.7 \times 10^{-6} \epsilon^{-3/2} \cr
\displaystyle \nu_{\parallel}\{e/i\}/n_i Z^2 \lambda_{ei}\} \approx 1.2
\times 10^{-4} \mu^{1/2} T^{-1/2} \epsilon^{-1} \lra 2.1 \times
10^{-9} \mu^{-1} T \epsilon^{-5/2} \cr
\halign{\smallskip Ion--electron \smallskip \cr
\displaystyle \nu_s\{i/e\}/n_e Z^2 \lambda_{ie}\} \approx 1.6 \times 10^{-9} \mu^{-1}
T^{-3/2} \lra 1.7 \times 10^{-4} \mu^{1/2} \epsilon^{-3/2} \cr
\displaystyle \nu_{\perp}\{i/e\}/n_e Z^2 \lambda_{ie}\} \approx 3.2
\times 10^{-9} \mu^{-1} T^{-1/2} \epsilon^{-1} \lra
1.8 \times 10^{-7} \mu^{-1/2} \epsilon^{-3/2} \cr
\displaystyle \nu_{\parallel}\{i/e\}/n_e Z^2 \lambda_{ie}\} \approx 1.6
\times 10^{-9} \mu^{-1} T^{-1/2} \epsilon^{-1} \lra
1.7 \times 10^{-4} \mu^{1/2} T \epsilon^{-5/2} \cr
\halign{\smallskip Ion--ion \smallskip \cr
\displaystyle \{\nu_s\{i/i\} \over n_i Z^2 Z'^2 \lambda_{ii'}\} \approx
6.8 \times 10^{-8} \{\mu^{1/2} \over \mu'\}
\left(1 + \{\mu' \over \mu\} \right) T^{-3/2} \lra \left(1 + \{\mu' \over \mu\} \right)
\left\{\mu^{1/2} \over \epsilon^{3/2}\right\} \cr
\displaystyle \{\nu_{\perp}\{i/i\} \over n_i Z^2 Z'^2 \lambda_{ii'}\} \approx
1.4 \times 10^{-7} \{\mu^{1/2} \over \mu'\} T^{-1/2}
\epsilon^{-1} \lra \left\{\mu^{1/2} \over \epsilon^{3/2}\right\} \cr
\displaystyle \{\nu_{\parallel}\{i/i\} \over n_i Z^2 Z'^2 \lambda_{ii'}\} \approx
6.8 \times 10^{-8} \{\mu^{1/2} \over \mu'\} T^{-1/2}
\epsilon^{-1} \lra \left\{\mu^{1/2} \over \epsilon^{3/2}\right\} \cr
\lra 9.0 \times 10^{-8} \mu^{1/2} \mu'^{-1} T \epsilon^{-5/2} \cr} \msk \halign{\cr
In the same limits, the energy transfer rate follows from the identity
 $\nu_{\parallel} \epsilon = 2 \nu_{\perp} - \nu_{\parallel} - \nu_{\parallel}'$ ,
except for the case of fast electrons or fast ions scattered by ions, where
the leading terms cancel. Then the appropriate forms are
 $\nu_{\parallel} \epsilon \rightarrow 4.2 \times 10^{-9} n_i Z^2 \lambda_{ei} \epsilon$ 
 $\left[ \epsilon^{-3/2} \mu^{-1} - 8.0 \times 10^{-4} (\mu T)^{-1/2} \epsilon^{-1} \right]$ ,
 $(\mu \ll 1)$ 
\fill\eggedend

```

have units $\text{cm}^3 \text{sec}^{-1}$. Test particle energy ϵ and field particle temperature T are both in eV; $\mu = m_i/m_p$ where m_p is the proton mass; Z is ion charge state; in electron-electron and ion-ion encounters, field particle quantities are distinguished by a prime. The two expressions given below for each rate hold for very slow ($x^{\alpha/\beta} \ll 1$) and very fast ($x^{\alpha/\beta} \gg 1$) test particles, respectively.

	<u>Slow</u>	<u>Fast</u>
Electron-electron		
$\nu_s^{e/e'}/n_{e'}\lambda_{ee'}$	$\approx 5.8 \times 10^{-6} T^{-3/2}$	$\longrightarrow 7.7 \times 10^{-6} \epsilon^{-3/2}$
$\nu_{\perp}^{e/e'}/n_{e'}\lambda_{ee'}$	$\approx 5.8 \times 10^{-6} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 7.7 \times 10^{-6} \epsilon^{-3/2}$
$\nu_{\parallel}^{e/e'}/n_{e'}\lambda_{ee'}$	$\approx 2.9 \times 10^{-6} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 3.9 \times 10^{-6} T \epsilon^{-5/2}$
Electron-ion		
$\nu_s^{e/i}/n_i Z^2 \lambda_{ei}$	$\approx 0.23 \mu^{3/2} T^{-3/2}$	$\longrightarrow 3.9 \times 10^{-6} \epsilon^{-3/2}$
$\nu_{\perp}^{e/i}/n_i Z^2 \lambda_{ei}$	$\approx 2.5 \times 10^{-4} \mu^{1/2} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 7.7 \times 10^{-6} \epsilon^{-3/2}$
$\nu_{\parallel}^{e/i}/n_i Z^2 \lambda_{ei}$	$\approx 1.2 \times 10^{-4} \mu^{1/2} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 2.1 \times 10^{-9} \mu^{-1} T \epsilon^{-5/2}$
Ion-electron		
$\nu_s^{i/e}/n_e Z^2 \lambda_{ie}$	$\approx 1.6 \times 10^{-9} \mu^{-1} T^{-3/2}$	$\longrightarrow 1.7 \times 10^{-4} \mu^{1/2} \epsilon^{-3/2}$
$\nu_{\perp}^{i/e}/n_e Z^2 \lambda_{ie}$	$\approx 3.2 \times 10^{-9} \mu^{-1} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 1.8 \times 10^{-7} \mu^{-1/2} \epsilon^{-3/2}$
$\nu_{\parallel}^{i/e}/n_e Z^2 \lambda_{ie}$	$\approx 1.6 \times 10^{-9} \mu^{-1} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 1.7 \times 10^{-4} \mu^{1/2} T \epsilon^{-5/2}$
Ion-ion		
$\frac{\nu_s^{i/i'}}{n_{i'} Z^2 Z'^2 \lambda_{ii'}}$	$\approx 6.8 \times 10^{-8} \frac{\mu'^{1/2}}{\mu} \left(1 + \frac{\mu'}{\mu}\right) T^{-3/2}$	$\longrightarrow 9.0 \times 10^{-8} \left(\frac{1}{\mu} + \frac{1}{\mu'}\right) \frac{\mu^{1/2}}{\epsilon^{3/2}}$
$\frac{\nu_{\perp}^{i/i'}}{n_{i'} Z^2 Z'^2 \lambda_{ii'}}$	$\approx 1.4 \times 10^{-7} \mu'^{1/2} \mu^{-1} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 1.8 \times 10^{-7} \mu^{-1/2} \epsilon^{-3/2}$
$\frac{\nu_{\parallel}^{i/i'}}{n_{i'} Z^2 Z'^2 \lambda_{ii'}}$	$\approx 6.8 \times 10^{-8} \mu'^{1/2} \mu^{-1} T^{-1/2} \epsilon^{-1}$	$\longrightarrow 9.0 \times 10^{-8} \mu^{1/2} \mu'^{-1} T \epsilon^{-5/2}$

In the same limits, the energy transfer rate follows from the identity

$$\nu_e = 2\nu_s - \nu_{\perp} - \nu_{\parallel},$$

except for the case of fast electrons or fast ions scattered by ions, where the leading terms cancel. Then the appropriate forms are

$$\nu_e^{i/i} \longrightarrow 4.2 \times 10^{-9} n_i Z^2 \lambda_{ei} \left[\epsilon^{-3/2} \mu^{-1} - 8.9 \times 10^4 (\mu/T)^{1/2} \epsilon^{-1} \exp(-1836 \mu \epsilon / T) \right] \text{sec}^{-1}$$

```

\input prolog
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\pageno=33
and

$$\nu_{\epsilon} \rightarrow 1.8 \times 10^{-7} n_i Z^2 \lambda_{ii} \left[ \epsilon^{-3/2} \mu^{1/2} / \mu' - 1.1 (\mu'/T)^{1/2} \epsilon^{-1} \exp(-\mu' \epsilon / T) \right] \text{sec}^{-1}$$

In general, the energy transfer rate  $\nu_{\epsilon}$  is positive for  $\epsilon > \epsilon_{\alpha}$  and negative for  $\epsilon < \epsilon_{\alpha}$ ,
where  $x \epsilon_{\alpha} = (m_{\beta}/m_{\alpha}) \epsilon_{\alpha} / T_{\beta}$  is the solution of  $\psi'(x \epsilon_{\alpha}) = (m_{\alpha}/m_{\beta}) \psi(x \epsilon_{\alpha})$ . The ratio  $\epsilon_{\alpha} / T_{\beta}$  is given for a number of specific  $\alpha$ ,  $\beta$  in the following table:
% BEGINNING OF TABLE

$$\begin{array}{c} \epsilon_{\alpha} / T_{\beta} \\ \hline \begin{array}{c} \text{He}^3 \\ \text{He}^4 \\ \text{H} \\ \text{D} \\ \text{T} \end{array} \end{array}$$


$$\begin{array}{c} 1.5 \\ 0.98 \\ 4.8 \times 10^{-3} \\ 2.6 \times 10^{-3} \\ 1.8 \times 10^{-3} \end{array}$$


$$\begin{array}{c} 1.4 \times 10^{-3} \\ 1.4 \times 10^{-3} \end{array}$$

\end{array}
When both species are near Maxwellian, with  $T_i \approx T_e$ , there are just two characteristic collision rates. For  $Z = 1$ ,

$$\nu_e \approx 2.9 \times 10^{-6} n \lambda_{Te}^{-3/2} \text{sec}^{-1}$$


$$\nu_i \approx 4.8 \times 10^{-8} n \lambda_{Ti}^{-3/2} \mu^{-1/2} \text{sec}^{-1}$$

\smallskip
\headfont Temperature Isotropization
\smallskip\indent
Isotropization is described by

$$\frac{dT_{\perp}}{dt} = - \frac{1}{2} \frac{dT_{\parallel}}{dt} = -\nu_T T_{\alpha} (T_{\perp} - T_{\parallel}),$$

where, if  $A \equiv T_{\perp}/T_{\parallel} - 1 > 0$ ,

$$\nu_T T_{\alpha} = \frac{2 \sqrt{\pi} e_{\alpha}^2 e_{\beta}^2 n_{\alpha} \lambda_{\alpha\beta} \beta_{\alpha} \over m_{\alpha}^{1/2} (k T_{\parallel})^{3/2} A^{-2} \left[ -3 + (A+3) \frac{\tan^{-1}(A^{1/2})}{A^{1/2}} \right] \text{right}}{A^{1/2}}$$

If  $A < 0$ ,  $\tan^{-1}(A^{1/2})/A^{1/2}$  is replaced by  $\tanh^{-1}(-A^{1/2})/(-A^{1/2})$ .
For  $T_{\perp} \approx T_{\parallel} \approx T$ ,

$$\nu_T \approx 8.2 \times 10^{-7} n \lambda T^{-3/2} \text{sec}^{-1}$$


$$\nu_{Ti} \approx 1.9 \times 10^{-3} n \lambda Z^2 \mu^{-1/2} T^{-3/2} \text{sec}^{-1}$$

\end{array}
\end{pre>

```


and

$$\nu_{\epsilon}^{i/i'} \longrightarrow 1.8 \times 10^{-7} n_{i'} Z^2 Z'^2 \lambda_{ii'} \left[\epsilon^{-3/2} \mu^{1/2} / \mu' - 1.1 (\mu' / T)^{1/2} \epsilon^{-1} \exp(-\mu' \epsilon / T) \right] \text{ sec}^{-1}.$$

In general, the energy transfer rate $\nu_{\epsilon}^{\alpha/\beta}$ is positive for $\epsilon > \epsilon_{\alpha}^*$ and negative for $\epsilon < \epsilon_{\alpha}^*$, where $x^* = (m_{\beta}/m_{\alpha})\epsilon_{\alpha}^*/T_{\beta}$ is the solution of $\psi'(x^*) = (m_{\alpha}/m_{\beta})\psi(x^*)$. The ratio $\epsilon_{\alpha}^*/T_{\beta}$ is given for a number of specific α, β in the following table:

α/β	i/e	$e/e, i/i$	e/p	e/D	$e/T, e/\text{He}^3$	e/He^4
$\frac{\epsilon_{\alpha}^*}{T_{\beta}}$	1.5	0.98	4.8×10^{-3}	2.6×10^{-3}	1.8×10^{-3}	1.4×10^{-3}

When both species are near Maxwellian, with $T_i \lesssim T_e$, there are just two characteristic collision rates. For $Z = 1$,

$$\begin{aligned} \nu_{\epsilon} &= 2.9 \times 10^{-6} n \lambda T_e^{-3/2} \text{ sec}^{-1}; \\ \nu_i &= 4.8 \times 10^{-8} n \lambda T_i^{-3/2} \mu^{-1/2} \text{ sec}^{-1}. \end{aligned}$$

Temperature Isotropization

Isotropization is described by

$$\frac{dT_{\perp}}{dt} = -\frac{1}{2} \frac{dT_{\parallel}}{dt} = -\nu_T^{\alpha} (T_{\perp} - T_{\parallel}),$$

where, if $A \equiv T_{\perp}/T_{\parallel} - 1 > 0$,

$$\nu_T^{\alpha} = \frac{2\sqrt{\pi} c_{\alpha}^2 c_{\beta}^2 n_{\alpha} \lambda_{\alpha\beta}}{m_{\alpha}^{1/2} (kT_{\parallel})^{3/2}} A^{-2} \left[-3 + (A + 3) \frac{\tan^{-1}(A^{1/2})}{A^{1/2}} \right].$$

If $A < 0$, $\tan^{-1}(A^{1/2})/A^{1/2}$ is replaced by $\tanh^{-1}(-A)^{1/2}/(-A)^{1/2}$. For $T_{\perp} \approx T_{\parallel} \equiv T$,

$$\begin{aligned} \nu_T^{\epsilon} &= 8.2 \times 10^{-7} n \lambda T^{-3/2} \text{ sec}^{-1}; \\ \nu_T^i &= 1.9 \times 10^{-8} n \lambda Z^2 \mu^{-1/2} T^{-3/2} \text{ sec}^{-1}. \end{aligned}$$

\input prolog \pageno=34
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\smallskip {\headfont Thermal Equilibration} \smallskip\indent
If the components of a plasma have different temperatures, but no relative drift, equilibration is described by

$$\frac{dT_\alpha}{dt} = \sum_\beta \bar{\nu}_{\alpha\beta} \epsilon_{\alpha\beta} AOB(T_\beta - T_\alpha),$$
where

$$\bar{\nu}_{\alpha\beta} \epsilon_{\alpha\beta} AOB = 1.8 \times 10^{-19} \{(m_\alpha m_\beta)^{1/2} Z_\alpha^2 Z_\beta^2 \nu_{\alpha\beta} \lambda_{\alpha\beta} / (m_\alpha T_\alpha + m_\beta T_\beta)^{3/2}\} \text{ cm}^3 \text{ sec}^{-1}.$$
For electrons and ions with $T_e \approx T_i \equiv T$, this implies

$$\bar{\nu}_{\alpha\beta} \epsilon_{\alpha\beta} \{e_i/n_i = \bar{\nu}_{\alpha\beta} \epsilon_{\alpha\beta} \{i/e\}/n_e = 3.2 \times 10^{-9} Z_\alpha^2 \lambda_{\alpha\beta} / \mu T^{3/2} \text{ cm}^3 \text{ sec}^{-1}\}.$$
\smallskip {\headfont Coulomb Logarithm}
\smallskip\indent
For test particles of mass m_α and charge Ze_α scattering off field particles of mass m_β and charge Ze_β , the Coulomb logarithm is defined as $\lambda_{\alpha\beta} = \ln(L_{\alpha\beta} / r_{\min})$. Here r_{\min} is the larger of $e_\alpha e_\beta / m_\alpha m_\beta \bar{u}^2$ and $\hbar/2 m_\alpha m_\beta \bar{u}$, averaged over both particle velocity distributions, where $m_\alpha m_\beta \bar{u} = m_\alpha m_\beta / (m_\alpha + m_\beta)$ and $\bar{u} = \sqrt{\langle v_\alpha^2 \rangle + \langle v_\beta^2 \rangle}$; $r_{\max} = (4\pi \sum_\gamma n_\gamma e_\gamma^2 / kT_\gamma)^{-1/2}$, where the summation extends over all species γ for which $\bar{u}^2 < \langle v_\gamma^2 \rangle$; kT_γ / m_γ . If this inequality cannot be satisfied, or if either $\bar{u} / \omega_{c\alpha} < r_{\min}$ or $\bar{u} / \omega_{c\beta} < r_{\max}$, the theory breaks down. Typically $\lambda_{\alpha\beta} \approx 10-20$. Corrections to the transport coefficients are $O(\lambda_{\alpha\beta}^{-1})$; hence the theory is good only to $\sim 10\%$ and fails when $\lambda_{\alpha\beta} \sim 1$.
\indent
The following cases are of particular interest:
\smallskip
(a) Thermal electron--electron collisions

$$\lambda_{ee} = 23 - \ln(n_e^{1/2} T_e^{-3/2}),$$

$$T_e \approx 10 \text{ eV};$$
\smallskip

$$= 24 - \ln(n_e^{1/2} T_e^{-1}),$$

$$T_e \approx 10 \text{ eV}.$$
(b) Electron--ion collisions

$$\lambda_{ei} = \lambda_{ie} = 23 - \ln\left(n_e^{1/2} Z T_e^{-3/2} \right),$$

$$T_{im_e/m_i} < T_e < 10 Z^2,$$
\smallskip

$$\lambda_{ei} = \lambda_{ie} = 24 - \ln\left(n_e^{1/2} T_e^{-1} \right),$$

$$T_i m_e/m_i < 10 Z^2, T_e < T_i$$
\smallskip

$$\lambda_{ei} = \lambda_{ie} = 30 - \ln\left(n_i^{1/2} T_i^{-3/2} Z^2 \mu^{-1} \right),$$

$$T_e < T_i Z m_e/m_i.$$
\input eject\end

Thermal Equilibration

If the components of a plasma have different temperatures, but no relative drift, equilibration is described by

$$\frac{dT_\alpha}{dt} = \sum_\beta \bar{\nu}_\epsilon^{\alpha/\beta} (T_\beta - T_\alpha),$$

where

$$\bar{\nu}_\epsilon^{\alpha/\beta} = 1.8 \times 10^{-19} \frac{(m_\alpha m_\beta)^{1/2} Z_\alpha^2 Z_\beta^2 n_\beta \lambda_{\alpha\beta}}{(m_\alpha T_\beta + m_\beta T_\alpha)^{3/2}} \text{ sec}^{-1}.$$

For electrons and ions with $T_e \approx T_i \equiv T$, this implies

$$\bar{\nu}_\epsilon^{e/i} / n_i = \bar{\nu}_\epsilon^{i/e} / n_e = 3.2 \times 10^{-9} Z^2 \lambda / \mu T^{3/2} \text{ cm}^3 \text{ sec}^{-1}.$$

Coulomb Logarithm

For test particles of mass m_α and charge $e_\alpha = Z_\alpha e$ scattering off field particles of mass m_β and charge $e_\beta = Z_\beta e$, the Coulomb logarithm is defined as $\lambda = \ln \Lambda \equiv \ln(r_{\max}/r_{\min})$. Here r_{\min} is the larger of $e_\alpha e_\beta / m_{\alpha\beta} \bar{u}^2$ and $\hbar/2m_{\alpha\beta} \bar{u}$, averaged over both particle velocity distributions, where $m_{\alpha\beta} = m_\alpha m_\beta / (m_\alpha + m_\beta)$ and $\mathbf{u} = \mathbf{v}_\alpha - \mathbf{v}_\beta$; $r_{\max} = (4\pi \sum_\gamma n_\gamma e_\gamma^2 / kT_\gamma)^{-1/2}$, where the summation extends over all species γ for which $\bar{u}^2 < v_{T_\gamma}^2 = kT_\gamma / m_\gamma$. If this inequality cannot be satisfied, or if either $\bar{u} \omega_{c\alpha}^{-1} < r_{\max}$ or $\bar{u} \omega_{c\beta}^{-1} < r_{\max}$, the theory breaks down. Typically $\lambda \approx 10$ –20. Corrections to the transport coefficients are $O(\lambda^{-1})$; hence the theory is good only to $\sim 10\%$ and fails when $\lambda \sim 1$.

The following cases are of particular interest:

(a) Thermal electron-electron collisions

$$\begin{aligned} \lambda_{ee} &= 23 - \ln(n_e^{1/2} T_e^{-3/2}), & T_e &\lesssim 10 \text{ eV}; \\ &= 24 - \ln(n_e^{1/2} T_e^{-1}), & T_e &\gtrsim 10 \text{ eV}. \end{aligned}$$

(b) Electron-ion collisions

$$\begin{aligned} \lambda_{ei} = \lambda_{ie} &= 23 - \ln(n_e^{1/2} Z T_e^{-3/2}), & T_i m_e / m_i < T_e < 10 Z^2 \text{ eV}; \\ &= 24 - \ln(n_e^{1/2} T_e^{-1}), & T_i m_e / m_i < 10 Z^2 \text{ eV} < T_e; \\ &= 30 - \ln(n_e^{1/2} T_i^{-3/2} Z^2 \mu^{-1}), & T_e < T_i Z m_e / m_i. \end{aligned}$$

```

\input prolog
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\pageno=35
(c) Mixed ion--ion collisions

$$\lambda_{ii'} = \lambda_{i'i} = 23 - \ln \left[ \frac{ZZ'(\mu + \mu')}{\mu T_{i'} + \mu' T_i} \left( \frac{n_i Z^2}{T_i} + \frac{n_{i'} Z'^2}{T_{i'}} \right)^{1/2} \right]$$

(d) Counterstreaming ions (relative velocity  $v_D = \beta_D c$ ) in the presence
of warm electrons,  $kT_i/m_i, kT_{i'}/m_{i'} < v_D^2 < kT_e/m_e$ 

$$\lambda_{ii'} = \lambda_{i'i} = 35 - \ln \left[ \frac{ZZ'(\mu + \mu')}{\mu \mu' \beta_D^2} \left( \frac{n_e}{T_e} \right)^{1/2} \right]$$

\smallskip
{\headfont Fokker-Planck Equation}
\smallskip

$$\left\{ \frac{Df}{Dt} \right\} \equiv \frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \mathbf{F} \cdot \nabla_{\mathbf{v}} f = \left( \frac{\partial f}{\partial t} \right)_{\text{coll}}$$

where  $\mathbf{F}$  is an external force field. The general form of the collision
integral is  $\left( \frac{\partial f}{\partial t} \right)_{\text{coll}} = -\sum_{\beta} \nabla_{\mathbf{v}} \cdot \mathbf{J}^{\beta}$ , with

$$\mathbf{J}^{\beta} = 2\pi \lambda_{\alpha\beta} \{ e_{\alpha}^2 e_{\beta}^2 \int m_{\alpha} \int d^3v' (u^2 \mathbf{I} - \mathbf{u}\mathbf{u}) u^{-3} \cdot \left[ \frac{m_{\beta}}{m_{\alpha}} f^{\beta}(\mathbf{v}) \nabla_{\mathbf{v}'} f^{\beta}(\mathbf{v}') - \frac{m_{\alpha}}{m_{\beta}} f^{\alpha}(\mathbf{v}') \nabla_{\mathbf{v}} f^{\alpha}(\mathbf{v}) \right] \}$$

(Landau form) where  $\mathbf{u} = \mathbf{v} - \mathbf{v}'$  and  $\mathbf{I}$  is the unit dyad, or
alternatively,

$$\mathbf{J}^{\beta} = 4\pi \lambda_{\alpha\beta} \{ e_{\alpha}^2 e_{\beta}^2 \int m_{\alpha}^2 \left[ f^{\alpha}(\mathbf{v}) \nabla_{\mathbf{v}} H(\mathbf{v}) - \frac{1}{2} \nabla_{\mathbf{v}} \cdot \left[ f^{\alpha}(\mathbf{v}) \nabla_{\mathbf{v}} G(\mathbf{v}) \right] \right] \}$$

where the Rosenbluth potentials are

$$G(\mathbf{v}) = \int f^{\beta}(\mathbf{v}') u d^3v'$$


$$H(\mathbf{v}) = \left( 1 + \frac{m_{\alpha}}{m_{\beta}} \right) \int f^{\beta}(\mathbf{v}') u^{-1} d^3v'$$

\vfilleject\end

```

(c) Mixed ion-ion collisions

$$\lambda_{ii'} = \lambda_{i'i} = 23 - \ln \left[\frac{ZZ'(\mu + \mu')}{\mu T_{i'} + \mu' T_i} \left(\frac{n_i Z^2}{T_i} + \frac{n_{i'} Z'^2}{T_{i'}} \right)^{1/2} \right].$$

(d) Counterstreaming ions (relative velocity $v_D = \beta_D c$) in the presence of warm electrons, $kT_i/m_i, kT_{i'}/m_{i'} < v_D^2 < kT_e/m_e$

$$\lambda_{ii'} = \lambda_{i'i} = 35 - \ln \left[\frac{ZZ'(\mu + \mu')}{\mu\mu'\beta_D^2} \left(\frac{n_e}{T_e} \right)^{1/2} \right].$$

Fokker-Planck Equation

$$\frac{Df^\alpha}{Dt} \equiv \frac{\partial f^\alpha}{\partial t} + \mathbf{v} \cdot \nabla f^\alpha + \mathbf{F} \cdot \nabla_{\mathbf{v}} f^\alpha = \left(\frac{\partial f^\alpha}{\partial t} \right)_{\text{coll}},$$

where \mathbf{F} is an external force field. The general form of the collision integral is $(\partial f^\alpha / \partial t)_{\text{coll}} = - \sum_\beta \nabla_{\mathbf{v}} \cdot \mathbf{J}^{\alpha/\beta}$, with

$$\begin{aligned} \mathbf{J}^{\alpha/\beta} = 2\pi\lambda_{\alpha\beta} \frac{e_\alpha^2 e_\beta^2}{m_\alpha} \int d^3v' (u^2 I - \mathbf{u}\mathbf{u}) u^{-3} \\ \cdot \left\{ \frac{1}{m_\beta} f^\alpha(\mathbf{v}) \nabla_{\mathbf{v}'} f^\beta(\mathbf{v}') - \frac{1}{m_\alpha} f^\beta(\mathbf{v}') \nabla_{\mathbf{v}} f^\alpha(\mathbf{v}) \right\} \end{aligned}$$

(Landau form) where $\mathbf{u} = \mathbf{v}' - \mathbf{v}$ and I is the unit dyad, or alternatively,

$$\mathbf{J}^{\alpha/\beta} = 4\pi\lambda_{\alpha\beta} \frac{e_\alpha^2 e_\beta^2}{m_\alpha^2} \left\{ f^\alpha(\mathbf{v}) \nabla_{\mathbf{v}} H(\mathbf{v}) - \frac{1}{2} \nabla_{\mathbf{v}} \cdot \left[f^\alpha(\mathbf{v}) \nabla_{\mathbf{v}} \nabla_{\mathbf{v}} H(\mathbf{v}) \right] \right\},$$

where the Rosenbluth potentials are

$$G(\mathbf{v}) = \int f^\beta(\mathbf{v}') u d^3v'$$

$$H(\mathbf{v}) = \left(1 + \frac{m_\alpha}{m_\beta} \right) \int f^\beta(\mathbf{v}') u^{-1} d^3v'.$$

```

\input prolog
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\pageno=36
If species  $\alpha$  is a weak beam (number and energy density small compared
with background) streaming through a Maxwellian plasma, then

$$\frac{dJ}{dt} = -\nu_s \nabla^2 v^2 \nabla f_\alpha + \frac{1}{2} (\nu_\perp \nabla^2 - \nu_\parallel \nabla) \cdot \nabla f_\alpha$$

\smallskip
{\headfont B-G-K Collision Operator}
\smallskip\indent
For distribution functions with no large gradients in velocity space, the
Fokker-Planck collision terms can be approximated according to

$$\frac{df_e}{dt} = \nu_{ee} (F_e - f_e) + \nu_{ei} (\bar{F}_i - f_i)$$


$$\frac{df_i}{dt} = \nu_{ie} (\bar{F}_i - f_i) + \nu_{ii} (F_i - f_i)$$

The respective slowing-down rates  $\nu_s$  given in the Relaxation Rate
section above can be used for  $\nu_{\alpha\beta}$ , assuming slow ions and fast
electrons, with  $\epsilon$  replaced by  $T_\alpha/T_e$ . For  $\nu_{ee}$  and
 $\nu_{ii}$ , one can equally well use  $\nu_\perp$  and the result is independent
to whether the slow- or fast-test particle limit is employed. The Maxwellian
 $F_\alpha$  and  $\bar{F}_\alpha$  are given by

$$F_\alpha = n_\alpha \left( \frac{m_e}{m_\alpha} \right)^{3/2} \exp \left( - \frac{m_e}{m_\alpha} \frac{v^2}{2T_\alpha} \right)$$


$$\bar{F}_\alpha = n_\alpha \left( \frac{m_e}{m_\alpha} \right)^{3/2} \exp \left( - \frac{m_e}{m_\alpha} \frac{v^2}{2T_\alpha} \right) \exp \left( - \frac{m_e}{m_\alpha} \frac{v^2}{2T_\alpha} \right)$$

where  $m_\alpha$ ,  $T_\alpha$  and  $n_\alpha$  are the species mass, temperature and
density, and  $\bar{F}_\alpha$  is the effective temperature obtained by taking
 $T_\alpha$  to be the effective temperature of the background plasma.
For  $\nu_{ei}$ ,  $\nu_{ie}$  and  $\nu_{ii}$ ,  $\bar{F}_i$  is the Maxwellian distribution
of the background plasma.
\smallskip
\indent
The Fokker-Planck collision terms can be approximated according to

$$\frac{df_e}{dt} = \nu_{ee} (F_e - f_e) + \nu_{ei} (\bar{F}_i - f_i)$$


$$\frac{df_i}{dt} = \nu_{ie} (\bar{F}_i - f_i) + \nu_{ii} (F_i - f_i)$$

where  $F_\alpha$  and  $\bar{F}_\alpha$  are the Maxwellian distribution functions of the
background plasma.

```

If species α is a weak beam (number and energy density small compared with background) streaming through a Maxwellian plasma, then

$$\mathbf{J}^{\alpha/3} = -\nu_s^{\alpha/3} \mathbf{v} f^\alpha - \frac{1}{2} \nu_\perp^{\alpha/3} v^2 \nabla_{\mathbf{v}} f^\alpha + \frac{1}{2} (\nu_\perp^{\alpha/3} - \nu_\parallel^{\alpha/3}) \mathbf{v} \mathbf{v} \cdot \nabla_{\mathbf{v}} f^\alpha.$$

B-G-K Collision Operator

For distribution functions with no large gradients in velocity space, the Fokker-Planck collision terms can be approximated according to

$$\frac{Df_i}{Dt} = \nu_{ei}(F_i - f_i) + \nu_{ii}(\bar{F}_i - f_i);$$

$$\frac{Df_e}{Dt} = \nu_{ei}(F_e - f_e) + \nu_{ee}(F_e - f_e).$$

The collisional slowing-down rates $\nu_s^{\alpha/3}$ given in the Relaxation Rate section can be used for ν_{ei} , assuming slow ions and fast electrons, with e replaced by T_e . (For ν_{ii} and ν_{ee} , one can equally well use ν_\perp , and the result is insensitive to whether the slow- or fast-test-particle limit is employed.) The Maxwellians F_i and F_e are given by

$$F_i = n_i \left(\frac{m_e}{2\pi k T_i} \right)^{3/2} \exp \left\{ - \left[\frac{m_e (\mathbf{v} - \mathbf{v}_i)^2}{2k T_i} \right] \right\};$$

$$F_e = n_e \left(\frac{m_e}{2\pi k T_e} \right)^{3/2} \exp \left\{ - \left[\frac{m_e (\mathbf{v} - \bar{\mathbf{v}}_e)^2}{2k T_e} \right] \right\},$$

where n_α and T_α are the number density, mean drift velocity, and effective temperature, obtained by taking moments of f_α . Some latitude in the definition of T_α and $\bar{\mathbf{v}}_\alpha$ is possible;²⁰ one choice is $T_i = T_i$, $T_e = T_e$, $\bar{\mathbf{v}}_i = \mathbf{v}_i$, $\bar{\mathbf{v}}_e = \mathbf{v}_e$.

Transport Coefficients

Transport equations for a multispecies plasma:

$$\frac{d^\alpha n_\alpha}{dt} + n_\alpha \nabla \cdot \mathbf{v}_\alpha = 0;$$

$$m_\alpha n_\alpha \frac{d^\alpha \mathbf{v}_\alpha}{dt} = -\nabla p_\alpha - \nabla \cdot \mathbf{P}_\alpha + Z_\alpha e n_\alpha \left[\mathbf{E} + \frac{1}{c} \mathbf{v}_\alpha \times \mathbf{B} \right] + \mathbf{R}_\alpha;$$

$$\frac{3}{2}n_\alpha \frac{d^\alpha kT_\alpha}{dt} + p_\alpha \nabla \cdot \mathbf{v}_\alpha = -\nabla \cdot \mathbf{q}_\alpha - P_\alpha : \nabla \mathbf{v}_\alpha + Q_\alpha.$$

Here $d^\alpha/dt \equiv \partial/\partial t + \mathbf{v}_\alpha \cdot \nabla$; $p_\alpha = n_\alpha kT_\alpha$, where k is Boltzmann's constant; $\mathbf{R}_\alpha = \sum_\beta \mathbf{R}_{\alpha\beta}$ and $Q_\alpha = \sum_\beta Q_{\alpha\beta}$, where $\mathbf{R}_{\alpha\beta}$ and $Q_{\alpha\beta}$ are respectively the momentum and energy gained by the α th species through collisions with the β th; P_α is the stress tensor; and \mathbf{q}_α is the heat flow.

The transport coefficients in a simple two-component plasma (electrons and singly charged ions) are tabulated below. Here \parallel and \perp refer to the direction of the magnetic field $\mathbf{B} = b\mathbf{B}$; $\mathbf{u} = \mathbf{v}_e - \mathbf{v}_i$ is the relative streaming velocity; $n_e = n_i \equiv n$; $\mathbf{j} = -ne\mathbf{u}$ is the current; $\omega_{ce} = 1.76 \times 10^7 B \text{ sec}^{-1}$ and $\omega_{ci} = (m_e/m_i)\omega_{ce}$ are the electron and ion gyrofrequencies, respectively; and the basic collisional times are taken to be

$$\tau_e = \frac{3\sqrt{m_e}(kT_e)^{3/2}}{4\sqrt{2\pi}n\lambda e^4} = 3.44 \times 10^5 \frac{T_e^{3/2}}{n\lambda} \text{ sec},$$

where λ is the Coulomb logarithm, and

$$\tau_i = \frac{3\sqrt{m_i}(kT_i)^{3/2}}{4\sqrt{\pi}n\lambda e^4} = 2.09 \times 10^7 \frac{T_i^{3/2}}{n\lambda} \mu^{1/2} \text{ sec}.$$

In the limit of large fields ($\omega_{ce}\tau_e \gg 1$, $\alpha = i, e$) the transport processes may be summarized as follows:²¹

momentum transfer	$\mathbf{R}_{ei} = -\mathbf{R}_{ie} \equiv \mathbf{R} = \mathbf{R}_u + \mathbf{R}_T$
frictional force	$\mathbf{R}_u = ne(\mathbf{j}_\parallel/\sigma_\parallel + \mathbf{j}_\perp/\sigma_\perp)$
electrical conductivities	$\sigma_\parallel = 2.0\sigma_\perp = 2.0 \frac{ne^2\tau_e}{m_e}$
thermal force	$\mathbf{R}_T = -0.71n\nabla_\parallel(kT_e) - \frac{3n}{2\omega_{ce}\tau_e}\mathbf{b} \times \nabla_\perp(kT_e)$
ion heating	$Q_i = \frac{3m_e}{m_i} \frac{nk}{\tau_e}(T_e - T_i)$
electron heating	$Q_e = -Q_i - \mathbf{R} \cdot \mathbf{u}$
ion heat flux	$\mathbf{q}_i = -\kappa'_\parallel \nabla_\parallel(kT_i) - \kappa'_\perp \nabla_\perp(kT_i) + \kappa'_\Lambda \mathbf{b} \times \nabla_\perp(kT_i)$
ion thermal conductivities	$\kappa'_\parallel = 3.9 \frac{nkT_i\tau_i}{m_i}$; $\kappa'_\perp = \frac{2nkT_i}{m_i\omega_{ci}^2\tau_i}$; $\kappa'_\Lambda = \frac{5nkT_i}{2m_i\omega_{ci}}$
electron heat flux	$\mathbf{q}_e = \mathbf{q}'_u + \mathbf{q}'_T$
frictional heat flux	$\mathbf{q}'_u = 0.71nkT_e\mathbf{u}_\parallel + \frac{3nkT_e}{2\omega_{ce}\tau_e}\mathbf{b} \times \mathbf{u}_\perp$

thermal gradient heat flux	$\mathbf{q}'_T = -\kappa'_{\parallel} \nabla_{\parallel}(kT_e) - \kappa'_{\perp} \nabla_{\perp}(kT_e) - \kappa'_{\wedge} \mathbf{b} \times \nabla_{\perp}(kT_e);$
electron thermal conductivities	$\kappa'_{\parallel} = 3.2 \frac{nkT_e \tau_e}{m_e}; \quad \kappa'_{\perp} = 4.7 \frac{nkT_e}{m_e \omega_{ce}^2 \tau_e}; \quad \kappa'_{\wedge} = \frac{5nkT_e}{2m_e \omega_{ce}^2};$
stress tensor (both species)	$P_{xx} = -\frac{\eta_0}{2}(W_{xx} + W_{yy}) - \frac{\eta_1}{2}(W_{xx} - W_{yy}) + \eta_3 W_{zz};$ $P_{yy} = -\frac{\eta_0}{2}(W_{xx} + W_{yy}) + \frac{\eta_1}{2}(W_{xx} - W_{yy}) + \eta_3 W_{zz};$ $P_{xy} = P_{yx} = -\eta_3 W_{xy} + \frac{\eta_3}{2}(W_{xx} - W_{yy});$ $P_{xz} = P_{zx} = \eta_2 W_{xz} - \eta_4 W_{yz};$ $P_{yz} = P_{zy} = \eta_2 W_{yz} + \eta_4 W_{xz};$ $P_{zz} = \eta_0 W_{zz};$

where the z axis is defined parallel to \mathbf{B}_0 :

ion viscosity	$\eta_0 = 0.96nkT_i \tau_i; \quad \eta'_1 = \frac{3nkT_i}{10\omega_{ci}^2 \tau_i}; \quad \eta'_2 = \frac{6nkT_i}{5\omega_{ci}^2 \tau_i};$ $\eta_3 = \frac{nkT_i}{2\omega_{ci}^2}; \quad \eta_4 = \frac{nkT_i}{\omega_{ci}^2};$
electron viscosity	$\eta'_0 = 0.73nkT_e \tau_e; \quad \eta'_1 = 0.51 \frac{nkT_e}{\omega_{ce}^2 \tau_e}; \quad \eta'_2 = 2.0 \frac{nkT_e}{\omega_{ce}^2 \tau_e};$ $\eta'_3 = -\frac{nkT_e}{2\omega_{ce}^2}; \quad \eta'_4 = -\frac{nkT_e}{\omega_{ce}^2};$

For both species the rate-of-strain tensor is defined as

$$W_{jk} = \frac{\partial v_j}{\partial x_k} + \frac{\partial v_k}{\partial x_j} - \frac{2}{3} \delta_{jk} \nabla \cdot \mathbf{v}.$$

When $\mathbf{B} = 0$ the following simplifications occur:

$$\mathbf{R}_0 = ne\mathbf{j}_T/\sigma_{\parallel}; \quad \mathbf{R}_T = -0.71n\nabla(kT_e); \quad \mathbf{q}_T = -\kappa'_{\parallel} \nabla(kT_e);$$

$$\mathbf{q}'_0 = 0.71nkT_e \mathbf{u}; \quad \mathbf{q}'_T = -\kappa'_{\parallel} \nabla(kT_e); \quad P_{jk} = -\eta_0 W_{jk},$$

for $\omega_{ce} \tau_e \gg 1 \gg \omega_{ci} \tau_i$, the electrons obey the high-field expressions and the ions obey the zero field expressions. Collisional transport theory is

applicable when (1) macroscopic time rates of change satisfy $d/dt \ll 1/\tau$, where τ is the longest collisional time scale, and (in the absence of a magnetic field) (2) macroscopic length scales L satisfy $L \gg l$, where $l = v\tau$ is the mean free path. In a strong field, $\omega_{ce} \tau \gg 1$, condition (2) is replaced by $L_{\parallel} \gg l$ and $L_{\perp} \gg \sqrt{Tr_e}$ ($L_{\perp} \gg r_e$ in a uniform field).

where L_{\parallel} is a macroscopic scale parallel to the field \mathbf{B} and L_{\perp} is the smaller of $B/|\nabla_{\perp} B|$ and the transverse plasma dimension. In addition, the standard transport coefficients are valid only when (3) the Coulomb logarithm satisfies $\lambda \gg 1$; (4) the electron gyroradius satisfies $r_e \gg \lambda_D$, or $8\pi n_e m_e c^2 \gg B^2$; (5) relative drifts $\mathbf{u} = \mathbf{v}_{\alpha} - \mathbf{v}_{\beta}$ between two species are small compared with the thermal velocities, i.e., $u^2 \ll kT_{\alpha}/m_{\alpha}, kT_{\beta}/m_{\beta}$; and (6) anomalous transport processes owing to microinstabilities are negligible.

Weakly Ionized Plasmas

Collision frequency for scattering of charged particles of species α by neutrals is

$$\nu_{\alpha} = n_0 \sigma_s^{\alpha/0} (kT_{\alpha}/m_{\alpha})^{1/2},$$

where n_0 is the neutral density and $\sigma_s^{\alpha/0}$ is the cross section, typically $\sim 5 \times 10^{-15} \text{ cm}^2$ and weakly dependent on temperature.

When the system is small compared with a Debye length, $L \ll \lambda_D$, the charged particle diffusion coefficients are

$$D_{\alpha} = kT_{\alpha}/m_{\alpha} \nu_{\alpha}.$$

In the opposite limit, both species diffuse at the ambipolar rate

$$D_A = \frac{\mu_i D_e - \mu_e D_i}{\mu_i - \mu_e} = \frac{(T_i + T_e) D_i D_e}{T_i D_e + T_e D_i},$$

where $\mu_i \approx e_i/m_i \nu_i$ is the mobility. The conductivity σ_{α} satisfies $\sigma_{\alpha} = e_i \mu_i n_i$.

In the presence of a magnetic field \mathbf{B} the scalars μ and σ become tensors,

$$\mathbf{J}^{\alpha} = \boldsymbol{\sigma}^{\alpha} \cdot \mathbf{E} = \sigma_{\parallel}^{\alpha} \mathbf{E}_{\parallel} + \sigma_{\perp}^{\alpha} \mathbf{E}_{\perp} + \sigma_{\Lambda}^{\alpha} \mathbf{E} \times \mathbf{b},$$

where $\mathbf{b} = \mathbf{B}/B$ and

$$\sigma_{\parallel}^{\alpha} = n_{\alpha} e_{\alpha}^2 / m_{\alpha} \nu_{\alpha};$$

$$\sigma_{\perp}^{\alpha} = \sigma_{\parallel}^{\alpha} \nu_{\alpha}^2 / (\nu_{\alpha}^2 + \omega_{c\alpha}^2);$$

$$\sigma_{\Lambda}^{\alpha} = \sigma_{\parallel}^{\alpha} \nu_{\alpha} \omega_{c\alpha} / (\nu_{\alpha}^2 + \omega_{c\alpha}^2).$$

Here σ_{\perp} and σ_{Λ} are the Pedersen and Hall conductivities, respectively.

```

\input prelog
\hoffset=1.125truein
\voffset=0.75truein
\hsize=6.25truein
\vsizer=9.0truein
\pageno=40
\centerline{\headfont APPROXIMATE MAGNITUDES}
\vskip1pt
\centerline{\headfont IN SOME TYPICAL PLASMAS}
\bsk % BEGINNING OF TABLE.
$$\vbox{\offinterlineskip \def\trule{\noalign{\hrule}}
\tabskip=0pt \halign to \hsize{
\vrule# \tabskip=3pt minus2pt\strut #\hfil &\vrule #&\hfil$#\hfil
&\vrule #&\hfil$#\hfil &\vrule #&\hfil$#\hfil
&\vrule #&\hfil$#\hfil &\vrule #&\hfil$#\hfil
&\vrule #&\hfil$#\hfil &\vrule #\tabskip=0pt\cr\trule
\tskc{7}{4pt}
&\hfil Plasma Type(\n\ {\rm cm})^{-3}[T\ {\rm eV}][\omega_{pe}\ {\rm sec}]^{-1}
[\lambda_{D}\ {\rm cm}][\lambda_{D}^{-3}/v_{ei}\ {\rm sec}]^{-1}&\cr
\tskc{7}{4pt} \trule \tskc{7}{1pt} \trule \tskc{7}{4pt}

&Interstellar gas |1 |1 |6\times10^{-4}\ph{0}| |\omega$7\times10^{-2}$\hfil |
4\times10^{-8}|7\times10^{-5}|&\cr \tskc{7}{3pt}
&Gaseous nebula |10^{-3} |1 |2\times10^{-6}\ph{0}| |20 |10^{-7} |
0\times10^{-2}&\cr \tskc{7}{3pt}
&Solar Corona |10^{-6} |10^{-2} |6\times10^{-7}\ph{0}| |7 |4\times10^{-8} |
0\times10^{-2}&\cr \tskc{7}{3pt}
&Diffuse hot plasma |10^{-12} |10^{-2} |6\times10^{-10}| |7\times10^{-3} |
4\times10^{-5}|40 &\cr \tskc{7}{3pt}
&Solar atmosphere, |10^{-14} |1 |6\times10^{-11}| |7\times10^{-5}| |40 |
2\times10^{-9}\ph{-}&\cr
&\hskip 1em gas discharge |11111&\cr \tskc{7}{3pt}
&Warm plasma |10^{-14} |10^{-6} |6\times10^{-11}| |2\times10^{-4}| |10^{-3} |10^{-7}&\cr
\tskc{7}{3pt}
&Hot plasma |10^{-14} |10^{-2} |6\times10^{-11}| |7\times10^{-4}|
4\times10^{-4}|4\times10^{-6}\ph{-}&\cr \tskc{7}{3pt}
&Other nuclear |10^{-18} |10^{-4} |2\times10^{-12}| |2\times10^{-3}| |10^{-7} |
1\times10^{-4}\ph{-}&\cr
&\hskip 1em plasma |11111&\cr \tskc{7}{3pt}
&Theta pinch |10^{-16} |10^{-2} |6\times10^{-12}| |7\times10^{-5}| |
4\times10^{-3}|2\times10^{-9}\ph{-}&\cr \tskc{7}{3pt}
&Pence hot plasma |10^{-16} |10^{-2} |6\times10^{-12}| |7\times10^{-6}| |
4\times10^{-2}|2\times10^{-10}\ph{.}&\cr \tskc{7}{3pt}
&Laser Plasma |10^{-20} |10^{-2} |6\times10^{-14}| |7\times10^{-7}| |40 |
2\times10^{-12}\ph{.}&\cr \tskc{7}{4pt}\trule}}$$
\bsk % END OF TABLE.
The diagram (facing) gives comparable information in graphical form.$^{-2}$
\fill\ject\end

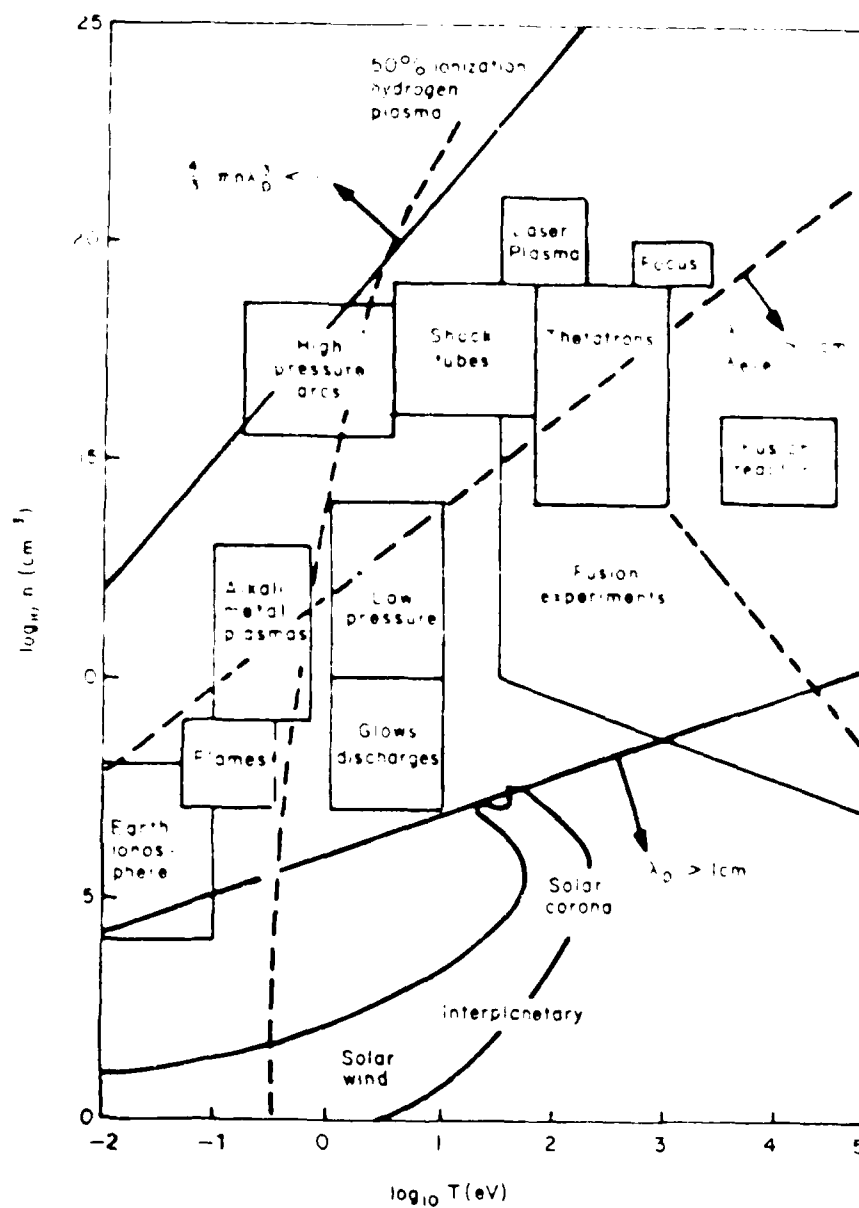
```

APPROXIMATE MAGNITUDES IN SOME TYPICAL PLASMAS

Plasma Type	$n \text{ cm}^{-3}$	$T \text{ eV}$	$\omega_{pe} \text{ sec}^{-1}$	$\lambda_D \text{ cm}$	$n\lambda_D^3$	$v_{ei} \text{ sec}^{-1}$
Interstellar gas	1	1	6×10^4	7×10^2	4×10^8	7×10^{-5}
Gaseous nebula	10^3	1	2×10^6	20	10^7	6×10^{-2}
Solar Corona	10^6	10^2	6×10^7	7	4×10^8	6×10^{-2}
Diffuse hot plasma	10^{12}	10^2	6×10^{10}	7×10^{-3}	4×10^5	40
Solar atmosphere, gas discharge	10^{14}	1	6×10^{11}	7×10^{-5}	40	2×10^9
Warm plasma	10^{14}	10	6×10^{11}	2×10^{-4}	10^3	10^7
Hot plasma	10^{14}	10^2	6×10^{11}	7×10^{-4}	4×10^4	4×10^6
Thermonuclear plasma	10^{15}	10^4	2×10^{12}	2×10^{-3}	10^7	5×10^4
Theta pinch	10^{16}	10^2	6×10^{12}	7×10^{-5}	4×10^3	3×10^8
Dense hot plasma	10^{18}	10^2	6×10^{13}	7×10^{-6}	4×10^2	2×10^{10}
Laser Plasma	10^{20}	10^2	6×10^{14}	7×10^{-7}	40	2×10^{12}

The diagram (facing) gives comparable information in graphical form.²²

The next page is an inserted figure



AD-A185 928

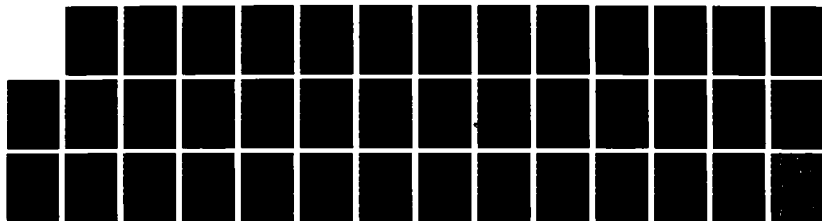
TEXTING THE FORMULARY: A COLLECTION OF EXAMPLES OF THE
USE OF TEX TO PROD. (U) NAVAL RESEARCH LAB WASHINGTON
DC D L BOOK ET AL. 86 OCT 87 NRL-RR-6844

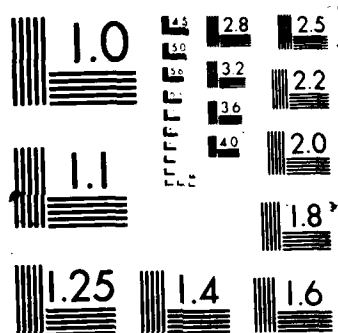
2/2

UNCLASSIFIED

F/G 20/9

NL





IONOSPHERIC PARAMETERS²³

The following tables give average nighttime values. Where two numbers are entered, the first refers to the lower and the second to the upper portion of the layer.

Quantity	E Region	F Region
Altitude (km)	90-160	160-500
Number density (m^{-3})	1.5×10^{10} - 3.0×10^{10}	5×10^{10} - 2×10^{11}
Height-integrated number density (m^{-2})	9×10^{14}	4.5×10^{15}
Ion-neutral collision frequency (sec^{-1})	2×10^3 - 10^2	0.5-0.05
Ion gyro-/collision frequency ratio κ_i	0.09-2.0	4.6×10^2 - 5.0×10^3
Ion Pederson factor $\kappa_i/(1+\kappa_i^2)$	0.09-0.5	2.2×10^{-3} - 2×10^{-4}
Ion Hall factor $\kappa_i^2/(1+\kappa_i^2)$	8×10^{-4} -0.8	1.0
Electron-neutral collision frequency	1.5×10^4 - 9.0×10^2	80-10
Electron gyro-/collision frequency ratio κ_e	4.1×10^2 - 6.9×10^3	7.8×10^4 - 6.2×10^5
Electron Pedersen factor $\kappa_e/(1+\kappa_e^2)$	2.7×10^{-3} - 1.5×10^{-4}	10^{-5} - 1.5×10^{-6}
Electron Hall factor $\kappa_e^2/(1+\kappa_e^2)$	1.0	1.0
Mean molecular weight	28-26	22-16
Ion gyrofrequency (sec^{-1})	180-190	230-300
Neutral diffusion coefficient ($\text{m}^2 \text{sec}^{-1}$)	30 - 5×10^3	10^5

The terrestrial magnetic field in the lower ionosphere at equatorial latitudes is approximately $B_0 = 0.35 \times 10^{-4}$ tesla. The earth's radius is $R_E = 6371$ km.

```

\input prolog \pageno=43
\hoffset=1.0truein\voffset=1.0truein\hsize=6.5truein\vsize=9.0truein
\centerline{{\headfont SOLAR PHYSICS PARAMETERS}$^{24}$} \msk
\vbox{\offinterlineskip \tabskip=0pt \halign to\hsize{
\vrule#\tabskip=3pt minus2pt&\strut#\hfil&\vrule#&\hfil$#\hfil&\vrule#
&$#\hfil&\vrule#&\hfil#\hfil&\vrule#\tabskip=0pt\cr\trule \tskc{4}{2pt}
&\hfil Parameter|\hbox{Symbol}|\hfil\hbox{Value}|\hbox{Units}&\cr
\tskc{4}{2pt} \trule \tskc{4}{1pt} \trule \tskc{4}{2pt}
&Total mass|M_\odot|1.99\times 10^{33}|g&\cr \tskc{4}{1.25pt}
&Radius|R_\odot|6.96\times 10^{10}|cm&\cr \tskc{4}{1.25pt}
&Surface gravity|g_\odot|2.74\times 10^4|cm\ts s^{-2}$&\cr \tskc{4}{1.25pt}
&Escape speed|v_\infty|6.18\times 10^7|cm\ts s^{-1}$&\cr \tskc{4}{1.25pt}
&Upward mass flux in spicules|\hbox{---}|1.6\times 10^{-9}|g\ts cm^{-2}$
\ts s^{-1}$&\cr \tskc{4}{1.25pt}
&Vertically integrated atmospheric density|\hbox{---}|\om\hfil4.28\hfil|
g\ts cm^{-2}$&\cr \tskc{4}{1.25pt}
&Sunspot magnetic field strength|E_{\rm max}|\om\hfil2500\hbox{--}3500\hfil|
G&\cr \tskc{4}{1.25pt}
&Surface temperature|T_0|\om\hfil6420\hfil|K&\cr \tskc{4}{1.25pt}
&Radiant power|\cal L_\odot|3.90\times 10^{33}|erg\ts s^{-1}$&\cr \tskc{4}{1.25pt}
&Radiant flux density|\cal F|6.41\times 10^{10}|erg\ts cm^{-2}$s^{-1}$&\cr
\tskc{4}{1.25pt}
&Optical depth at 500\ts nm, measured|\tau_{500}|\om\hfil0.99\hfil|\hbox{---}&\cr
&\hskip 1em from photosphere||&\cr \tskc{4}{1.25pt}
&Astronomical unit (radius of earth's orbit)|\rm AU|1.50\times 10^{13}|cm&\cr
\tskc{4}{1.25pt}
&Solar constant (intensity at 1\ts AU)|f|1.39\times 10^6|
erg\ts cm^{-2}$\ts s^{-1}$&\cr \tskc{4}{2pt}\trule}} \msk
{\headfont Chromosphere and Corona}$^{25}$} \ssk
\vbox{\offinterlineskip \tabskip=0pt \halign to\hsize{
\vrule#\tabskip=3pt minus2pt&\strut#\hfil&\vrule#&\hfil$#\hfil&\vrule#
&\hfil$#\hfil&\vrule#&\hfil$#\hfil&\vrule#\tabskip=0pt\cr\trule \tskc{4}{2pt}
&\om|\hbox{Quiet}|\hbox{Coronal}|\hbox{Active}&\cr \bs{4pt}
&\hfil Parameter (Units)|\om|\om|\om&\cr \bs{4pt}
&\om|\hbox{Sun}|\hbox{Hole}|\hbox{Region}&\cr
\tskc{4}{2pt} \trule \tskc{4}{1pt} \trule \tskc{4}{2pt}
&Chromospheric radiation losses||&\cr
&\hskip 1em (erg\ts cm^{-2}$\ts s^{-1}$)||&\cr \tskc{4}{1pt}
&\hskip 2em Low chromosphere|2\times 10^6|2\times 10^6|\approx 10^7&\cr \tskc{4}{1pt}
&\hskip 2em Middle chromosphere|2\times 10^6|2\times 10^6|10^7&\cr \tskc{4}{1pt}
&\hskip 2em Upper chromosphere|3\times 10^5|3\times 10^5|2\times 10^6&\cr \tskc{4}{1pt}
&\hskip 2em Total|4\times 10^6|4\times 10^6|\approx 2\times 10^7&\cr \tskc{4}{1.5pt}
&Transition layer pressure (dyne\ts cm^{-2}$)|0.2|0.07|2&\cr \tskc{4}{1.5pt}
&Coronal temperature (K) at 1.1\ts R_\odot$|1.1\hbox{--}1.6\times 10^6|10^6|
2.5\times 10^6&\cr \tskc{4}{1.5pt}
&Coronal energy losses (erg\ts cm^{-2}$\ts s^{-1}$)||&\cr \tskc{4}{1pt}
&\hskip 2em Conduction|2\times 10^5|6\times 10^4|10^5\hbox{--}10^7&\cr \tskc{4}{1pt}
&\hskip 2em Radiation|10^5|10^4|5\times 10^6&\cr \tskc{4}{1pt}
&\hskip 2em Solar Wind|\approx 10^5|7\times 10^5|<10^5&\cr \tskc{4}{1pt}
&\hskip 2em Total|3\times 10^5|8\times 10^5|10^7&\cr \tskc{4}{1.5pt}
&Solar wind mass loss (g\ts cm^{-2}$\ts s^{-1}$)|\approx 10^2|2\times 10^{-11}|
2\times 10^{-10}|<4\times 10^{-11}&\cr \tskc{4}{2pt}\trule}}
\vfill\ject\end

```

SOLAR PHYSICS PARAMETERS²⁴

Parameter	Symbol	Value	Units
Total mass	M_{\odot}	1.99×10^{33}	g
Radius	R_{\odot}	6.96×10^{10}	cm
Surface gravity	g_{\odot}	2.74×10^4	cm s^{-2}
Escape speed	v_{∞}	6.18×10^7	cm s^{-1}
Upward mass flux in spicules	—	1.6×10^{-9}	$\text{g cm}^{-2} \text{ s}^{-1}$
Vertically integrated atmospheric density	—	4.28	g cm^{-2}
Sunspot magnetic field strength	B_{max}	2500-3500	G
Surface temperature	T_0	6420	K
Radiant power	\mathcal{L}_{\odot}	3.90×10^{33}	erg s^{-1}
Radiant flux density	\mathcal{F}	6.41×10^{10}	$\text{erg cm}^{-2} \text{ s}^{-1}$
Optical depth at 500nm, measured from photosphere	τ_{500}	0.99	
Astronomical unit (radius of earth's orbit)	AU	1.50×10^{13}	cm
Solar constant (intensity at 1 AU)	f	1.39×10^6	$\text{erg cm}^{-2} \text{ s}^{-1}$

Chromosphere and Corona²⁵

Parameter (Units)	Quiet Sun	Coronal Hole	Active Region
Chromospheric radiation losses ($\text{erg cm}^{-2} \text{ s}^{-1}$)			
Low chromosphere	2×10^6	2×10^6	$\gtrsim 10^7$
Middle chromosphere	2×10^6	2×10^6	10^7
Upper chromosphere	3×10^5	3×10^5	2×10^6
Total	4×10^6	4×10^6	$\gtrsim 2 \times 10^7$
Transition layer pressure (dyne cm^{-2})	0.2	0.07	2
Coronal temperature (K) at 1.1 R .	1.1 1.6×10^6	10^6	2.5×10^6
Coronal energy losses ($\text{erg cm}^{-2} \text{ s}^{-1}$)			
Conduction	2×10^5	6×10^4	10^5 10^7
Radiation	10^5	10^4	5×10^6
Solar Wind	$\lesssim 5 \times 10^4$	7×10^5	$< 10^5$
Total	3×10^5	8×10^5	10^7
Solar wind mass loss ($\text{g cm}^{-2} \text{ s}^{-1}$)	$\lesssim 2 \times 10^{-11}$	2×10^{-10}	$< 4 \times 10^{-11}$

```

\input prolog
\hoffset=1.25truein
\voffset=1.0truein
\hsize=6.0truein
\vsize=9.0truein
\pageno=44
\centerline{{\headfont THERMONUCLEAR FUSION}\$^{26}\$}
\bigskip
Natural abundance of isotopes:
\smallskip\nointerlineskip
\$\vbox{{\halign{{#\quad&##\$}\hfil}\cr
hydrogen&n_D^{\phantom{3}}/n_H^{\phantom{3}}=1.5\times10^{-4}\cr
helium&n_{\rm He^3}/n_{\rm He^4}=1.3\times10^{-6}\cr
\ncolalign{\vskip1.5pt}
lithium&n_{\rm Li^6}/n_{\rm Li^7}=0.08\cr}}\$
\halign{{#\quad\quad&##\$}\hfil&##\$}\hfil\cr
Mass ratios: &m_e/m_D &=2.72\times10^{-4} = 1/3670 \cr
&(m_e/m_D)^{1/2} &=1.65\times10^{-2} = 1/60.6 \cr
&m_e/m_T &=1.82\times10^{-4} = 1/5496 \cr
&(m_e/m_T)^{1/2} &=1.35\times10^{-2} = 1/74.1 \cr
\medskip
Absorbed radiation dose is measured in rads: 1 rad = 10\$^{-2}\$ \ts erg\ts g\$^{-1}\$.
The curie (abbreviated Ci) is a measure of radioactivity: 1curie =
\$3.7\times10^{10}\$ \ts counts\ts sec\$^{-1}\$.
\medskip
Fusion reactions (branching ratios are correct for energies near the cross
section peaks; a negative yield means the reaction is endothermic):\$^{27}\$
\smallskip{\def\yields#1{\lower4.5pt\hbox{{\$overrightarrow{\ph{0}}#1\ph{0}}\$}}
% THE \yields MACRO IS USED TO DRAW THE 'YIELDS' SYMBOL '--->' WITH
% A NUMBER UNDER IT.
\halign{{\indent#\quad&#\quad&}\hfil&#\hfil&#\hfil\cr
&(1a)&D + D\yields{50\%}&T(1.01\ts MeV) + p(3.02\ts MeV)\cr
&(1b)&&\yields{50\%}&He^3(0.82\ts MeV) + n(2.45\ts MeV)\cr
&(2) &D + T\yields{\ph{50\%}}&He^4(3.5\ts MeV) + n(14.1\ts MeV)\cr
&(3) &D + He^3\yields{\ph{50\%}}&He^4(3.6\ts MeV) + p(14.7\ts MeV)\cr
&(4) &T + T\yields{\ph{50\%}}&He^4\$ + 2n + 11.3\ts MeV\cr
&(5a)&He^3\$ + T\yields{51\%}&He^4\$ + p + n + 12.1\ts MeV\cr
&(5b)&&\yields{43\%}&He^4(4.8\ts MeV) + D(9.5\ts MeV)\cr
&(5c)&&\yields{\ph{56\%}}&He^5(2.4\ts MeV) + p(11.9\ts MeV)\cr
&(6) &p + Li^6\yields{\ph{50\%}}&He^4(1.7\ts MeV) + He^3(2.3\ts MeV)\cr
&(7a)&p + Li^7\yields{20\%}&2 He^4\$ + 17.3\ts MeV\cr
&(7b)&&\yields{80\%}&Be^7\$ + n \$- \$ 1.6\ts MeV\cr
&(8) &D + Li^6\yields{\ph{50\%}}&2He^4\$ + 22.4\ts MeV\cr
&(9) &p + B^{11}\yields{\ph{50\%}}&3 He^4\$ + 8.7\ts MeV\cr
&(10)&n + Li^6\yields{\ph{50\%}}&He^4(2.1\ts MeV) + T(2.7\ts MeV)\cr
\medskip
The total cross section in barns as a function of \$E\$, the energy in keV
of the incident particle (the first ion on the left side of Eqs. (1)-(10)),
assuming the target ion at rest, can be fitted by\$^{28}\$
\mathrel{\mathop{\rule{0pt}{1.2ex}}}{\sigma}_T(E)=[A_0+1/\{[(A_4-A_2E)^2+1]\}]\{-1\}A_2\exp[-E/(A_1+
E/(1/2)^{-1})]\right]\}!$
\end{pre>

```


THERMONUCLEAR FUSION²⁶

Natural abundance of isotopes:

hydrogen	$n_D/n_H = 1.5 \times 10^{-4}$
helium	$n_{\text{He}^3}/n_{\text{He}^4} = 1.3 \times 10^{-6}$
lithium	$n_{\text{Li}^6}/n_{\text{Li}^7} = 0.08$

Mass ratios:	$m_e/m_D = 2.72 \times 10^{-4} = 1/3670$
	$(m_e/m_D)^{1/2} = 1.65 \times 10^{-2} = 1/60.6$
	$m_e/m_T = 1.82 \times 10^{-4} = 1/5496$
	$(m_e/m_T)^{1/2} = 1.35 \times 10^{-2} = 1/74.1$

Absorbed radiation dose is measured in rads: 1 rad = 10^2 erg g^{-1} . The curie (abbreviated Ci) is a measure of radioactivity: 1 curie = $3.7 \times 10^{10} \text{ counts sec}^{-1}$.

Fusion reactions (branching ratios are correct for energies near the cross section peaks; a negative yield means the reaction is endothermic):²⁷

- (1a) $\text{D} + \text{D} \xrightarrow{50\%} \text{T}(1.01 \text{ MeV}) + \text{p}(3.02 \text{ MeV})$
- (1b) $\xrightarrow{50\%} \text{He}^3(0.82 \text{ MeV}) + \text{n}(2.45 \text{ MeV})$
- (2) $\text{D} + \text{T} \longrightarrow \text{He}^4(3.5 \text{ MeV}) + \text{n}(14.1 \text{ MeV})$
- (3) $\text{D} + \text{He}^3 \longrightarrow \text{He}^4(3.6 \text{ MeV}) + \text{p}(14.7 \text{ MeV})$
- (4) $\text{T} + \text{T} \longrightarrow \text{He}^4 + 2\text{n} + 11.3 \text{ MeV}$
- (5a) $\text{He}^3 + \text{T} \xrightarrow{51\%} \text{He}^4 + \text{p} + \text{n} + 12.1 \text{ MeV}$
- (5b) $\xrightarrow{43\%} \text{He}^4(4.8 \text{ MeV}) + \text{D}(9.5 \text{ MeV})$
- (5c) $\xrightarrow{6\%} \text{He}^5(2.4 \text{ MeV}) + \text{p}(11.9 \text{ MeV})$
- (6) $\text{p} + \text{Li}^6 \longrightarrow \text{He}^4(1.7 \text{ MeV}) + \text{He}^3(2.3 \text{ MeV})$
- (7a) $\text{p} + \text{Li}^7 \xrightarrow{20\%} 2 \text{ He}^4 + 17.3 \text{ MeV}$
- (7b) $\xrightarrow{80\%} \text{Be}^7 + \text{n} - 1.6 \text{ MeV}$
- (8) $\text{D} + \text{Li}^6 \longrightarrow 2\text{He}^4 + 22.4 \text{ MeV}$
- (9) $\text{p} + \text{B}^{11} \longrightarrow 3 \text{ He}^4 + 8.7 \text{ MeV}$
- (10) $\text{n} + \text{Li}^6 \longrightarrow \text{He}^4(2.1 \text{ MeV}) + \text{T}(2.7 \text{ MeV})$

The total cross section in barns as a function of E , the energy in keV of the incident particle [the first ion on the left side of Eqs. (1) (5)], assuming the target ion at rest, can be fitted by²⁸

$$\sigma_T(E) = \frac{A_5 + [(A_4 - A_3 E)^2 + 1]^{-1} A_2}{E [\exp(A_1 E^{-1/2}) - 1]}$$

The listing for page 45 begins on the next page.

```

\input prolog
\hsize=6.5truein
\hoffset=1.125truein
\pageno=45
\nointerlineskip
where the Duane coefficients  $A_j$  for the principle fusion reactions are
as follows:
\ssk % BEGINNING OF FIRST TABLE.

$$\begin{array}{c} \text{Table 1: Duane coefficients } A_j \text{ for principle fusion reactions} \\ \text{Reactions: } D-D, D-T, D-He^3, T-T, T-He^3, a+b \end{array}$$

\ssk

```

Reaction rates $\overline{\sigma v}$ (in cm^3/sec), averaged over Maxwellian distributions:

```

\ssk % BEGINNING OF SECOND TABLE.

$$\begin{array}{c} \text{Table 2: Reaction rates } \overline{\sigma v} \text{ (in cm}^3/\text{sec)} \\ \text{Reactions: } D-D, D-T, D-He^3, T-T, T-He^3, a+b \end{array}$$

\ssk

```

$$(\overline{\sigma v})_{DD} = 2.33 \times 10^{-14} T^{-2/3} \exp(-18.76 T^{-1/2})$$

$$\text{cm}^3 \text{sec}^{-1};$$

$$(\overline{\sigma v})_{DT} = 3.68 \times 10^{-12} T^{-2/3} \exp(-19.94 T^{-1/2})$$

$$\text{cm}^3 \text{sec}^{-1},$$

where T is measured in keV.

medskip

The power density released in the form of charged particles is

smallskip

$$P_{DD} = 3.3 \times 10^{-13} n_D^2 (\overline{\sigma v})_{DD} \text{ watts cm}^{-3}$$

(including the subsequent

4.625 true in D-T reaction);

$$P_{DT} = 5.6 \times 10^{-13} n_D n_T (\overline{\sigma v})_{DT} \text{ watts cm}^{-3};$$

medskip

$$P_{DHe} = 2.9 \times 10^{-12} n_D \langle \sigma v \rangle_{DHe} (\overline{\sigma v})_{DHe} \text{ watts cm}^{-3}.$$

will reject end

where the Duane coefficients A_j for the principle fusion reactions are as follows:

	D-D (1a)	D-D (1b)	D-T (2)	D-He ³ (3)	T-T (4)	T-He ³ (5a-c)
A_1	46.097	47.88	45.95	89.27	38.39	123.1
A_2	372	482	50200	25900	448	11250
A_3	4.36×10^{-4}	3.08×10^{-4}	1.368×10^{-2}	3.98×10^{-3}	1.02×10^{-3}	0
A_4	1.220	1.177	1.076	1.297	2.09	0
A_5	0	0	409	647	0	0

Reaction rates $\overline{\sigma v}$ (in $\text{cm}^3 \text{sec}^{-1}$), averaged over Maxwellian distributions:

Temperature (keV)	D-D (1a + 1b)	D-T (2)	D-He ³ (3)	T-T (4)	T-He ³ (5a-c)
1.0	1.5×10^{-22}	5.5×10^{-21}	10^{-26}	3.3×10^{-22}	10^{-28}
2.0	5.4×10^{-21}	2.6×10^{-19}	1.4×10^{-23}	7.1×10^{-21}	10^{-25}
5.0	1.8×10^{-19}	1.3×10^{-17}	6.7×10^{-21}	1.4×10^{-19}	2.1×10^{-22}
10.0	1.2×10^{-18}	1.1×10^{-16}	2.3×10^{-19}	7.2×10^{-19}	1.2×10^{-20}
20.0	5.2×10^{-18}	4.2×10^{-16}	3.8×10^{-18}	2.5×10^{-18}	2.6×10^{-19}
50.0	2.1×10^{-17}	8.7×10^{-16}	5.4×10^{-17}	8.7×10^{-18}	5.3×10^{-18}
100.0	4.5×10^{-17}	8.5×10^{-16}	1.6×10^{-16}	1.9×10^{-17}	2.7×10^{-17}
200.0	8.8×10^{-17}	6.3×10^{-16}	2.4×10^{-16}	4.2×10^{-17}	9.2×10^{-17}
500.0	1.8×10^{-16}	3.7×10^{-16}	2.3×10^{-16}	8.4×10^{-17}	2.9×10^{-16}
1000.0	2.2×10^{-16}	2.7×10^{-16}	1.8×10^{-16}	8.0×10^{-17}	5.2×10^{-16}

For low energies ($T \lesssim 25 \text{ keV}$) the data may be represented by

$$(\overline{\sigma v})_{DD} = 2.33 \times 10^{-14} T^{-2/3} \exp(-18.76 T^{-1/3}) \text{ cm}^3 \text{ sec}^{-1};$$

$$(\overline{\sigma v})_{DT} = 3.68 \times 10^{-12} T^{-2/3} \exp(-19.94 T^{-1/3}) \text{ cm}^3 \text{ sec}^{-1},$$

where T is measured in keV.

The power density released in the form of charged particles is

$$P_{DD} = 3.3 \times 10^{-13} n_D^2 (\overline{\sigma v})_{DD} \text{ watt cm}^{-3} \text{ (including the subsequent D-T reaction);}$$

$$P_{DT} = 5.6 \times 10^{-13} n_D n_T (\overline{\sigma v})_{DT} \text{ watt cm}^{-3};$$

$$P_{D\text{He}^3} = 2.9 \times 10^{-12} n_D n_{\text{He}^3} (\overline{\sigma v})_{D\text{He}^3} \text{ watt cm}^{-3}.$$

```

\input prolog
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein
\pageno=46
\centerline{\headfont RELATIVISTIC ELECTRON BEAMS}
\medskip
\indent
Here  $\gamma = (1 - \beta^2)^{-1/2}$  is the relativistic scaling factor;
quantities in analytic formulas are expressed in SI or cgs units, as indicated;
in numerical formulas,  $I$  is in amperes (A),  $B$  is in gauss (G), electron
linear density  $N$  is in  $\text{cm}^{-1}$ , and temperature, voltage and energy are
in MeV;  $\beta_z = v_z/c$ ;  $k$  is Boltzmann's constant.
\medskip
Relativistic electron gyroradius:

$$r_e = \frac{\hbar c}{eB} (\gamma^2 - 1)^{-1/2} \quad (\text{in cgs}) =$$


$$1.70 \times 10^{-3} (\gamma^2 - 1)^{-1/2} B^{-1} \quad (\text{in cm}).$$


Relativistic electron energy:

$$W = \hbar c \gamma = 0.511 \gamma \quad (\text{in MeV}).$$


Bennett pinch condition:

$$I^2 = 2Nk(T_e + T_i) c^2 \quad (\text{in cgs}) = 3.20 \times 10^{-4} N(T_e + T_i) \quad (\text{in A}^2).$$


Alfvén-Lawson limit:

$$I_A = \frac{\hbar c^3}{e} \beta_z \gamma \quad (\text{in cgs}) = (4\pi \hbar c / \mu_0 e) \beta_z \gamma \quad (\text{in SI})$$


$$= 1.70 \times 10^{-4} \beta_z \gamma \quad (\text{in A}).$$


The ratio of net current to  $I_A$  is

$$\frac{I}{I_A} = \frac{\nu}{\gamma}.$$

Here  $\nu = N r_e$  is the Budker number, where  $r_e = e^2 / \hbar c^2 = 2.82 \times 10^{-13} \text{ cm}$  is the classical electron radius. Beam electron number density is

$$n_b = 2.08 \times 10^{-8} J \beta^{-1} \quad (\text{in cm}^{-3}),$$

where  $J$  is the current density in  $\text{A/cm}^2$ . For a uniform beam of
radius  $a$  (in cm),

$$n_b = 6.63 \times 10^{-7} I a^{-2} \beta^{-1} \quad (\text{in cm}^{-3}),$$

and

$$\frac{2r_e}{a} = \frac{\nu}{\gamma}.$$

\wfil\reject\end

```

RELATIVISTIC ELECTRON BEAMS

Here $\gamma = (1 - \beta^2)^{-1/2}$ is the relativistic scaling factor: quantities in analytic formulas are expressed in SI or cgs units, as indicated: in numerical formulas, I is in amperes (A), B is in gauss (G), electron linear density N is in cm^{-1} , and temperature, voltage and energy are in MeV; $\beta_z = v_z/c$; k is Boltzmann's constant.

Relativistic electron gyroradius:

$$r_e = \frac{mc^2}{eB}(\gamma^2 - 1)^{1/2} \text{ (cgs)} = 1.70 \times 10^3(\gamma^2 - 1)^{1/2} B^{-1} \text{ cm.}$$

Relativistic electron energy:

$$W = mc^2\gamma = 0.511\gamma \text{ MeV.}$$

Bennett pinch condition:

$$I^2 = 2Nk(T_e + T_i)c^2 \text{ (cgs)} = 3.20 \times 10^{-4}N(T_e + T_i) \text{ A}^2.$$

Alfvén-Lawson limit:

$$I_A = (mc^3/e)\beta_z\gamma \text{ (cgs)} = (4\pi mc/\mu_0 e)\beta_z\gamma \text{ (SI)} = 1.70 \times 10^4\beta_z\gamma \text{ A.}$$

The ratio of net current to I_A is

$$\frac{I}{I_A} = \frac{\nu}{\gamma}.$$

Here $\nu = Nr_e$ is the Budker number, where $r_e = e^2/mc^2 = 2.82 \times 10^{-13} \text{ cm}$ is the classical electron radius. Beam electron number density is

$$n_b = 2.08 \times 10^8 J \beta^{-1} \text{ cm}^{-3},$$

where J is the current density in A cm^{-2} . For a uniform beam of radius a (in cm),

$$n_b = 6.63 \times 10^7 I a^{-2} \beta^{-1} \text{ cm}^{-3},$$

and

$$\frac{2r_e}{a} = \frac{\nu}{\gamma}.$$

```

\input prolog
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsizer=9.0truein
\pageno=47
Child's law: (non-relativistic) space-charge-limited current density between
parallel plates with voltage drop  $V$  and separation  $d$  (in cm) is
 $J=2.34 \times 10^{-3} V^{3/2} d^{-2}$ , \rm A/cm-2.
\smallskip
The saturated parapotential current (magnetically self-limited flow along
equi-potentials in pinched diodes and transmission lines) is{29}
 $I_p=8.5 \times 10^{-3} G \gamma \ln \left[ \frac{\gamma + (\gamma^2 - 1)^{1/2}}{1} \right]$ , \rm A,
where  $G$  is a geometrical factor depending on the diode structure:
\smallskip
\begin{array}{l}
\text{for parallel plane cathode and anode} \\
\text{width } w, \text{ separation } d; \\
\text{displaystyle } G = \frac{\ln(R_2/R_1)}{2\pi d} \text{ for cylinders of} \\
\text{radius } R_1 \text{ (inner) and } R_2 \text{ (outer);} \\
\text{for conical cathode of radius } R_c, \text{ maximum} \\
\text{separation } d_0 \text{ (at } r=R_c \text{) from plane} \\
\text{anode.}
\end{array}
\smallskip
For  $\beta \rightarrow 0$ ; ( $\gamma \rightarrow \infty$ ), both  $I_A$  and  $I_p$  vanish.
\smallskip
The condition for a longitudinal magnetic field  $B_z$  to suppress
filamentation in a beam of current density  $J$  (in A/cm-2) is
 $B_z > 47 \beta \gamma^{1/2} J$ , \rm G.
Voltage registered by Rogowski coil of minor cross-sectional area  $A$ ,  $n$ 
turns, major radius  $a$ , inductance  $L$ , external resistance  $R$  and
capacitance  $C$  (all in SI):
\begin{array}{l}
\text{externally integrated } V = (1/RC)(nA\mu_0 I/2\pi a); \\
\text{self-integrating } V = (R/L)(nA\mu_0 I/2\pi a) = RI/n.
\end{array}
X-ray production, target with average atomic number  $Z$ ; ( $V$  approxlt
1 MeV):
\smallskip\nointerlineskip
\begin{array}{l}
\text{equiv } \text{x-ray power/beam power} = 7 \times 10^{-4} ZV \\
\text{ray dose at 1 meter generated by an e-beam depositing total charge } Q \\
\text{while } V \text{ goes from } V_{\text{max}} \text{ in material with charge state } Z: \\
\text{equiv } \text{dose} = 2.3 \times 10^{-4} Z^{1/2} Q \text{ rads.}
\end{array}
\smallskip\text{end}

```


Child's law: (non-relativistic) space-charge-limited current density between parallel plates with voltage drop V and separation d (in cm) is

$$J = 2.34 \times 10^3 V^{3/2} d^{-2} \text{ A cm}^{-2}.$$

The saturated parapotential current (magnetically self-limited flow along equipotentials in pinched diodes and transmission lines) is²⁹

$$I_p = 8.5 \times 10^3 G \gamma \ln [\gamma + (\gamma^2 - 1)^{1/2}] \text{ A},$$

where G is a geometrical factor depending on the diode structure:

$$\begin{aligned} G &= \frac{w}{2\pi d} && \text{for parallel plane cathode and anode} \\ &&& \text{of width } w, \text{ separation } d; \\ G &= \left(\ln \frac{R_2}{R_1} \right)^{-1} && \text{for cylinders of radii } R_1 \text{ (inner) and } R_2 \text{ (outer);} \\ G &= \frac{R_c}{d_0} && \text{for conical cathode of radius } R_c, \text{ maximum} \\ &&& \text{separation } d_0 \text{ (at } r = R_c \text{) from plane anode.} \end{aligned}$$

For $\beta \rightarrow 0$ ($\gamma \rightarrow 1$), both I_A and I_p vanish.

The condition for a longitudinal magnetic field B_z to suppress filamentation in a beam of current density J (in A cm^{-2}) is

$$B_z > 47\beta_z(\gamma J)^{1/2} \text{ G}.$$

Voltage registered by Rogowski coil of minor cross-sectional area A , n turns, major radius a , inductance L , external resistance R and capacitance C (all in SI):

$$\begin{aligned} \text{externally integrated} & \quad V = (1/RC)(nA\mu_0 I/2\pi a); \\ \text{self-integrating} & \quad V = (R/L)(nA\mu_0 I/2\pi a) = RI/n. \end{aligned}$$

X-ray production, target with average atomic number Z ($V \lesssim 5 \text{ MeV}$):

$$\eta \equiv \text{x-ray power/beam power} = 7 \times 10^{-4} ZV.$$

X-ray dose at 1 meter generated by an e-beam depositing total charge Q coulombs while $V \geq 0.84V_{\text{max}}$ in material with charge state Z :

$$D = 150V_{\text{max}}^{2.8} Q Z^{1/2} \text{ rads}.$$

```

\input prolog
\voffset=1.0truein\hoffset=0.25truein\ysize=9.0truein\hsize=6.5truein
\pageno=48
\centerline{\headfont{BEAM INSTABILITIES}\$~{30}\$}
\bksk % BEGINNING OF TABLE.
\vbox{\tabskip=0pt \offinterlineskip \halign to\hsize
{\strut#\vrule#\tabskip=3pt plus3pt minus2pt&\hfil\vrule#
&\$~{\hfil\vrule#\hfil\vrule#\tabskip=0pt\cr \trule \tska{3}{3pt}
\hfil Name\om\hfil Conditions\hfil\hfil Saturation Mechanism&\cr
\tska{3}{3pt} \trule \tska{3}{1pt} \trule \tska{3}{3pt}
|Electron-|V_d>\ov{V}_e\j=1,2|Electron trapping until &\cr
|\quad electron|\om|\quad\$~{\ov{V}_e}\sim V_d\$~\cr \tska{3}{6pt}
|Buneman|V_d>(M/m)^{1/3}\ov{V}_e|Electron trapping until&\cr
|\om|V_d>\ov{V}_e|\quad\$~{\ov{V}_e}\sim V_d\$~\cr \tska{3}{6pt}
|Beam-plasma|V_b>(n_p/n_b)^{1/3}\ov{V}_b|Trapping of beam electrons&\cr
\tska{3}{6pt}
|Weak beam-|V_b<(n_p/n_b)^{1/3}\ov{V}_b|Quasilinear or nonlinear&\cr
|\quad plasma|\om|\quad (mode coupling)&\cr \tska{3}{6pt}
|Beam-plasma|\ov{V}_e>V_b>\ov{V}_b|Quasilinear or nonlinear&\cr
|\quad (hot-electron)|\om|\om&\cr \tska{3}{6pt}
|Ion acoustic|T_e\gg T_i,\;V_d\gg C_s|Quasilinear, ion tail form-&\cr
|\om|\om|\quad ation, nonlinear scattering,&\cr
|\om|\om|\quad or resonance broadening.&\cr \tska{3}{6pt}
|Anisotropic|T_e\perp>2T_e\parallel|Isotropization&\cr \bs{1pt}
|\quad temperature|\om|\om&\cr
|\quad (hydro)|\om|\om&\cr \tska{3}{6pt}
|Ion cyclotron|V_d>20\ov{V}_e (for)|Ion heating&\cr
|\om|\om\hfil\$T_e\approx T_i\$|\om&\cr \tska{3}{6pt}
|Beam-cyclotron|V_d>C_s|Resonance broadening&\cr
|\quad (hydro)|\om|\om&\cr \tska{3}{6pt}
|Modified two-|V_d<(1+\beta)^{1/2}V_A|Trapping&\cr
|\quad stream (hydro)|V_d>C_s|\om&\cr \tska{3}{6pt}
|Ion-ion (equal|U<2(1+\beta)^{1/2}V_A|Ion trapping&\cr
|\quad beams)|\om|\om&\cr
\tska{3}{6pt}
|Ion-ion (equal|U<2C_s|Ion trapping&\cr
|\quad beams)|\om|\om&\cr \tska{3}{3pt} \trule}}

```

For nomenclature, see p. 50.

\vfil\egroup\end

BEAM INSTABILITIES³⁰

Name	Conditions	Saturation Mechanism
Electron-electron	$V_d > \bar{V}_{ej}, j = 1, 2$	Electron trapping until $\bar{V}_{ej} \sim V_d$
Buneman	$V_d > (M/m)^{1/3} \bar{V}_i,$ $V_d > \bar{V}_e$	Electron trapping until $\bar{V}_e \sim V_d$
Beam-plasma	$V_b > (n_p/n_b)^{1/3} \bar{V}_b$	Trapping of beam electrons
Weak beam-plasma	$V_b < (n_p/n_b)^{1/3} \bar{V}_b$	Quasilinear or nonlinear (mode coupling)
Beam-plasma (hot-electron)	$\bar{V}_e > V_b > \bar{V}_b$	Quasilinear or nonlinear
Ion acoustic	$T_e \gg T_i, V_d \gg C_s$	Quasilinear, ion tail formation, nonlinear scattering, or resonance broadening.
Anisotropic temperature (hydro)	$T_{e\perp} > 2T_{e\parallel}$	Isotropization
Ion cyclotron	$V_d > 20\bar{V}_i$ (for $T_e \approx T_i$)	Ion heating
Beam-cyclotron (hydro)	$V_d > C_s$	Resonance broadening
Modified two-stream (hydro)	$V_d < (1 + \beta)^{1/2} V_A,$ $V_d > C_s$	Trapping
Ion-ion (equal beams)	$U < 2(1 + \beta)^{1/2} V_A$	Ion trapping
Ion-ion (equal beams)	$U < 2C_s$	Ion trapping

For nomenclature, see p. 50.

```

\input prolog \pageno=49
\voffset=1.0truein\hoffset=0.25truein\ysize=9.0truein\hsize=6.5truein
\ vbox{\tabskip=0pt \offinterlineskip \def\ds{\displaystyle}
% HERE, \ds IS USED FOR 'DISPLAYSTYLE'.
\halign to\hsize{\strut#\vrule#\tabskip=3pt plus3pt minus2pt#\hfil#\vrule#
&\hfil$#\hfil#\vrule#\hfil$#\hfil#\vrule#\hfil$#\hfil#\vrule#
&\hfil$#\hfil#\vrule#\tabskip=0pt\cr\trule
\om&height3pt&\om\multispan7&\cr
\om\multispan7\hfil Parameters of Most Unstable Mode\hfil&\cr
\om&height3pt&\om\multispan7&\cr
\om\om&\multispan9\hrulefill\cr \tska{5}{2pt} \bs{0.75ex}
% THIS USE OF \hrulefill IS MORE GENERAL AND ELEGANT IN MANY CASES
% THAN THE 'HARD-WIRED' \hrule USED UNDER 'Dimension' ON PAGE 10.
|\hfil\raise1ex\hbox{Name}\om\om\{\rm Wave}\{\rm Group}&\cr
|\om\{\rm Growth\ Rate}\{\rm Frequency}\{\rm Number}\{\rm Velocity}&\cr
\tska{5}{2pt} \trule \tska{5}{1pt} \trule \tska{5}{2pt}
|Electron-\ds{1\over 2}\omega_e|0|\ds{0.9\omega_e\over V_d}|0&\cr \bs{5pt}
|\quad electron|\om\om\om\om&\cr \tska{5}{2pt}
|Buneman|\ds{0.7\left(\frac{m}{M}\right)^{1/3}\omega_e|
\ds{0.4\left(\frac{m}{M}\right)^{1/3}\omega_e|\ds{\omega_e
\over V_d}|\ds{2\over 3}V_d&\cr \tska{5}{2pt}
|Beam-plasma|\ds{0.7\left(\frac{n_b}{n_p}\right)^{1/3}\omega_e|\om$\omega_e-$
\hfil|\ds{\omega_e\over V_b}|\ds{2\over 3}V_b&\cr
|\om\om\om\hfil$\ds{0.4\left(\frac{n_b}{n_p}\right)^{1/3}\omega_e|
\om\om&\cr \tska{5}{2pt}
|Weak beam-\ds{\frac{n_b}{2n_p}}\left(\frac{V_b}{V_b}\right)^2\omega_e
|\omega_e|\ds{\omega_e\over V_b}|\ds{3\over 2}\frac{V_e}{V_b}&\cr \bs{6pt}
|\quad plasma|\om\om\om\om&\cr \tska{5}{2pt}
|Beam-plasma|\ds\left(\frac{n_b}{n_p}\right)^{1/2}\frac{V_e}{V_b}\omega_e
|\ds{\frac{V_b}{V_e}}\omega_e|\lambda_D^{-1}|V_b&\cr \bs{5pt}
|\quad (hot-electron)|\om\om\om\om&\cr \tska{5}{2pt}
|Ion acoustic|\ds\left(\frac{m}{M}\right)^{1/2}\omega_i|\ds{\omega_i
|\lambda_D^{-1}|C_s&\cr \tska{5}{2pt}
|Anisotropic|\Omega_e\omega_e\cos\theta\sim\Omega_e|r_e^{-1}|
\ov{V}_e\perp&\cr \bs{1pt}
|\quad temperature|\om\om\om\om&\cr
|\quad (hydro)|\om\om\om\om&\cr \tska{5}{2pt}
|Ion cyclotron|0.1\Omega_i|1.2\Omega_i|r_i^{-1}|\ds{1\over 3}\ov{V}_i&\cr
\tska{5}{2pt}
|Beam-cyclotron|\ds{0.7\Omega_e|\ds{n\Omega_e|0.7
|\lambda_D^{-1}|\approx V_d;&\cr \bs{1pt}
|\quad (hydro)|\om\om\om\approx C_s\phi;&\cr \tska{5}{2pt}
|Modified two-\ds{1\over 2}\Omega_H|\ds{0.9\Omega_H|\ds{1.7
\Omega_H\over V_d}|\ds{1\over 2}V_d&\cr \bs{5pt}
|\quad stream|\om\om\om\om&\cr
|\quad (hydro)|\om\om\om\om&\cr \tska{5}{2pt}
|Ion-ion (equal|0.4\Omega_H|0|\ds{1.2\Omega_H\over U}|0&\cr \bs{4pt}
|\quad beams)|\om\om\om\om&\cr \tska{5}{2pt}
|Ion-ion (equal|\ds{0.4\omega_i|0|\ds{1.2\omega_i\over U}|0&\cr \bs{4pt}
|\quad beams)|\om\om\om\om&\cr \tska{5}{2pt}\trule}}
For nomenclature, see p. 50.
\vfil\egroup\end

```

Name	Parameters of Most Unstable Mode			
	Growth Rate	Frequency	Wave Number	Group Velocity
Electron-electron	$\frac{1}{2}\omega_e$	0	$0.9\frac{\omega_e}{V_d}$	0
Buneman	$0.7\left(\frac{m}{M}\right)^{1/3}\omega_e$	$0.4\left(\frac{m}{M}\right)^{1/3}\omega_e$	$\frac{\omega_e}{V_d}$	$\frac{2}{3}V_d$
Beam-plasma	$0.7\left(\frac{n_b}{n_p}\right)^{1/3}\omega_e$	$\omega_e -$ $0.4\left(\frac{n_b}{n_p}\right)^{1/3}\omega_e$	$\frac{\omega_e}{V_b}$	$\frac{2}{3}V_b$
Weak beam-plasma	$\frac{n_b}{2n_p}\left(\frac{V_b}{\bar{V}_b}\right)^2\omega_e$	ω_e	$\frac{\omega_e}{V_b}$	$\frac{3\bar{V}_e^2}{V_b}$
Beam-plasma (hot-electron)	$\left(\frac{n_b}{n_p}\right)^{1/2}\frac{\bar{V}_e}{V_b}\omega_e$	$\frac{V_b}{\bar{V}_e}\omega_e$	λ_D^{-1}	V_b
Ion acoustic	$\left(\frac{m}{M}\right)^{1/2}\omega_i$	ω_i	λ_D^{-1}	C_s
Anisotropic temperature (hydro)	Ω_e	$\omega_e \cos \theta \sim \Omega_e$	r_e^{-1}	$\bar{V}_{e\perp}$
Ion cyclotron	$0.1\Omega_i$	$1.2\Omega_i$	r_i^{-1}	$\frac{1}{3}\bar{V}_i$
Beam-cyclotron (hydro)	$0.7\Omega_e$	$n\Omega_e$	$0.7\lambda_D^{-1}$	$\gtrsim V_d;$ $\lesssim C_s$
Modified two-stream (hydro)	$\frac{1}{2}\Omega_H$	$0.9\Omega_H$	$1.7\frac{\Omega_H}{V_d}$	$\frac{1}{2}V_d$
Ion-ion (equal beams)	$0.4\Omega_H$	0	$1.2\frac{\Omega_H}{U}$	0
Ion-ion (equal beams)	$0.4\omega_i$	0	$1.2\frac{\omega_i}{U}$	0

For nomenclature, see p. 50.

```

\input prolog \pageno=50
\hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein\indent
In the preceding tables, subscripts $e$, $i$, $d$, $b$, $p$ stand for
'elec\tron,' 'ion,' 'drift,' 'beam,' and 'plasma,' respectively.
Thermal velocities are denoted by a bar. In addition, the following are used: \ssk
\halign {$#\hfil\quad#\hfil\quad#\hfil\quad#\hfil\cr
m &electron mass &r_e, r_i &gyroradius \cr
M &ion mass &\beta &plasma/magnetic energy \cr
V &velocity &\quad &density ratio \cr
T &temperature &V_A &Alfvén speed \cr
n_e, n_i &number density &\Omega_e, \Omega_i &gyrofrequency \cr
n &harmonic number &\Omega_H &hybrid gyrofrequency, \cr
C_s=(T_e/M)^{1/2} &ion sound speed &\quad &{\Omega_H}^2=\Omega_e\Omega_i \cr
\omega_e, \omega_i &plasma frequency &U &relative drift velocity of \cr
\lambda_D &Debye length &\quad &two ion species \cr} \bsk
\centerline{\headfont LASERS} \msk
{\headfont System Parameters} \ssk
Efficiencies and power levels are approximately state-of-the-art (1987).$^{31}$\ssk
\vbox{\offinterlineskip \tabskip=Opt \halign to \hsize{
\vrule# \tabskip=3pt plus 2pt minus 2pt&\strut #\hfil
&\vrule #&\hfil#\hfil &\vrule #&\hfil#\hfil
&\vrule #&\hfil#\hfil &\vrule #&\hfil#\hfil
&\vrule #\tabskip=Opt\cr\trule
height 2pt&\oml\oml\oml\multispan3 &\cr
&\oml\oml\oml\multispan3 Power levels available (W) &\cr \bs{6.5pt}
&\hfil Type{\displaystyle{\rm Wavelength\atop(\mu m)}}{\displaystyle{\rm Efficiency\atop(\%)}}\cr
&\oml\oml\cr \bs{8pt}
&\oml\oml\oml\multispan3=Opt\multispan3&\cr
\lnoalign{\vskip-0.4ex \moveright3.215truein \vbox{\hrule width2.79truein}}
\tskc{5}{3pt} &\oml\oml\oml\Pulsed/CW&\cr
\tskc{5}{3pt} \trule \tskc{5}{1pt} \trule \tskc{5}{2pt}
&CO$_2$|10.6|0.01--0.02|>2\times 10^{13}$|>10^5$&\cr
&\oml\oml(pulsed)|\oml\oml\cr \tskc{5}{1pt}
&CO|5|0.4|>10^9$|>100$&\cr \tskc{5}{1pt}
&Holmium|2.06|0.03|>10^7$|30&\cr \tskc{5}{1pt}
&Iodine|1.315|0.003|>10^{12}$|--&\cr \tskc{5}{1pt}
&Nd-glass,|1.06|0.001--0.06|sim 10^{14}$|1--10^3$&\cr
&\quad YAG|\oml\oml(10-beam system)|\oml\oml\cr \tskc{5}{1pt}
&\llap{*}Color|1--4|10^{-3}$|>10^6$|1&\cr
&\quad center|\oml\oml\oml\oml\cr \tskc{5}{1pt}
&\llap{*}OPD|0.7--0.9|10^{-3}$|10^6$|1&\cr \tskc{5}{1pt}
&Ruby|0.6943|<10^{-3}$|10^{10}$|1&\cr \tskc{5}{1pt}
&He-Ne|0.6328|10^{-4}$|1--50\times 10^{-3}$&\cr \tskc{5}{1pt}
&\llap{*}Argon ion|0.45--0.6|10^{-3}$|5\times 10^4$|1--10&\cr \tskc{5}{1pt}
&Nd$_2$|0.3371|0.001--0.05|10^5$--10^6$|--&\cr \tskc{5}{1pt}
&\llap{*}Dye|0.3--1.1|10^{-3}$|>10^6$|140&\cr \tskc{5}{1pt}
&Kr-F|0.26|0.08|>10^9$|--&\cr \tskc{5}{1pt}
&Xe|0.175|0.02|>10^8$|--&\cr \tskc{5}{2pt} \trule \lnoalign{\ssk}
\oml\hbox{*Tunable sources}\hidewidth\cr}} \ssk % END OF TABLE.
YAG stands for Yttrium--Aluminum Garnet and OPD for Optical Parametric
Oscillator.
\vfll\ject\end

```

In the preceding tables, subscripts *e*, *i*, *d*, *b*, *p* stand for "electron," "ion," "drift," "beam," and "plasma," respectively. Thermal velocities are denoted by a bar. In addition, the following are used:

m	electron mass	r_e, r_i	gyroradius
M	ion mass	β	plasma/magnetic energy
V	velocity		density ratio
T	temperature	V_A	Alfvén speed
n_e, n_i	number density	Ω_e, Ω_i	gyrofrequency
n	harmonic number	Ω_H	hybrid gyrofrequency,
$C_s = (T_e/M)^{1/2}$	ion sound speed		$\Omega_H^2 = \Omega_e \Omega_i$
ω_e, ω_i	plasma frequency	U	relative drift velocity of
λ_D	Debye length		two ion species

LASERS

System Parameters

Efficiencies and power levels are approximately state-of-the-art (1987).³¹

Type	Wavelength (μm)	Efficiency	Power levels available (W)	
			Pulsed	CW
CO ₂	10.6	0.01-0.02 (pulsed)	$> 2 \times 10^{13}$	$> 10^5$
CO	5	0.4	$> 10^9$	> 100
Holmium	2.06	0.03	$> 10^7$	30
Iodine	1.315	0.003	$> 10^{12}$	
Nd-glass, YAG	1.06	0.001-0.06	$\sim 10^{14}$ (10-beam system)	1- 10^3
*Color center	1-4	10^{-3}	$> 10^6$	1
*OPO	0.7-0.9	10^{-3}	10^6	1
Ruby	0.6943	$< 10^{-3}$	10^{10}	1
He-Ne	0.6328	10^{-4}		1- 50×10^{-3}
*Argon ion	0.45-0.60	10^{-3}	5×10^4	1-10
N ₂	0.3371	0.001-0.05	10^5 - 10^6	
*Dye	0.3-1.1	10^{-3}	$> 10^6$	140
Kr-F	0.26	0.08	$> 10^9$	
Xenon	0.175	0.02	$> 10^8$	

*Tunable sources

YAG stands for Yttrium Aluminum Garnet and OPO for Optical Parametric Oscillator.

```

\input prolog
\pageno=51
\hoffset=1.25truein
\voffset=1.0truein
\hsize=6.0truein
\vsize=9.0truein
{\headfont Formulas}
\medskip\indent
An e-m wave with {\bf k} $\parallel$ {\bf B} has an index of refraction
given by

$$n_{\pm} = [1 - \omega_{pe}^2 \cos^2 \theta / (\omega(\omega \mp \omega_{ce}))]^{\pm 1/2},$$

where  $\pm$  refers to the helicity. The rate of change of polarization
angle  $\theta$  as a function of displacement  $s$  (Faraday rotation) is given
by

$$d\theta/ds = (k/2)(n_{-} - n_{+}) = 2.36 \times 10^{-4} N_e B^2 \cos^2 \theta,$$

where  $N_e$  is the electron number density,  $B$  is the field strength, and
 $f$  is the wave frequency, all in cgs.
\smallskip\indent
The quiver velocity of an electron in an e-m field of angular frequency  $\omega$ 
is

$$v_0 = eE_{\max} / m\omega = 25.6 I^{1/2} \lambda_0 \text{ cm sec}^{-1}$$

in terms of the laser flux  $I = cE_{\max}^2 \cos^2 \theta / 8\pi$ , with  $I$  in
watt/cm2, laser wavelength  $\lambda_0$  in  $\mu\text{m}$ . The ratio of quiver
energy to thermal energy is

$$W_{\text{qu}} / W_{\text{th}} = m_e v_0^2 / 2kT = 1.81 \times 10^{-13} \lambda_0^2 I / T,$$

where  $T$  is given in eV. For example, if  $I = 10^{15} \text{ W cm}^{-2}$ ,
 $\lambda_0 = 1 \mu\text{m}$ ,  $T = 2 \text{ keV}$ , then  $W_{\text{qu}} / W_{\text{th}} \approx 0.1$ .
\smallskip\indent
Pondermotive force:

$$\mathbf{F} = -\nabla \langle E^2 \rangle / 8\pi N_c,$$

where

$$N_c = 1.1 \times 10^{21} \lambda_0^{-2} \text{ cm}^{-3}.$$

\indent
For uniform illumination of a lens with  $F$ , the diameter  $d$ 
at focus (85% of the energy) and the depth of focus  $l$  (distance to first
zero in intensity) are given by

$$d \approx 2.44 F \lambda \theta / \theta_{DL} \quad \text{and} \quad l \approx \pm 2 F^2 \lambda \theta / \theta_{DL}.$$

Here  $\theta$  is the beam divergence containing 85% of energy and
 $\theta_{DL}$  is the diffraction-limited divergence:

$$\theta_{DL} = 2.44 \lambda / b,$$

where  $b$  is the aperture. These formulas are modified for nonuniform (such
as Gaussian) illumination of the lens or for pathological laser profiles.
\fill\egroup\end

```


Formulas

An e-m wave with $\mathbf{k} \parallel \mathbf{B}$ has an index of refraction given by

$$n_{\pm} = [1 - \omega_{pe}^2 / \omega(\omega \mp \omega_{ce})]^{1/2},$$

where \pm refers to the helicity. The rate of change of polarization angle θ as a function of displacement s (Faraday rotation) is given by

$$d\theta/ds = (k/2)(n_- - n_+) = 2.36 \times 10^4 N B f^{-2} \text{ cm}^{-1},$$

where N is the electron number density, B is the field strength, and f is the wave frequency, all in cgs.

The quiver velocity of an electron in an e-m field of angular frequency ω is

$$v_0 = eE_{\max}/m\omega = 25.6 I^{1/2} \lambda_0 \text{ cm sec}^{-1}$$

in terms of the laser flux $I = cE_{\max}^2/8\pi$, with I in watt/cm², laser wavelength λ_0 in μm . The ratio of quiver energy to thermal energy is

$$W_{\text{qu}}/W_{\text{th}} = m_e v_0^2 / 2kT = 1.81 \times 10^{-13} \lambda_0^2 I / T,$$

where T is given in eV. For example, if $I = 10^{15} \text{ W cm}^{-2}$, $\lambda_0 = 1 \mu\text{m}$, $T = 2 \text{ keV}$, then $W_{\text{qu}}/W_{\text{th}} \approx 0.1$.

Pondermotive force:

$$\mathcal{F} = N \nabla \langle E^2 \rangle / 8\pi N_e,$$

where

$$N_e = 1.1 \times 10^{21} \lambda_0^{-2} \text{ cm}^{-3}.$$

For uniform illumination of a lens with f -number F , the diameter d at focus (85% of the energy) and the depth of focus l (distance to first zero in intensity) are given by

$$d \approx 2.44 F \lambda \theta / \theta_{DL} \quad \text{and} \quad l \approx \pm 2 F^2 \lambda \theta / \theta_{DL}.$$

Here θ is the beam divergence containing 85% of energy and θ_{DL} is the diffraction-limited divergence:

$$\theta_{DL} = 2.44 \lambda / b,$$

where b is the aperture. These formulas are modified for nonuniform (such as Gaussian) illumination of the lens or for pathological laser profiles.

```

\input prolog
\hoffset=1.0truein
\voffset=1.0truein
\hsize=6.125truein
\vsize=9.0truein
\pageno=52
\parskip=5pt
\centerline{\headfont ATOMIC PHYSICS AND RADIATION}
\bigskip\inndent
Energies and temperatures are in eV; all other units are cgs except where
noted.  $Z$  is the charge state ( $Z=0$  refers to a neutral atom); the subscript
 $e$  labels electrons.  $N$  refers to number density,  $n$  to principal quantum
number. Asterisk superscripts on level population densities denote local
thermodynamic equilibrium (LTE) values. Thus  $N_n \text{hbox{*}}$  is the LTE number
density of atoms (or ions) in level  $n$ .

Characteristic atomic collision cross section:

$$\pi a_0^2 = 8.80 \times 10^{-17} \text{ cm}^2. \text{leqno(1)}$$

Binding energy of outer electron in level labelled by quantum numbers  $n, l$ :

$$E_{\infty}^Z(n,l) = -\{Z^2 E_{\infty}^H \over (n - \Delta_l)^2\}, \text{leqno(2)}$$

where  $E_{\infty}^H = 13.6 \text{ eV}$  is the hydrogen ionization energy and
 $\Delta_l = 0.751 \times 10^{-5}$ ,  $\Delta_l \approx 5$ , is the quantum defect.
\medskip
{\headfont Excitation and Decay}

\inndent
Cross section (Bethe approximation) for electron excitation by dipole allowed
transition  $m \rightarrow n$  (Refs. 32, 33):

$$\sigma_{mn} = 2.36 \times 10^{-13} f_{nm} g(n,m) \over \epsilon \Delta E_{nm} \text{ cm}^2, \text{leqno(3)}$$

where  $f_{nm}$  is the oscillator strength,  $g(n,m)$  is the Gaunt factor,
 $\epsilon$  is the incident electron energy, and  $\Delta E_{nm} = E_n - E_m$ .

Electron excitation rate averaged over Maxwellian velocity distribution,
 $X_{mn} = N_e \langle \sigma_{mn} v \rangle$  (Refs. 34, 35):

$$X_{mn} = 1.6 \times 10^{-5} f_{nm} \langle g(n,m) \rangle N_e \over \Delta E_{nm} T_e^{-1/2} \exp\left(-\Delta E_{nm} \over T_e\right) \text{ sec}^{-1}, \text{leqno(4)}$$

where  $\langle g(n,m) \rangle$  denotes the thermal averaged Gaunt factor
(generally  $\sim 1$  for atoms,  $\sim 0.2$  for ions).
\fill\egroup\end

```

ATOMIC PHYSICS AND RADIATION

Energies and temperatures are in eV; all other units are cgs except where noted. Z is the charge state ($Z = 0$ refers to a neutral atom); the subscript e labels electrons. N refers to number density, n to principal quantum number. Asterisk superscripts on level population densities denote local thermodynamic equilibrium (LTE) values. Thus N_n^* is the LTE number density of atoms (or ions) in level n .

Characteristic atomic collision cross section:

$$(1) \quad \pi a_0^2 = 8.80 \times 10^{-17} \text{ cm}^2.$$

Binding energy of outer electron in level labelled by quantum numbers n, l :

$$(2) \quad E_\infty^Z(n, l) = -\frac{Z^2 E_\infty^H}{(n - \Delta_l)^2},$$

where $E_\infty^H = 13.6 \text{ eV}$ is the hydrogen ionization energy and $\Delta_l = 0.75l^{-5}$, $l \gtrsim 5$, is the quantum defect.

Excitation and Decay

Cross section (Bethe approximation) for electron excitation by dipole allowed transition $m \rightarrow n$ (Refs. 32, 33):

$$(3) \quad \sigma_{m \rightarrow n} = 2.36 \times 10^{-13} \frac{f_{nm} g(n, m)}{\epsilon \Delta E_{nm}} \text{ cm}^2,$$

where f_{nm} is the oscillator strength, $g(n, m)$ is the Gaunt factor, ϵ is the incident electron energy, and $\Delta E_{nm} = E_n - E_m$.

Electron excitation rate averaged over Maxwellian velocity distribution, $X_{mn} = N_e \langle \sigma_{mn} v \rangle$ (Refs. 34, 35):

$$(4) \quad X_{mn} = 1.6 \times 10^{-5} \frac{f_{nm} \langle g(n, m) \rangle N_e}{\Delta E_{nm} T_e^{1/2}} \exp \left(-\frac{\Delta E_{nm}}{T_e} \right) \text{ sec}^{-1}.$$

where $\langle g(n, m) \rangle$ denotes the thermal averaged Gaunt factor (generally ~ 1 for atoms, ~ 0.2 for ions).

prolog
hoffset=1.0truein
voffset=1.0truein
hsize=6.125truein
vsize=9.0truein
pageno=53
parskip=5pt
Rate for electron collisional deexcitation:

$$Y_{nm} = (N_m / N_n) X_{mn} \leq (5)$$
Here $(N_m / N_n) = (g_m / g_n) \exp(-\Delta E_{nm} / T_e)$ is the Boltzmann relation for level population densities, where g_n is the statistical weight of level n .

Rate for spontaneous decay $n \rightarrow m$ (Einstein A coefficient)³⁴

$$A_{nm} = 4.3 \times 10^{-7} (g_n / g_m) f_{nm} (\Delta E_{nm})^2, \text{ (sec)}^{-1} \leq (6)$$

Intensity emitted per unit volume from the transition $n \rightarrow m$ in an optically thin plasma:

$$I_{nm} = 1.6 \times 10^{-19} A_{nm} N_n \Delta E_{nm}, \text{ (watt/cm)}^3 \leq (7)$$

Condition for steady state in a corona model:

$$N_{0n} \langle \sigma_{0n} \rangle v = N_n A_{n0}, \leq (8)$$
where the ground state is labelled by a zero subscript.

Hence for a transition $n \rightarrow m$ in ions, where $\langle g(n,0) \rangle \approx 0.2$,

$$I_{nm} = 5.1 \times 10^{-25} \{ f_{nm} g_0 N_{e0} / g_m T_e^{1/2} \} \left(\Delta E_{nm} / \Delta E_{n0} \right)^3 \exp \left(- \Delta E_{n0} / T_e \right), \text{ (watt/cm)}^3 \leq (9)$$

\medskip
{\headfont Ionization and Recombination}
\medskip

\inndent

In a general time-dependent situation the number density of the charge state Z satisfies

$$\frac{dN(Z)}{dt} = N_e \left[-S(Z)N(Z) - \alpha(Z)N(Z) \right] \leq (10)$$

\vskip-0.3truein

$$+ S(Z-1)N(Z-1) + \alpha(Z+1)N(Z+1) \big].$$
Here $S(Z)$ is the ionization rate. The recombination rate $\alpha(Z)$ has the form $\alpha(Z) = \alpha_r(Z) + N_e \alpha_3(Z)$, where α_r and α_3 are the radiative and three-body recombination rates, respectively.

\vfill\rejection

Rate for electron collisional deexcitation:

$$(5) \quad Y_{nm} = (N_m^*/N_n^*)X_{mn}.$$

Here $N_m^*/N_n^* = (g_m/g_n) \exp(\Delta E_{nm}/T_e)$ is the Boltzmann relation for level population densities, where g_n is the statistical weight of level n .

Rate for spontaneous decay $n \rightarrow m$ (Einstein A coefficient)³⁴

$$(6) \quad A_{nm} = 4.3 \times 10^7 (g_n/g_m) f_{nm} (\Delta E_{nm})^2 \text{ sec}^{-1}.$$

Intensity emitted per unit volume from the transition $n \rightarrow m$ in an optically thin plasma:

$$(7) \quad I_{nm} = 1.6 \times 10^{-19} A_{nm} N_n \Delta E_{nm} \text{ watt/cm}^3.$$

Condition for steady state in a corona model:

$$(8) \quad N_0 N_e \langle \sigma_{0n} v \rangle = N_n A_{n0},$$

where the ground state is labelled by a zero subscript.

Hence for a transition $n \rightarrow m$ in ions, where $\langle g(n, 0) \rangle \approx 0.2$,

$$(9) \quad I_{nm} = 5.1 \times 10^{-25} \frac{f_{nm} g_0 N_e N_0}{g_m T_e^{1/2}} \left(\frac{\Delta E_{nm}}{\Delta E_{n0}} \right)^3 \exp \left(- \frac{\Delta E_{n0}}{T_e} \right) \frac{\text{watt}}{\text{cm}^3}.$$

Ionization and Recombination

In a general time-dependent situation the number density of the charge state Z satisfies

$$(10) \quad \frac{dN(Z)}{dt} = N_e \left[-S(Z)N(Z) - \alpha(Z)N(Z) + S(Z-1)N(Z-1) + \alpha(Z+1)N(Z+1) \right].$$

Here $S(Z)$ is the ionization rate. The recombination rate $\alpha(Z)$ has the form $\alpha(Z) = \alpha_r(Z) + N_e \alpha_3(Z)$, where α_r and α_3 are the radiative and three-body recombination rates, respectively.

\input prolog
 \hoffset=1.0truein
 \voffset=1.0truein
 \hsize=6.125truein
 \vsize=9.0truein
 \pageno=54
 \parskip=5pt
 Classical ionization cross-section^{36} for any atomic shell j
 $\sigma_{\text{sig}} = 6 \times 10^{-14} \text{ b}_{\text{ij}} g_j(x) / \{U_j\}^2 \text{ cm}^2$. ^{11}
 Here b_{ij} is the number of shell electrons; U_j is the binding energy
 of the ejected electron; $x = \epsilon / U_j$, where ϵ is the incident
 electron energy; and g is a universal function with a minimum value g_{min}
 approx 0.2 at x approx 4.

Ionization from ion ground state, averaged over Maxwellian electron
 distribution, for $0.02 \text{ approx } T_e / E_{\infty} \text{ approx } 100$ (Ref. 35):
 $S(Z) = 10^{-5} \{ (T_e / E_{\infty})^{1/2} / (E_{\infty})^{3/2} (6.0 + T_e / E_{\infty}) \} \exp \left(- \{ E_{\infty} / T_e \} \right) \text{ cm}^3 / \text{sec}$,
^{12}
 where E_{∞} is the ionization energy.

Electron-ion radiative recombination rate $\left(e + M(Z) \rightarrow M(Z-1) + h\nu \right)$ for $T_e / Z^2 \text{ approx } 400$ (Ref. 37):
 $\alpha_r(Z) = 5.2 \times 10^{-14} Z \left(\{ E_{\infty} / T_e \} \right)^{1/2} \left[0.43 + \{ 1 / 2 \} \ln(E_{\infty} / T_e) \right] \text{ cm}^3 / \text{sec}$.
^{13}
 $\alpha_r \approx 0.469 (E_{\infty} / T_e)^{-1/3} \text{ cm}^3 / \text{sec}$.

For $1 \text{ eV} < T_e / Z^2 < 15 \text{ eV}$, this becomes approximately^{35}
 $\alpha_r(Z) = 2.7 \times 10^{-13} Z^2 \{ T_e \}^{-1/2} \text{ cm}^3 / \text{sec}$. ^{14}

Collisional (three-body) recombination rate for singly ionized plasma:^{38}
 $\alpha_3 = 3.75 \times 10^{-27} \{ T_e \}^{-4.5} \text{ cm}^6 / \text{sec}$. ^{15}

Photoionization cross section for ions in level n, l (short-wavelength limit):
 $\sigma_{\text{ph}}(n, l) = 1.34 \times 10^{-16} Z^5 / n^3 K^{7+2l} \text{ cm}^2$,
^{16}
 where K is the wavenumber in Rydbergs ($1 \text{ Rydberg} = 1.0974 \times 10^7 \text{ cm}^{-1}$).
 \fill eject\end

Classical ionization cross-section³⁶ for any atomic shell j

$$(11) \quad \sigma_i = 6 \times 10^{-14} b_j g_j(x) / U_j^2 \text{ cm}^2.$$

Here b_j is the number of shell electrons; U_j is the binding energy of the ejected electron; $x = \epsilon/U_j$, where ϵ is the incident electron energy; and g is a universal function with a minimum value $g_{\min} \approx 0.2$ at $x \approx 4$.

Ionization from ion ground state, averaged over Maxwellian electron distribution, for $0.02 \lesssim T_e/E_\infty^Z \lesssim 100$ (Ref. 35):

$$(12) \quad S(Z) = 10^{-5} \frac{(T_e/E_\infty^Z)^{1/2}}{(E_\infty^Z)^{3/2}(6.0 + T_e/E_\infty^Z)} \exp\left(-\frac{E_\infty^Z}{T_e}\right) \text{ cm}^3/\text{sec},$$

where E_∞^Z is the ionization energy.

Electron-ion radiative recombination rate ($e + N(Z) \rightarrow N(Z-1) + h\nu$) for $T_e/Z^2 \lesssim 400 \text{ eV}$ (Ref. 37):

$$(13) \quad \alpha_r(Z) = 5.2 \times 10^{-14} Z \left(\frac{E_\infty^Z}{T_e}\right)^{1/2} \left[0.43 + \frac{1}{2} \ln(E_\infty^Z/T_e) + 0.469(E_\infty^Z/T_e)^{-1/3}\right] \text{ cm}^3/\text{sec}.$$

For $1 \text{ eV} < T_e/Z^2 < 15 \text{ eV}$, this becomes approximately³⁵

$$(14) \quad \alpha_r(Z) = 2.7 \times 10^{-13} Z^2 T_e^{-1/2} \text{ cm}^3/\text{sec}.$$

Collisional (three-body) recombination rate for singly ionized plasma:³⁸

$$(15) \quad \alpha_3 = 8.75 \times 10^{-27} T_e^{-4.5} \text{ cm}^6/\text{sec}.$$

Photoionization cross section for ions in level n, l (short-wavelength limit):

$$(16) \quad \sigma_{\text{ph}}(n, l) = 1.64 \times 10^{-16} Z^5 / n^3 K^{7+2l} \text{ cm}^2,$$

where K is the wavenumber in Rydbergs ($1 \text{ Rydberg} = 1.0974 \times 10^5 \text{ cm}^{-1}$).

```

\input prolog
\hoffset=1.0truein
\voffset=1.0truein
\hsize=6.125truein
\vsize=9.0truein
\pageno=55
\parskip=5pt
{\headfont Ionization Equilibrium Models}

\inndent
Saha equilibrium:$^{39}$

$$\frac{N_{eN_1}(Z)}{N_{eN}(Z-1)} = 6.0 \times 10^{21} \frac{g_1^Z T_e^{-3/2}}{g_{n-1}^{Z-1} \exp\left(-\frac{E_{\infty}^Z(n,1)}{T_e}\right)} \text{ cm}^{-3},$$

\leqno{(17)}
where $g_n^Z$ is the statistical weight for level $n$ of charge state
$Z$ and $E_{\infty}^Z(n,1)$ is the ionization energy of the neutral atom
initially in level $(n, 1)$, given by Eq.(2).

In a steady state at high electron density,

$$\frac{N_{eN_1}(Z)}{N_{eN}(Z-1)} = \frac{S(Z-1)}{\alpha_3}, \text{ } \leqno{(18)}$$

a function only of $T$.

Conditions for LTE:$^{39}$

(a) Collisional and radiative excitation rates for a level $n$ must satisfy

$$Y_{nm} \approx 10 A_{nm}, \text{ } \leqno{(19)}$$


(b) Electron density must satisfy

$$N_e \approx 7 \times 10^{18} Z^{-7} n^{-17/2} (T/E_{\infty}^Z)^{1/2} \text{ cm}^{-3},$$

\leqno{(20)}

Steady state condition in corona model:

$$\frac{N(Z-1)}{N(Z)} = \frac{\alpha_r}{S(Z-1)}, \text{ } \leqno{(21)}$$


Corona model is applicable if$^{40}$

$$10^{12} t_I^{-1} < N_e < 10^{16} T_e^{-7/2} \text{ cm}^{-3}, \text{ } \leqno{(22)}$$

where $t_I$ is the ionization time.
\fillreject\end

```


Ionization Equilibrium Models

Saha equilibrium:³⁹

$$(17) \quad \frac{N_e N_1^*(Z)}{N_n^*(Z-1)} = 6.0 \times 10^{21} \frac{g_1^Z T_e^{3/2}}{g_n^{Z-1}} \exp \left(-\frac{E_\infty^Z(n, l)}{T_e} \right) \text{ cm}^{-3}.$$

where g_n^Z is the statistical weight for level n of charge state Z and $E_\infty^Z(n, l)$ is the ionization energy of the neutral atom initially in level (n, l) , given by Eq. (2).

In a steady state at high electron density,

$$(18) \quad \frac{N_e N^*(Z)}{N^*(Z-1)} = \frac{S(Z-1)}{\alpha_3},$$

a function only of T .

Conditions for LTE:³⁹

(a) Collisional and radiative excitation rates for a level n must satisfy

$$(19) \quad Y_{nm} \gtrsim 10 A_{nm}.$$

(b) Electron density must satisfy

$$(20) \quad N_e \gtrsim 7 \times 10^{18} Z^7 n^{-17/2} (T/E_\infty^Z)^{1/2} \text{ cm}^{-3}.$$

Steady state condition in corona model:

$$(21) \quad \frac{N(Z-1)}{N(Z)} = \frac{\alpha_r}{S(Z-1)}.$$

Corona model is applicable if⁴⁰

$$(22) \quad 10^{12} t_I^{-1} < N_e < 10^{16} T_e^{7/2} \text{ cm}^{-3},$$

where t_I is the ionization time.

```

\input prolog
\hoffset=1.0truein
\voffset=1.0truein
\hsize=6.125truein
\vsize=9.0truein
\pageno=56
\parskip=5pt
{\headfont Radiation}

\inndent
{\it M.\ B.}  Energies and temperatures are in eV; all other quantities are in
cgs units except where noted.  $Z$ is the charge state ($Z=0$ refers to a
neutral atom); the subscript $e$ labels electrons.  $N$ is number density.

Average radiative decay rate of a state with principal quantum number $n$
is

$$A_n = \sum_m A_{nm} = 1.6 \times 10^{10} Z^4 n^{-9/2} \text{ sec.} \quad \text{\leqno{(23)}}$$


Natural linewidth  $(\Delta E \text{ in eV})$ :

$$\Delta E \approx \hbar / \Delta t = 4.14 \times 10^{-15} \text{ eV} \cdot \text{sec} \quad \text{\leqno{(24)}}$$

where  $\Delta t$  is the lifetime of the line.

Doppler width:

$$\Delta \lambda / \lambda = 7.7 \times 10^{-5} (T/\mu)^{1/2} \quad \text{\leqno{(25)}}$$

where  $\mu$  is the mass of the emitting atom or ion scaled by the proton mass.

Optical depth for a Doppler-broadened line:  $\tau$ 

$$\tau = 1.76 \times 10^{-13} \lambda (Mc^2/kT)^{1/2} NL = 5.4 \times 10^{-9} \lambda (\mu/T)^{1/2} NL \quad \text{\leqno{(26)}}$$


where  $\lambda$  is the wavelength and  $L$  is the physical depth of the plasma;
 $M$ ,  $N$ , and  $T$  are the mass, number density, and temperature of the absorber;
 $\mu$  is  $M$  divided by the proton mass.  Optically thin means  $\tau < 1$ .

Resonance absorption cross section at center of line:

$$\sigma_{\lambda=\lambda_c} = 5.6 \times 10^{-13} \lambda^2 / \Delta \lambda \quad \text{\leqno{(27)}}$$


Wien displacement law (wavelength of maximum black-body emission):

$$\lambda_{\max} = 2.50 \times 10^{-5} T^{-1} \text{ cm} \quad \text{\leqno{(28)}}$$


Radiation from the surface of a black body at temperature  $T$ :

$$W = 1.03 \times 10^5 T^4 \text{ watt/cm}^2 \quad \text{\leqno{(29)}}$$

\vfil\eject\end

```

Radiation

N. B. Energies and temperatures are in eV; all other quantities are in cgs units except where noted. Z is the charge state ($Z = 0$ refers to a neutral atom); the subscript e labels electrons. N is number density.

Average radiative decay rate of a state with principal quantum number n is

$$(23) \quad A_n = \sum_{m < n} A_{nm} = 1.6 \times 10^{10} Z^4 n^{-9/2} \text{ sec.}$$

Natural linewidth (ΔE in eV):

$$(24) \quad \Delta E \Delta t = h = 4.14 \times 10^{-15} \text{ eV sec.}$$

where Δt is the lifetime of the line.

Doppler width:

$$(25) \quad \Delta \lambda / \lambda = 7.7 \times 10^{-5} (T/\mu)^{1/2},$$

where μ is the mass of the emitting atom or ion scaled by the proton mass.

Optical depth for a Doppler-broadened line:³⁹

$$(26) \quad \tau = 1.76 \times 10^{-13} \lambda (Mc^2/kT)^{1/2} NL = 5.4 \times 10^{-9} \lambda (\mu/T)^{1/2} NL,$$

where λ is the wavelength and L is the physical depth of the plasma; M , N , and T are the mass, number density, and temperature of the absorber; μ is M divided by the proton mass. Optically thin means $\tau < 1$.

Resonance absorption cross section at center of line:

$$(27) \quad \sigma_{\lambda=\lambda_c} = 5.6 \times 10^{-13} \lambda^2 / \Delta \lambda \text{ cm}^2.$$

Wien displacement law (wavelength of maximum black-body emission):

$$(28) \quad \lambda_{\text{max}} = 2.50 \times 10^{-5} T^{-1} \text{ cm.}$$

Radiation from the surface of a black body at temperature T :

$$(29) \quad W = 1.03 \times 10^5 T^4 \text{ watt/cm}^2.$$

```

\input prolog
\hoffset=1.0truein
\voffset=1.0truein
\hsize=6.125truein
\vsize=9.0truein
\pageno=57
\parskip=5pt
Bremsstrahlung from hydrogen-like plasma:$^{26}$
$$P_{\rm Br}=1.69\times 10^{-32}N_eT_e^{-1/2}\sum\left[Z^2N(Z)\right],{\rm watt/cm}^3,\leqno{(30)}$$
where the sum is over all ionization states $$$.

Bremsstrahlung optical depth:$^{41}$
$$\tau=5.0\times 10^{-38}N_eN_iZ^2\overline{g}LT^{-7/2},\leqno{(31)}$$
where  $\overline{g}\approx 1.2$  is an average Gaunt factor and  $L$  is the
physical path length.

Inverse bremsstrahlung absorption coefficient$^{42}$ for radiation of angular
frequency  $\omega$ :

$$\kappa = 3.1\times 10^{-7}Z^2n_e\ln\Lambda T^{-3/2}\omega^{-2}(1-\omega_p^2/\omega^2)^{-1/2},{\rm cm}^{-1};\leqno{(32)}$$

here  $\Lambda$  is the electron thermal velocity divided by  $v$ , where  $v$  is
the larger of  $\omega$  and  $\omega_p$  multiplied by the larger of  $Ze^2/kT$  and
 $\hbar/(mkT)^{1/2}$ .

Recombination (free-bound) radiation:

$$P_r=1.69\times 10^{-32}N_eT_e^{-1/2}\sum\left[E_{\infty}^{Z-1}\over T_e\right],{\rm watt/cm}^3.\leqno{(33)}$$


Cyclotron radiation$^{26}$ in magnetic field  $B$ :

$$P_c=6.21\times 10^{-28}B^2N_eT_e,{\rm watt/cm}^3.\leqno{(34)}$$


For  $N_e kT_e = N_i kT_i = B^2/16\pi$  ( $\beta=1$ , isothermal plasma),$^{26}$

$$P_c=5.00\times 10^{-38}N_e^2T_e^{-2},{\rm watt/cm}^3.\leqno{(35)}$$


Cyclotron radiation energy loss  $\epsilon$ -folding time for a single electron:$^{41}$

$$t_c\approx\{9.0\times 10^{-8}B^{-2}\over 2.5+\gamma},{\rm sec},\leqno{(36)}$$

where  $\gamma$  is the kinetic plus rest energy divided by the rest energy
 $mc^2$ .

Number of cyclotron harmonics$^{41}$ trapped in a medium of finite depth  $L$ :

$$m_{\rm tr}=(57\beta BL)^{1/6},\leqno{(37)}$$

where  $\beta=8\pi MkT/B^2$ .

Line radiation is given by summing Eq. (9) over all species in the plasma.
\fill\egroup\end

```

Bremsstrahlung from hydrogen-like plasma:²⁶

$$(30) \quad P_{\text{Br}} = 1.69 \times 10^{-32} N_e T_e^{1/2} \sum [Z^2 N(Z)] \text{ watt/cm}^3.$$

where the sum is over all ionization states Z .

Bremsstrahlung optical depth:⁴¹

$$(31) \quad \tau = 5.0 \times 10^{-38} N_e N_i Z^2 \bar{g} L T^{-7/2},$$

where $\bar{g} \approx 1.2$ is an average Gaunt factor and L is the physical path length.

Inverse bremsstrahlung absorption coefficient⁴² for radiation of angular frequency ω :

$$(32) \quad \kappa = 3.1 \times 10^{-7} Z n_e^2 \ln \Lambda T^{-3/2} \omega^{-2} (1 - \omega_p^2/\omega^2)^{1/2} \text{ cm}^{-1};$$

here Λ is the electron thermal velocity divided by V , where V is the larger of ω and ω_p multiplied by the larger of $Z e^2/kT$ and $\hbar/(mkT)^{1/2}$.

Recombination (free-bound) radiation:

$$(33) \quad P_r = 1.69 \times 10^{-32} N_e T_e^{1/2} \sum \left[Z^2 N(Z) \left(\frac{E_\infty^{Z-1}}{T_e} \right) \right] \text{ watt/cm}^3.$$

Cyclotron radiation²⁶ in magnetic field \mathbf{B} :

$$(34) \quad P_c = 6.21 \times 10^{-28} B^2 N_e T_e \text{ watt/cm}^3.$$

For $N_e kT_e = N_i kT_i = B^2/16\pi$ ($\beta = 1$, isothermal plasma).²⁶

$$(35) \quad P_c = 5.00 \times 10^{-38} N_e^2 T_e^2 \text{ watt/cm}^3.$$

Cyclotron radiation energy loss e -folding time for a single electron:⁴¹

$$(36) \quad t_c \approx \frac{9.0 \times 10^8 B^{-2}}{2.5 + \gamma} \text{ sec.}$$

where γ is the kinetic plus rest energy divided by the rest energy mc^2 .

Number of cyclotron harmonics⁴¹ trapped in a medium of finite depth L :

$$(37) \quad m_{\text{tr}} = (57\beta BL)^{1/6},$$

where $\beta = 8\pi NkT/B^2$.

Line radiation is given by summing Eq. (9) over all species in the plasma.

\magnification=1728 \tolerance=1000
 \hoffset=1.25truein\voffset=1.0truein\hsize=6.0truein\vsize=9.0truein
 \input pointsize \sevenpoint
 \font\headfont=cmbx5 scaled \magstep2
 \def\ital#1{{\it #1}} \def\nocorr{\kern 0pt}
 \pageno=58 \parindent=15pt
 \def\item#1{\smallskip\hskip0pt\spacefactor=1000
 \hang\llap{#1\enspace}\ignorespaces}
 \centerline{\headfont REFERENCES}
 \bigskip\indent

Most of the formulas and data in this collection are well known and for all practical purposes are in the "public domain." The books and articles cited below are intended primarily not for the purpose of giving credit to the original workers, but $\{\textit{ital}\}$ (1) to guide the reader to sources containing related material and $\{\textit{ital}\}$ (2) to indicate where to find derivations, explanations, examples, etc., which have been omitted from this compilation. Additional material can also be found in D.L. Book, WRL Memorandum Report No. 3332 (1977).

\smallskip\hrenchspacing
 \item{1} See M. Abramowitz and I. A. Stegun, Eds., \ital{Handbook of Mathematical Functions} (Dover, New York, 1968), pp. 1-3, for a tabulation of some mathematical constants not available on pocket calculators.

\item{2} H. W. Gould, "Note on Some Binomial Coefficient Identities of Ehrenbaum," J. Math. Phys. {\bf 10}, 49 (1969); H. W. Gould and J. Kaucký, "Evaluation of a Class of Binomial Coefficient Summations," J. Comb. Theory {\bf 1}, 233 (1966).

\item{3} R. S. Newberger, "New Sum Rule for Products of Bessel Functions with Application to Plasma Physics," J. Math. Phys. {\bf 23}, 1278 (1982); {\bf 24}, 2283 (1983).

\item{4} P. M. Morse and H. Feshbach, \ital{Methods of Theoretical Physics} (McGraw-Hill Book Co., New York, 1953), Vol. II, pp. 47-52 and pp. 680-690.

\item{5} W. H. Bayard, "A Collection of Vector Formulas," Princeton University, Princeton, NJ, 1966 (unpublished), and personal communication (1977).

\item{6} See \ital{Fundamental Units and Symbols} 1973, report of the Special Committee of the Royal Society, 2nd Edition (Royal Society, London, 1975) for a discussion of nomenclature in SI units.

\item{7} E. R. Cohen and P. M. Taylor, "Least Squares Adjustment of the Fundamental Constants," J. Phys. Chem. Ref. Data {\bf 2}, 663 (1973).

\item{8} E. M. Verbeke, "Dimensionally Correct Transformations between Different Systems of Units," Amer. J. Phys. {\bf 36}, 1132 (1968).

\item{9} J. Stratton, \ital{Electromagnetic Theory} (McGraw-Hill Book Co., New York, 1941), p. 10.

\item{10} See \ital{IEEE Reference Data for Engineers} (IEEE Press, New York, 1974), pp. 1-10. These definitions are International Telecommunications Union (ITU) standards.

REFERENCES

Most of the formulas and data in this collection are well known and for all practical purposes are in the "public domain." The books and articles cited below are intended primarily not for the purpose of giving credit to the original workers, but (1) to guide the reader to sources containing related material and (2) to indicate where to find derivations, explanations, examples, etc., which have been omitted from this compilation. Additional material can also be found in D. L. Book, NRL Memorandum Report No. 3332 (1977).

1. See M. Abramowitz and I. A. Stegun, Eds., *Handbook of Mathematical Functions* (Dover, New York, 1968), pp. 1-3, for a tabulation of some mathematical constants not available on pocket calculators.
2. H. W. Gould, "Note on Some Binomial Coefficient Identities of Rosenbaum," *J. Math. Phys.* **10**, 49 (1969); H. W. Gould and J. Kaucky, "Evaluation of a Class of Binomial Coefficient Summations," *J. Comb. Theory* **1**, 233 (1966).
3. B. S. Newberger, "New Sum Rule for Products of Bessel Functions with Application to Plasma Physics," *J. Math. Phys.* **23**, 1278 (1982); **24**, 2250 (1983).
4. P. M. Morse and H. Feshbach, *Methods of Theoretical Physics* (McGraw-Hill Book Co., New York, 1953), Vol. I, pp. 47-52 and pp. 656-666.
5. W. D. Hayes, "A Collection of Vector Formulas," Princeton University, Princeton, N.J., 1956 (unpublished), and personal communication (1977).
6. See *Quantities, Units and Symbols*, report of the Symbols Committee of the Royal Society, 2nd Edition (Royal Society, London, 1975) for a discussion of nomenclature in SI units.
7. E. R. Cohen and B. N. Taylor, "Least-Squares Adjustment of the Fundamental Constants," *J. Phys. Chem. Ref. Data* **2**, 663 (1973).
8. E. S. Weibel, "Dimensionally Correct Transformations between Different Systems of Units," *Amer. J. Phys.* **36**, 1130 (1968).
9. J. Stratton, *Electromagnetic Theory* (McGraw-Hill Book Co., New York, 1941), p. 508.
10. *ITT Reference Data for Engineers* (Eams and Co., Indianapolis, IN, 1975), p. I-1. These definitions are International Telecommunications Union (ITU) Standards.
11. H. E. Thomas, *Handbook of Microwave Techniques and Equipment* (Prentice-Hall, Englewood Cliffs, N.J., 1972), p. 9. Further subdivisions are defined in Ref. 10, p. I-3.

- \item{11.}H. E. Thomas, \ital{Handbook of Microwave Techniques and Equipment} (Pren\~tice-Hall, Englewood Cliffs, NJ, 1972), p. 9. Further subdivisions are defined in Ref. 10, p. I--3.
- \item{12.}J. F. Catchpole and G. Fulford, Ind. and Eng. Chem. {\bf 58}, 47 (1966); reprinted in recent editions of the \ital{Handbook of Chemistry and Physics} (Chemical Rubber Co., Cleveland, OH) on pp. F306--323.
- \item{13.}W. D. Hayes, "The Basic Theory of Gasdynamic Discontinuities," in \ital{Fundamentals of Gas Dynamics\nocorr}, Vol. III, \ital{High Speed Aerodynamics and Jet Propulsion\nocorr}, H. W. Emmons, Ed. (Princeton University Press, Princeton, NJ, 1958).
- \item{14.}W. B. Thompson, \ital{An Introduction to Plasma Physics} (Addison-Wesley Publishing Co., Reading, MA, 1962), pp. 86--95.
- \item{15.}L. D. Landau and E. M. Lifshitz, \ital{Fluid Mechanics} (Addison-Wesley Publishing Co., Reading, MA, 1959), pp. 319--331.
- \item{16.}The \mathcal{Z} function is tabulated in B. D. Fried and S. D. Conte, \ital{The Plasma Dispersion Function} (Academic Press, New York, 1961).
- \item{17.}R. W. Landau and S. Cuperman, "Stability of Anisotropic Plasmas to Almost-Perpendicular Magnetosonic Waves," J. Plasma Phys. {\bf 6}, 495 (1971).
- \item{18.}B. D. Fried, C. L. Hedrick, J. McCune, "Two-Pole Approximation for the Plasma Dispersion Function," Phys. Fluids {\bf 11}, 249 (1968).
- \item{19.}B. A. Trubnikov, "Particle Interactions in a Fully Ionized Plasma," \ital{Reviews of Plasma Physics\nocorr}, Vol. 1 (Consultants Bureau, New York, 1965), p. 105.
- \item{20.}J. M. Greene, "Improved Bhatnagar--Gross--Krook Model of Elec\~tron-Ion Collisions," Phys. Fluids {\bf 16}, 2022 (1973).
- \item{21.}S. I. Braginskii, "Transport Processes in a Plasma," \ital{Reviews of Plasma Physics\nocorr}, Vol. 1 (Consultants Bureau, New York, 1965), p. 205.
- \item{22.}J. Sheffield, \ital{Plasma Scattering of Electromagnetic Radiation} (Academic Press, New York, 1975), p. 6 (after J. W. Paul).
- \item{23.}K. H. Lloyd and G. H. \~arendel, "Numerical Modeling of the Drift and Deformation of Ionospheric Plasma Clouds and of their Interaction with Other Layers of the Ionosphere," J. Geophys. Res. {\bf 78}, 7389 (1973).
- \item{24.}C. W. Allen, \ital{Astrophysical Quantities\nocorr}, 2nd edition (Athlone Press, London, 1963), p. 161.
- \item{25.}G. L. Withbridge and R. W. Hayes, "Mass and Energy Flow in the Solar Chromosphere and Corona," Ann. Rev. Astrophys. {\bf 15}, 363 (1977).
- \item{26.}S. Glasstone and R. H. Lovberg, \ital{Controlled Thermonuclear Reactions} (Van Nostrand, New York, 1960), Chapt. 2.
- \item{27.}References to experimental measurements of branching ratios and cross sections are listed in F. K. McGowan, et al., Nucl. Data Tables {\bf A6},

12. J. P. Catchpole and G. Fulford, *Ind. and Eng. Chem.* **58**, 47 (1966); reprinted in recent editions of the *Handbook of Chemistry and Physics* (Chemical Rubber Co., Cleveland, OH) on pp. F306-323.
13. W. D. Hayes, "The Basic Theory of Gasdynamic Discontinuities," in *Fundamentals of Gas Dynamics*, Vol. III, *High Speed Aerodynamics and Jet Propulsion*, H. W. Emmons, Ed. (Princeton University Press, Princeton, NJ, 1958).
14. W. B. Thompson, *An Introduction to Plasma Physics* (Addison-Wesley Publishing Co., Reading, MA, 1962), pp. 86-95.
15. L. D. Landau and E. M. Lifshitz, *Fluid Mechanics* (Addison-Wesley Publishing Co., Reading, MA, 1959), pp. 319-331.
16. The Z function is tabulated in B. D. Fried and S. D. Conte, *The Plasma Dispersion Function* (Academic Press, New York, 1961).
17. R. W. Landau and S. Cuperman, "Stability of Anisotropic Plasmas to Almost-Perpendicular Magnetosonic Waves," *J. Plasma Phys.* **6**, 495 (1971).
18. B. D. Fried, C. L. Hedrick, J. McCune, "Two-Pole Approximation for the Plasma Dispersion Function," *Phys. Fluids* **11**, 249 (1968).
19. B. A. Trubnikov, "Particle Interactions in a Fully Ionized Plasma," *Reviews of Plasma Physics*, Vol. 1 (Consultants Bureau, New York, 1965), p. 105.
20. J. M. Greene, "Improved Bhatnagar Gross Krook Model of Electron-Ion Collisions," *Phys. Fluids* **16**, 2022 (1973).
21. S. I. Braginskii, "Transport Processes in a Plasma," *Reviews of Plasma Physics*, Vol. 1 (Consultants Bureau, New York, 1965), p. 205.
22. J. Sheffield, *Plasma Scattering of Electromagnetic Radiation* (Academic Press, New York, 1975), p. 6 (after J. W. Paul).
23. K. H. Lloyd and G. Härendel, "Numerical Modeling of the Drift and Deformation of Ionospheric Plasma Clouds and of their Interaction with Other Layers of the Ionosphere," *J. Geophys. Res.* **78**, 7389 (1973).
24. C. W. Allen, *Astrophysical Quantities*, 2nd edition (Athlone Press, London, 1963), p. 161.
25. G. L. Withbroe and R. W. Noyes, "Mass and Energy Flow in the Solar Chromosphere and Corona," *Ann. Rev. Astrophys.* **15**, 363 (1977).
26. S. Glasstone and R. H. Lovberg, *Controlled Thermonuclear Reactions* (Van Nostrand, New York, 1960), Chapt. 2.
27. References to experimental measurements of branching ratios and cross sections are listed in F. K. McGowan, et al., *Nucl. Data Tables*

353 (1969); {\bf A8}, 199 (1970). The yields listed in the table are calculated directly from the mass defect.

\item{28.}G.~H. Miley, H.~Towner and N.~Ivich, \ital{Fusion Cross Section and Reactivities\nocorr}, Rept. COO-2218-17 (University of Illinois, Urbana, IL, 1974).

\item{29.}J.~M. Creedon, "Relativistic Brillouin Flow in the High ν/γ Limit," J.~Appl. Phys. {\bf 46}, 2946 (1975).

\item{30.}See, for example, A.~B. Mikhailovskii, \ital{Theory of Plasma Instabilities\nocorr} Vol.~I (Consultants Bureau, New York, 1974). The table on pp.~48--49 was compiled by K.~Papadopoulos.

\item{31.}Table prepared from data compiled by J.~M. McMahon (personal communication, 1987).

\item{32.}M.~J. Seaton, "The Theory of Excitation and Ionization by Electron Impact," in \ital{Atomic and Molecular Processes\nocorr}, D.~R. Bates, Ed. (New York, Academic Press, 1962), Chapt. 11.

\item{33.}H.~Van Regemorter, "Rate of Collisional Excitation in Stellar Atmospheres," Astrophys.~J. {\bf 136}, 906 (1962).

\item{34.}A.~C. Kolb and R.~W.~P. McWhirter, "Ionization Rates and Power Loss from θ -Pinches by Impurity Radiation," Phys. Fluids {\bf 7}, 519 (1964).

\item{35.}R.~W.~P. McWhirter, "Spectral Intensities," in \ital{Plasma Diagnostic Techniques\nocorr}, R.~H. Huddlestone and S.~L. Leonard, Eds. (Academic Press, New York, 1965).

\item{36.}M.~Gryzinski, "Classical Theory of Atomic Collisions I. Theory of Inelastic Collision," Phys. Rev. {\bf 138A}, 336 (1965).

\item{37.}M.~J. Seaton, "Radiative Recombination of Hydrogen Ions," Mon. Not. Roy. Astron. Soc. {\bf 119}, 81 (1959).

\item{38.}Ya.~B. Zel'dovich and Yu.~P. Raizer, \ital{Physics of Shock Waves and High-Temperature Hydrodynamic Phenomena} (Academic Press, New York, 1966), Vol.~I, p.~407.

\item{39.}H.~R. Griem, \ital{Plasma Spectroscopy} (Academic Press, New York, 1966).

\item{40.}T.~F. Stratton, "X-Ray Spectroscopy," in \ital{Plasma Diagnostic Techniques\nocorr}, R.~H. Huddlestone and S.~L. Leonard, Eds. (Academic Press, New York, 1965).

\item{41.}G.~Bekefi, \ital{Radiation Processes in Plasmas} (Wiley, New York, 1966).

\item{42.}T.~W. Johnston and J.~M. Dawson, "Correct Values for High-Frequency Power Absorption by Inverse Bremsstrahlung in Plasmas," Phys. Fluids {\bf 16}, 722 (1973).
\vfill\ejct\end

- A6**, 353 (1969); **A8**, 199 (1970). The yields listed in the table are calculated directly from the mass defect.
28. G. H. Miley, H. Towner and N. Ivich, *Fusion Cross Section and Reactivities*, Rept. COO-2218-17 (University of Illinois, Urbana, IL, 1974).
 29. J. M. Creedon, "Relativistic Brillouin Flow in the High ν/γ Limit," *J. Appl. Phys.* **46**, 2946 (1975).
 30. See, for example, A. B. Mikhailovskii, *Theory of Plasma Instabilities* Vol. I (Consultants Bureau, New York, 1974). The table on pp. 48-49 was compiled by K. Papadopoulos.
 31. Table prepared from data compiled by J. M. McMahon (personal communication, 1987).
 32. M. J. Seaton, "The Theory of Excitation and Ionization by Electron Impact," in *Atomic and Molecular Processes*, D. R. Bates, Ed. (New York, Academic Press, 1962), Chapt. 11.
 33. H. Van Regemorter, "Rate of Collisional Excitation in Stellar Atmospheres," *Astrophys. J.* **136**, 906 (1962).
 34. A. C. Kolb and R. W. P. McWhirter, "Ionization Rates and Power Loss from θ -Pinches by Impurity Radiation," *Phys. Fluids* **7**, 519 (1964).
 35. R. W. P. McWhirter, "Spectral Intensities," in *Plasma Diagnostic Techniques*, R. H. Huddleston and S. L. Leonard, Eds. (Academic Press, New York, 1965).
 36. M. Gryzinski, "Classical Theory of Atomic Collisions I. Theory of Inelastic Collision," *Phys. Rev.* **138A**, 336 (1965).
 37. M. J. Seaton, "Radiative Recombination of Hydrogen Ions," *Mon. Not. Roy. Astron. Soc.* **119**, 81 (1959).
 38. Ya. B. Zel'dovich and Yu. P. Raizer, *Physics of Shock Waves and High-Temperature Hydrodynamic Phenomena* (Academic Press, New York, 1966), Vol. I, p. 407.
 39. H. R. Griem, *Plasma Spectroscopy* (Academic Press, New York, 1966).
 40. T. F. Stratton, "X-Ray Spectroscopy," in *Plasma Diagnostic Techniques*, R. H. Huddleston and S. L. Leonard, Eds. (Academic Press, New York, 1965).
 41. G. Bekefi, *Radiation Processes in Plasmas* (Wiley, New York, 1966).
 42. T. W. Johnston and J. M. Dawson, "Correct Values for High-Frequency Power Absorption by Inverse Bremsstrahlung in Plasmas," *Phys. Fluids* **16**, 722 (1973).

END

12-87

DTIC